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Viking Lander Atlas of Mars

Sidney Liebes, Jr.
Stanford University
Stanford, California

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To Tim Mutch ... explorer

PREFACE

The 1976 Viking scientific exploratory mission to Mars successfully placed two identical spacecraft in orbit and two identical landers on the surface of the planet. This report presents half-size reproductions of the extensive sets of systematic map products that have been generated for the two Mars Viking landing sites from stereo pairs of images radioed back to Earth from the landers. These maps span from the immediate foreground to the remote limits of ranging capability, several hundred meters from the spacecraft. The maps are of two kinds: 1) elevation contour, and 2) vertical profile.

The report is written with the following aims in mind: 1) To provide, in convenient form, relatively inexpensive reduced-size reproductions of the complete set of nearly 200 true-scaled maps, in the expectation that the reduced-size versions will suffice for certain requirements. 2) To describe the maps, and discuss

their unorthodox appearance that results principally from the positions of the cameras relative to the Martian surface. 3) To make available an organized version of the complete set of maps that will enable a convenient means of identifying the individual sheets for which true-scale copies might be desired. 4) To provide a companion resource to collections of true-scale sheets, from which one may index into the true-scale collection. Individual copies of the true-scale maps are available from the National Space Science Data Center.

This report is divided into three parts, and a set of appendices: Part I contains background and explanatory material important for understanding and utilizing the map collections; less essential material is deferred to appendices at the end of the report. Part II contains the collection of lander 1 maps. Part III contains the collection for lander 2.

ACKNOWLEDGMENTS

There is a lengthy history to the production of the maps contained in this report. The Stanford Stereo System, especially constructed for the creation of maps from digital images of the surface of Mars, has been previously described [Liebes, 1977]. The many individuals who contributed directly to the development and support of that system are noted in the earlier report. I wish particularly to single out, however, the contribution made both to the success of the stereo system and to the creation of the map data by Arnold A. Schwartz, then of the Image Processing Laboratory, Jet Propulsion Laboratory, Pasadena, CA. Arnie wrote all of the substantial computer code that supported the stereo mapping system, and subsequently spent countless hours commanding the computer while the stereophotogrammetry of the Martian surface was being performed. Joseph W. Berry, then also at the Image Processing Lab at JPL, good naturedly provided relief to Arnie during a portion of the map production. All of the stereophotogrammetry for the systematic maps was performed by the author.

The range data sets, which are the prime photogrammetric product, represent the spatial location of each of the recorded points on every contour or vertical profile by their three-dimensional position in a spacecraft specific system of coordinates. After the creation of these range data sets, at the JPL Image Processing Laboratory, magnetic tapes of the sets were transported to Stanford University, where the laborious tasks of writing code to read the sets, designing the format for the map products, and executing the extensive production activity was undertaken.

Two individuals played a particularly key role in the tasks of program design, implementation and execution, and in the production of the map masters at Stanford. In this regard, I wish to note the exceptional contribution made by Stuart McDonald, largely while he was an undergraduate student in computer science at the University of California at Berkeley. Stuart spent the summer of 1979 designing and implementing the code that enabled us to trace through the range data sets, and implementing most of the code for the formatting of the contour maps. His work was creative, efficient, meticulous and dedicated. John C. Gilbert, then a masters student in computer science at Stanford, adapted Stuart's code to the processing of the vertical profiles and implemented the format for the the plotting of these sheets.

This was an extensive task in itself, and John executed this task with the greatest of skill and dedication. A Stanford freshman, Ernest Adams, provided not only valuable assistance in elements of the general code work, but additionally did much of the adaptation of Stuart's code for the reading of mosaic line and sample values out of the range data sets and formatting of the camera perspectives of the contour maps that are included as a part of the collection of map products.

Hans Moravec, then a PhD graduate student in computer science at Stanford, provide invaluable assistance to us, initially in writing code to enable us to read the IBM formatted range data sets on our DEC PDP-10 computer, and subsequently in educating us to the use of his computer graphics and character font routines on the Varian electrostatic plotter employed in the production of the map masters. I am deeply grateful to Robert Tucker, of the Stanford Medical School SUMEX computer staff, for his skilled and generous assistance to us in reformatting the magnetic tape data, as received from the Jet Propulsion Laboratory, so that it could be read and transmitted between DEC computer systems at Stanford.

Dr. Henry J. Moore, of the U.S. Geological Survey, Menlo Park, CA served as a member of the physical properties team for the Viking lander. Hank meticulously made visual measurements of camera pixel positions for conjugate points in stereo image pairs and performed extensive hand calculations of 3-D field point coordinates, both to support his own research interests and to help check the performance of the Stanford Stereo System. Hank has additionally conducted spot checks of the systematic elevation contour maps for both landers. Hank's invaluable collaboration is deeply appreciated.

And finally, I wish to acknowledge my deep appreciation to Prof. Elliott C. Levinthal of Stanford University for his collegueship and support during the Viking years, and for his general supervision of the production of the Viking lander camera mosaics that constituted the input imagery for the photogrammetric processing that is the subject of this report.

Support for the production of these maps from the previously generated computer readable range data sets was provided under contracts NAS1-9682 and JPL-955249, and grants NSG-7538 and NAGW-128 (Mars Data Analysis Program).

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PART I BACKGROUND AND EXPLANATION OF MAPS

1. INTRODUCTION

1.1 The Viking Mission

The Viking Mission launched two unmanned spacecraft toward Mars in the summer of 1975, for the purpose of scientific exploration. Each craft consisted of an orbiter and a lander. The craft accomplished Mars orbital insertion seven weeks apart in the summer of 1976. The landing site certification processes resulted in revision of the preselected landing sites for both landers. Useful points of entry into the earlier scientific literature on the mission are the special issue of the *Journal of Geophysical Research*, vol 82, Sept. 30, 1977, and vol 84, Dec. 30, 1979. The reader interested in the dramatic accomplishments of the flight phase of the mission — especially the remarkable human and engineering achievements involved in solving problems that arose — is referred to the paper of Martin and Young [Martin, 1976].

1.2 Lander Locations

The areographic coordinate [de Vaucouleurs, 1973] locations for the Viking landers have been established to be as follows [Davies, 1978; Morris, 1980]:

Viking lander 1: lat 22.483° N., long 47.968° W.

Viking lander 2: lat 47.966° N., long 225.736° W.

1.3 Lander Cameras

Among the complement of equipment on each lander were a pair of identical panoramic cameras [Patterson, 1977]. These cameras enabled extensive stereoscopic imaging of the landing sites. The camera was a point scanning device capable of acquiring data from 40° above the horizon to 60° below, through nearly 360° of azimuth. The camera was configured about a vertical optical axis. A photosensor array, consisting of a tightly clustered set of 12 light-sensitive silicon photodiodes, was located in the focal region below the lens. The diode apertures were of two sizes, corresponding to angular resolutions of 0.12° and 0.04°. Four of these were "high resolution" 0.04° broad band (black

and white) diodes, each stepped at a slightly different focal distance from the lens. Virtually all of the imaging data used for mapping purposes was acquired at 0.04° resolution. The choice of diode was dictated by the range span of the section of the landscape to be imaged.

Mounted in the object space immediately above the lens was a mirror, with a capability to nod about a horizontal axis. This mirror enabled the Martian landscape to be imaged upon the photosensor array. In the standard mode of operation, the mirror nodded five times per second. As the image of the scene was swept across the photosensor array, the diode selected to be active would sample the image 512 times, at angular spacings equal to the diode resolution. Each of the sampled light intensity levels was digitized at 6 bits, corresponding to 64 shades of gray. At the end of a downward scan through the scene, the entire camera would rotate one angular resolution element about the vertical axis, and the process would be repeated for the next scan line. In this manner, in the high resolution 0.04°-mode, an approximately 20° vertical acquisition would be achieved. The vertical acquisition window could be commanded to image centers at 10° vertical intervals. The start and stop azimuth positions for the rotation of the camera about the vertical axis were under command control. Adjustable offset and gain controls could be set for brightness and contrast conditions. The images were radioed back to Earth either directly from the lander, or more commonly via a relay link through the more powerful transmitter of the orbiter.

The camera photogrammetric reference points were located approximately 0.82 m apart at equal heights above the lander deck, and at 1.3 m above the nominal landing plane. In the course of the mission, practically all of the visible panorama was imaged in high resolution at different times of day by both cameras of both landers. Thereby, virtually all of the scene that was visible in common to both cameras was imaged stereoscopically.

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2. MAPPING THE MARTIAN SURFACE

2.1 The Stereo Mapping System

A versatile stereoscopic mapping system was developed jointly by Stanford University and the Jet Propulsion Laboratory [Liebes, 1977] to support the lander cameras. This system enabled the prompt generation of precise topographic maps in support of mission critical acquisition of surface sample material for the several miniature laboratories on board the landers. It additionally supported use of the sampler arm for Martian surface physical properties experiments [Moore, 1978]. When pressure abated to deliver map products in support of sample arm activities, a systematic mapping was undertaken of the entire stereoscopically imaged landing sites, from the immediate foreground to the remote limits of ranging capability, over 100 m from the landers.

The ranging and mapping system consisted of the Stanford Stereo Station, a computer, image display hardware, and a substantial computer program, RANGER. The Stanford Stereo Station supported a pair of video monitors arranged so that they could be viewed through a panning stereoscope. Computer resident digital stereo imaging data could be called forth for display upon the video monitors. A photogrammetrist, peering into the stereoscope, perceived a 3-D model of the Martian relief. The computer could be commanded to project upon the monitors an artificial 3-space mark, consisting of a pair of appropriately coupled video dot overlay cursors. The photogrammetrist, stereoscopically viewing the displayed images, could employ a hand-input device in the form of a trackball to move the 3-D cursor through the space of the visualized stereoscopic model. The cursor could be constrained to arbitrary mathematical surfaces in the model. The systematic maps contained within this report were generated by commanding the computer to constrain the 3-space mark to a selected member of a set of horizontal or vertical surfaces, for the purpose of enabling the generation of elevation contours or vertical profiles, respectively. The task of the photogrammetrist then became the traditional one of moving the mark along his perception of the intersection of the mathematical surface of constraint with the Martian relief. In this manner, digitized versions of the contours

or profiles of any desired kind could be generated and stored. The fundamental system product was a computer readable and plottable, range data set (RDS). The RDS represented contour and vertical profile lines as sequences of 3-D points. The record for each point consisted of its three spatial coordinates, plus the corresponding scan line and sample values associated with the image locations in each of the images — a set of seven values. Secondary products included maps, stereo projections of maps, and image mosaics overlaid with contours and profiles.

2.2 Input Images to the Stereo System

The stereo mapping system was designed to operate on images from arbitrary sources, exhibiting arbitrary mappings between recording systems and data storage formats. In particular it could operate, for the Viking lander, upon either raw geometrically untransformed images, taken with any of the photodiodes, or upon high resolution computer assembled mosaics, constructed from images taken at three different times of day — morning, mid-day and evening [Levinthal, 1980]. All of the sampler arm support work was performed with raw, geometrically untransformed images. The mosaics served as the input images to the stereo mapping system for the generation of the systematic maps. These mosaics were compiled from numerous raw images acquired with each of the four high resolution diodes. Each elemental input image was geometrically transformed to compensate for diode-specific geometric calibration characteristics prior to incorporation into a mosaic [Wolf, 1981]. The initial mosaics used in the generation of the systematic map data manifested lander tilt, evidenced by a the sinuous warp of the horizon. Later, all mosaics were transformed to the "detilted" form they would have exhibited if the camera were rotated about its photogrammetric reference point so as to remove the rotation induced by lander tilt. The stereo system could operate equally well on both "tilted" or "detilted" mosaics, and indeed both were used, according to availability, in generating the systematic RDSs. All of the mosaics presented later in this report are of the "detilted" form.

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3. EXPLANATION OF MAPS

3.1 Systematic Map Types

Two different types of comprehensive systematic maps were developed:

Elevation Contours: Systematic contour map data, consisting of intersections with the surface of Mars of planes oriented perpendicular to the local gravity vector. These extend from the immediate foreground to the remote limits of ranging capability, in excess of 100 m range, for the front and back of both landers.

Vertical Profiles: Systematic vertical profiles, consisting of intersections with the surface of Mars of planes radiating out from the lander and containing the local gravity vector. These are generated at 5°-azimuth intervals, from the immediate foreground to the remote limits of ranging capability, for the front and back of both landers.

3.2 General Comment

The systematic map data, derived as it is from two cameras located approximately 0.82 m apart and roughly 1.3 m above the nominal Martian surface, is of a quite unorthodox character. The stereo ranging accuracy is approximately quadratically dependent upon range, with absolute single point ranging accuracy varying from approximately ± 1 cm near the lander to roughly ± 20 m at 100 m range, with rapid deterioration beyond (see Appendix D; also [Liebes, 1977]). Relative accuracy of the data, out to a few tens of meters, is estimated to be an order of magnitude better than the absolute accuracy.

Given that the ranging accuracy of the map data is highly range dependent, it was adopted as a production policy for the elevation contour maps that the complement of scales should be such that somewhere within the map collection there should appear a sheet upon which any contour segment could be examined with the full level of meaningful resolution. This policy was partially relaxed for the vertical profiles.

Though the operation of the mapping system and the production of RDSs could, in principle, be accomplished by a single individual, in practise it invariably involved two people, one to command the computer and the other to perform the stereophotogrammetry. Efficient operation required knowledge both of the program RANGER and the operating system. It was not infrequent that problems of one sort or another related to either the program or the support system would arise. Furthermore, the photogrammetrist was kept busy generating cartographic data, making judgements about where to go next, be it to continue a line or profile, to window to different portions of the images, etc.

3.2.1 Maps and Support Products

Several products have been prepared to aid in the interpretation of the maps. In order to enable the reader to associate the lines of the contour and vertical profile maps with the corresponding physical features on Mars, there are included at the beginning of each collection of maps reproductions of the stereo pairs of computer assembled mosaics used to generate the maps. These mosaics exhibit overlays of the entire set of elevation con-

tour or vertical profile lines. As the perspective of the contour maps is orthographic from above, the perspective of the vertical profile plots is normal to the individual profile planes, and the perspective of the cameras is oblique from the surface, a little practise will be required to make the association between corresponding lines in the maps and mosaic overlays. The map overlay lines on the mosaics are not annotated with either elevation or azimuth values. However, companion products have been included along with the mosaics. They consist of a reprojection of the map data as they would appear to camera number 1, the left hand of the two cameras (looking outward from the front of the lander). The lines in the reprojections are labeled in one instance with the contour-elevation and in the other with profile-azimuth values. Since these "camera perspective" representations of the annotated lines have the same projective appearance as the mosaic overlays, one may readily determine the parameter values of lines in the overlay images.

3.3 Elevation Contour Maps

Each of the elevation contour map grids is standardized to a 50-cm-square frame (true-scale), corresponding to a 25-cm-square frame in the half-size reproductions contained in this report. The contour map production policy lead to the generation of single sheets for the larger scale numbers. However, multiple sheets were required for the smaller scale numbers. The number of sheets required for each of the landers differed, due to variations in landing site topography and lander tilt. There are 64 elevation contour sheets for lander 1 and 66 for lander 2.

The coordinate system of the maps is the Local Gravity-Normal (LGN) system. The LGN system is orthogonal and right-handed. The LGN z-axis is directed vertically upward toward the local Martian zenith. The LGN x- and y-axes are horizontal. The origin is approximately 28 cm below the nominal landing surface, at a point approximately centrally located beneath the lander. The LGN x-axis is directed horizontally to the right on the maps. The intercamera baseline is nearly parallel to the LGN x-axis. The LGN y-axis is directed upward on the sheets, and points approximately forward from the lander. The LGN origin coincides with that of the Lander Aligned Coordinate System (LACS). Further details regarding coordinate systems are presented in Appendix A.

The production policy lead to the creation of elevation contour maps at eleven different scales: 1:1, 1:2, 1:5, 1:10, 1:20, 1:50, 1:100, 1:200, 1:500, 1:1000, and 1:2000. The number of sheets required at any given scale is variable. For each scale the sheet complement is a contiguous, non-overlapping array.

Only single sheets were produced at the larger scale numbers, and for these the lander has been "sheet-centered," that is, located in the center of the sheet. Multiple sheets have been produced for the smaller scale numbers, and for these sheets the lander is "sheet-cornered," that is, located at the mutual corners of a real or virtual subset quadruple of the complement of sheets.

Figure 1 illustrates an example of an elevation contour map sheet. Immediately below the Viking lander number, at bottom center, the scale and a sheet array designator are indicated. The scale is valid for the true-scale sheets. Since the sample sheet

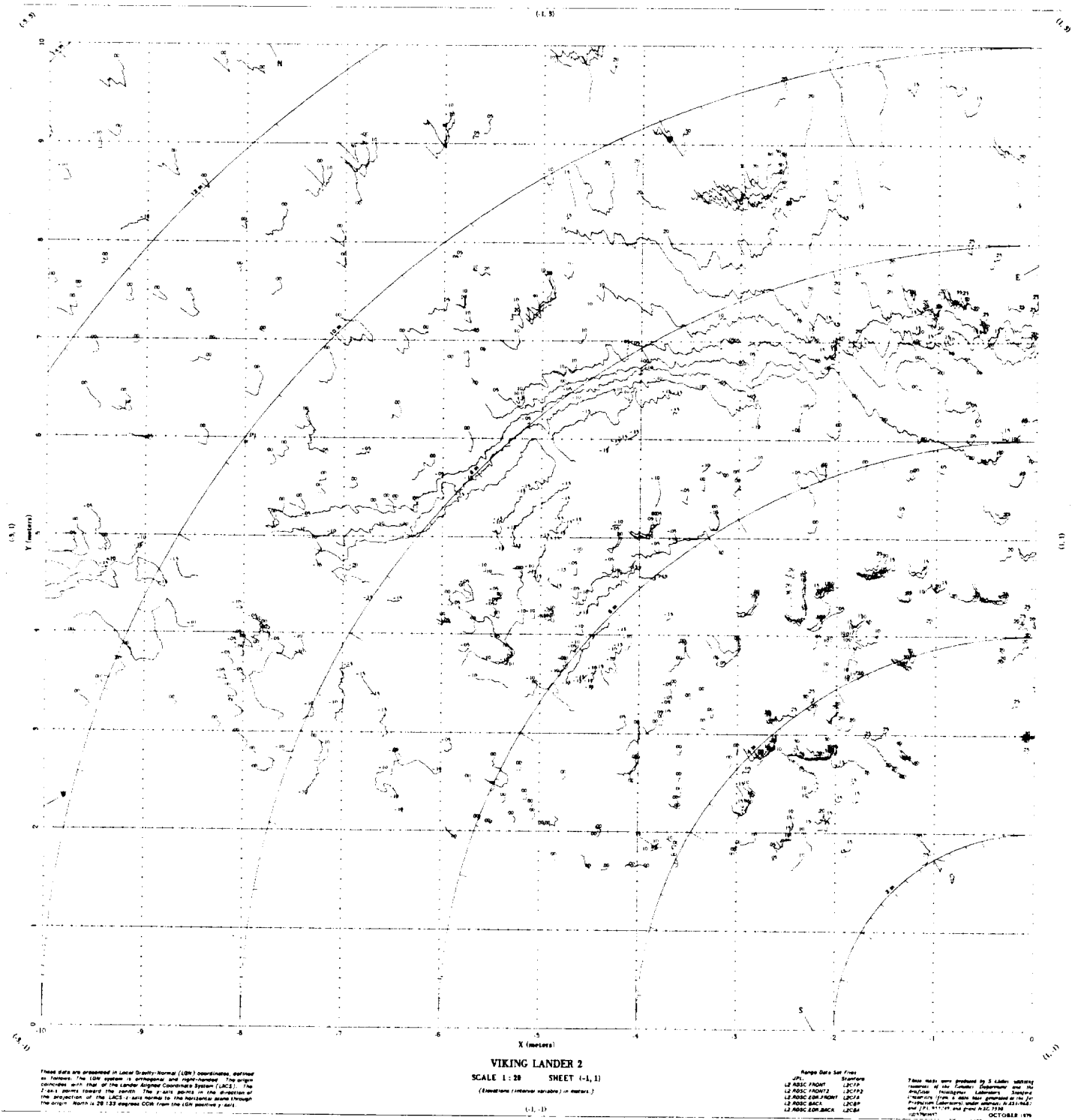


Fig. 1. Example of an elevation contour map sheet.

is indicated to be "SCALE 1:20," the actual scale of the sample sheet, reproduced here at half-size, is 1:40.

The sheet array designator consists of a number pair enclosed in parentheses, e.g. "SHEET (-1, 1)." This designator indicates at sheet scale the position of the sheet center, relative to the lander, in units of half-sheet dimensions. The first member of the designator number pair is equal to the number of half-sheet dimensions that the subject sheet center is displaced in the LGN x-direction, i.e. to the right of the lander; the second number of the pair is equal to the number of half-sheet dimensions that the subject sheet center is displaced in the LGN y-direction, that is forward from the lander. A sign reversal designates a displacement in the opposite direction. This scheme leads to designators composed of odd number pairs if the lander is sheet-cornered for the set, and to designators composed of even number pairs if the set is characterized by sheet-center lander placement. Thus, e.g., in the case of the larger scale numbers, where only a single sheet is produced, the lander is centered on the sheet, and the sheet designator is (0,0). The designator for the sample sheet is (-1, 1). The "-1" indicates that the center of the sheet is one half-sheet dimension to the left of the lander and the "+1" indicates that it is one half-sheet dimension in front of the lander. Thus, for this sheet the lander is located at the lower right corner of the sheet; the lander is "sheet-cornered" for this set of sheets. As the front of the lander faces toward the top edge of the sheet, i.e., in the positive LGN y-direction, and the intercamera baseline is nearly parallel to the LGN x-axis, camera 1 is located at $x_{LGN} \approx -0.41$ m, $y_{LGN} \approx 0$ m, and camera 2 is located at $x_{LGN} \approx +0.41$ m, $y_{LGN} \approx 0$ m).

The block of information in the lower left corner of the sheet describes the coordinate system of the maps -- the LGN system -- and indicates the direction of north relative to that system.

The block of information to the right of the sheet designator lists the names of the range data set files from which the maps have been derived. As the naming convention is different at Stanford from that at JPL, both sets of file names have been indicated. The file names for all the RDSs used in generating the contour and vertical profile maps for both landers are listed in sect. 3 of Appendix C.

As all sheets are standardized to a 50 cm grid boundary (true-scale), a standard interior grid pattern has been adopted. Grid lines are at 5 cm (true-scale) spacing, and grid ticks at 5 mm (true-scale) intervals. The absolute significance of these marks is of course determined by the scale.

The sheet exhibits a set of arcs (or circles) concentric with the lander, at 10-sheet-cm (true-scale) intervals in radius, and each labeled with the associated horizontal range from the lander. The arcs also bear a circumferential indexing pattern that denotes the directions of the systematic vertical profiles that have been generated at 5°-azimuth intervals around the lander. The azimuthal fiducial marks have not been labeled. To have done so would have unnecessarily cluttered the often already busy sheets, and would risk overwriting elevation values. The indexing pattern makes it easy to establish the azimuth values. The marks at positions that are integral multiples of 15° have enhanced length relative to the intermediate pairs of 5° marks.

Elevation values for the contour lines are indicated in meters, relative to the LGN origin. These values are equal to

the LGN z-coordinate associated with the contouring plane. The cardinal points of the compass are indicated on the inner margin of the grid frame; these are relative to a compass rose origin located at the center of the sheet.

At the sides and corners of the sheets will be found pointers to the sheet designators for the neighboring sheets. The pointers are present regardless of whether or not the particular neighbor has been created.

A tabulation, immediately preceding the collection of elevation contour sheets indicates both the complement of sheets that have been generated for each of the scales, and the page numbers upon which the individual sheets may be found.

3.3.1 Idiosyncrasies of the Contour Maps

The contour maps, generated as they are from surface based cameras located approximately 0.82 m apart and approximately 1.3 m above the Martian surface, exhibit several unorthodox characteristics that bear comment:

- 1) Of course, only those portions of the scene that can viewed by both cameras can be stereoscopically mapped. Much of the near field and some portions of the horizon are obscured by hardware from the view of one or both cameras. Reference to the "Skyline Drawings" contained in Appendix B will reveal the nature and extent of the obscuration. Contour lines, which physically must all be closed, can only rarely be completed on the maps, as generally only relatively short segments can be followed stereoscopically (or for that matter monoscopically) without interruption.
- 2) Only the facing masks of rocks and the more remote undulations can be mapped. Thus, e.g., rocks are developed either as single or multiple partial contours, indicating the portions of the rocks that could be jointly imaged by both cameras. The entire process of map printing was performed by computer. As the subroutine that layed down the elevation values on the maps was a simple one, there are occasions, especially for steeply inclined surfaces, where the elevation values will pile on top of one another, or otherwise be overwritten and become unreadable. In these instances, the elevation values may be determined by reference to the camera perspective representation of the maps, as discussed in sect. 3.2.1.
- 3) The ranging accuracy falls off roughly quadratically with range, deteriorating rapidly as one approaches 100 meters range. This leads to noisy data at larger ranges (as discussed in further detail below, and in Appendices D and E).
- 4) The elevation contour intervals are variable with range, reflecting the the deterioration of spatial resolution and ranging accuracy with range.
- 5) A nonlinear coupling was used between the hand-input trackball device and the 3-space mark employed to generate the map data, in order a) to enable the approach to large distances with a reasonable number of rolls of the trackball, and b) to facilitate acquisition of 3-D data at a spatial density reasonably related to that of the 3-D information extractable from the images. The individual points constituting the development of the digital contour lines were generally recorded at a quantum granularity somewhat below the theoretical limit of point-ranging resolution inherent in the imagery, or equivalently that of the raster dis-

play system, which mapped 1:1 to the imaged points. At the larger ranges, where the range uncertainty is relatively great, the quantization of the radial distance becomes conspicuous in the maps. At larger ranges, the maps are characterized by features developing a rather exaggerated "u-shape-like" form in a direction extended toward the lander. This artifact in the data is explained in Appendix E. It results from a combination of the two phenomena, one having to do with the nature of the parametrization of trackball motion relative to Martian 3-space displacement of the ranging mark, and the other with the lack of visual feedback from cursor motion when stretching the limits of ranging resolution. The magnitude of the distortion of remote features is generally comparable with the theoretical ranging error at the distance involved. The primary purpose in working so far from the lander, rather than limiting the mapping to smaller values of range, was to milk the data for whatever it was worth, i.e., to establish elevations and rough range positions for features that could not individually be detailed.

6) It is to be noted that the maps are entirely unedited, being reproduced precisely as they were recorded in the original generation process. We have been encouraged to "smooth" the lines to suppress the quantum jaggies. In response, to this, a simple-minded smoothing has been applied, in that the line elements of the plots represent a connection of the midpoints between recorded data points. This smoothing algorithm had minimal visible effect on the maps due to the fact that at the larger range values, where the effect of smoothing would be most evident, the absolute lateral quantization is much finer than that for the radial quantization. Thus, there would generally be a number of lateral jumps for each radial jump. The result was for there to be almost no visible influence on the data. It is suggested that while a more sophisticated smoothing might be soothing, with familiarity the jaggies will not only not be annoying, but will in fact alert the user to the magnitude of the range uncertainty inherent in the data, thereby discouraging unjustified inferences regarding shape and accuracy.

3.4 Vertical Profile Maps

The vertical profile maps are all plotted within the same 50-cm-frame size (true-scale) that was adopted as the standard for the elevation contour maps. As with the contour maps, the vertical profile maps are also reproduced at half-size. The coordinate system of the vertical profile plots is the same as that for the elevation contour sheets, namely, the Local Gravity-Normal (LGN) system.

Systematic vertical profiles have been generated for both landers at 5°-azimuth intervals around the vertically oriented LGN z-axis. The angular orientation of the profiles is measured clockwise from a reference of 0° in the direction of the LGN negative y-axis (the direction generally backward relative to the lander). Each of the profiles is run from the closest possible

point to the lander to the remote limits of ranging capability. The policy followed for the elevation contours, namely, that the choice of scales should be such that each plotted line should be viewable at resolution on some sheet, was compromised for the profile sheets. The profiles are plotted at three different scales: 1:10, 1:100 and 1:1000. For each scale, the landing site is divided into four quadrants and also into a scale-dependent number of annular zones, the latter corresponding to spans of horizontal range. For any given scale, all of the vertical profiles associated with a given quadrant and span of horizontal range are plotted on the same sheet. This entails plotting the profiles for nineteen different values of azimuth on the same sheet.

The number of annular zones and the span of horizontal range for the annular zones associated with each of the three scales is as follows: For scale 1:10, there are four annular zones in front and three in back, with horizontal range spans of 0-5 m (front only), 5-10 m, 10-15 m, and 15-20 m. For scale 1:100, there are two zones, with spans of 0-50 m, and 50-100 m. For scale 1:1000, there is a single range of 0-500 m. The surface of Mars in range 0-5 m in back of the lander is obscured by lander hardware. Thus there are 26 vertical profile sheets generated for each lander, consisting of seven for each of the forward quadrants and six for each of the back.

Figure 2 shows an example of a half-size version of a vertical profile plot for Viking lander 1. The sheet has been generated for scale 1:20, a span of horizontal range from 0 to 50 m, and span of azimuth from 90° to 180°, inclusive (the left forward quadrant).

The abscissa indicates the horizontal range, in meters. The relative elevation scale of the vertical profiles is indicated in the upper left-hand portion of the sheet; it is 1:1 with the scale for the horizontal range. In order to geometrically separate the various profiles from one another the zeros of reference for the elevation values are displaced from one another along the ordinate. Each is offset vertically from its neighbor by 2 cm (true-scale.) The reference lines are each labeled with the corresponding azimuth values for the profiles. In order to facilitate associating individual profile plots with their corresponding zero elevation reference lines, and additionally to facilitate unambiguous profile tracking at crossovers, three different values of line thickness have been used, in cyclic fashion, in plotting the profiles. The line weight associated with any particular profile is indicated by the thickened portion of the corresponding elevation reference line, in the neighborhood of the first vertical bar inside the plot frame from the left hand side.

No tabulation of the kind preceding the elevation contour maps has been prepared. In the instance of the vertical profiles, the sheets are more conveniently ordered for indexing. Namely, they are first grouped by quadrant in clockwise order starting from the back left quadrant, then within each quadrant by order of decreasing scale number, and within each scale by order of increasing range.

2861 ~~INTERCOMING~~

4. ELEVATION CONTOURS

4.1 Elevation Contour Mosaic Overlay Stereo Pairs

This section contains mosaic stereo pairs of images into which have been overlayed the systematic elevation contour lines. The lines are extremely fine, and as such do not reproduce well in the half-tone figures of this report. The reader interested in seriously working with the map data should consider obtaining photographic print versions of the half-tone overlay images contained in this report (see Appendix F). Once generated, the RDSs — representing the geometric descriptions of the contour and vertical profile data sets — are no longer directly coupled to the particular mosaics used for their generation, e.g. morning, noon, or evening. For logistical reasons related to the scope of the project, all of the contour and vertical profile RDSs have been overlayed uniformly only into morning mosaics. However, both morning and evening mosaics were used in generating the RDSs. Since the individual image segment windows that went into the various mosaics were not the same, and furthermore since the sampler arm rested in various positions when the contributing image segments were acquired, there will be instances where small regions of an overlay will exhibit contour/profile lines in the absence of underlying imagery, and vice versa. Additionally, the times of day of the overlay mosaic will not always be the

most favorable for viewing either the topography in stereo, or for showing the overlay lines in good contrast. The images, as reproduced, represent two-fold enlargements from portions of 8"x10" duplicate negatives. The image identification numbers are included in the figure captions for reference purposes.

A stereo viewer will be required in order for the reader to view the images of this report stereoscopically. Even with a viewer, not all parts of the images can be stereoscopically fused. Difficulties arise in the near field to either side of center, due to the relative vertical parallax, local relative rotation, and projected size differences. Kenneth L. Jones has created sets of stereo images, call VLSTEREO, that suppress the vertical parallax differentials in first approximation [Levinthal, 1980]. However, not all of the elevation contours and vertical profile overlays have been created in the VLSTEREO form. Thus, rather than present some images in standard form and others in VLSTEREO, the present report contains only the standard form. Readers interested in obtaining photographic quality versions of images contained or referenced in this report, or any of the various other lander image data forms are referred to the report of Levinthal and Jones [Levinthal, 1980] for product descriptions and ordering information.

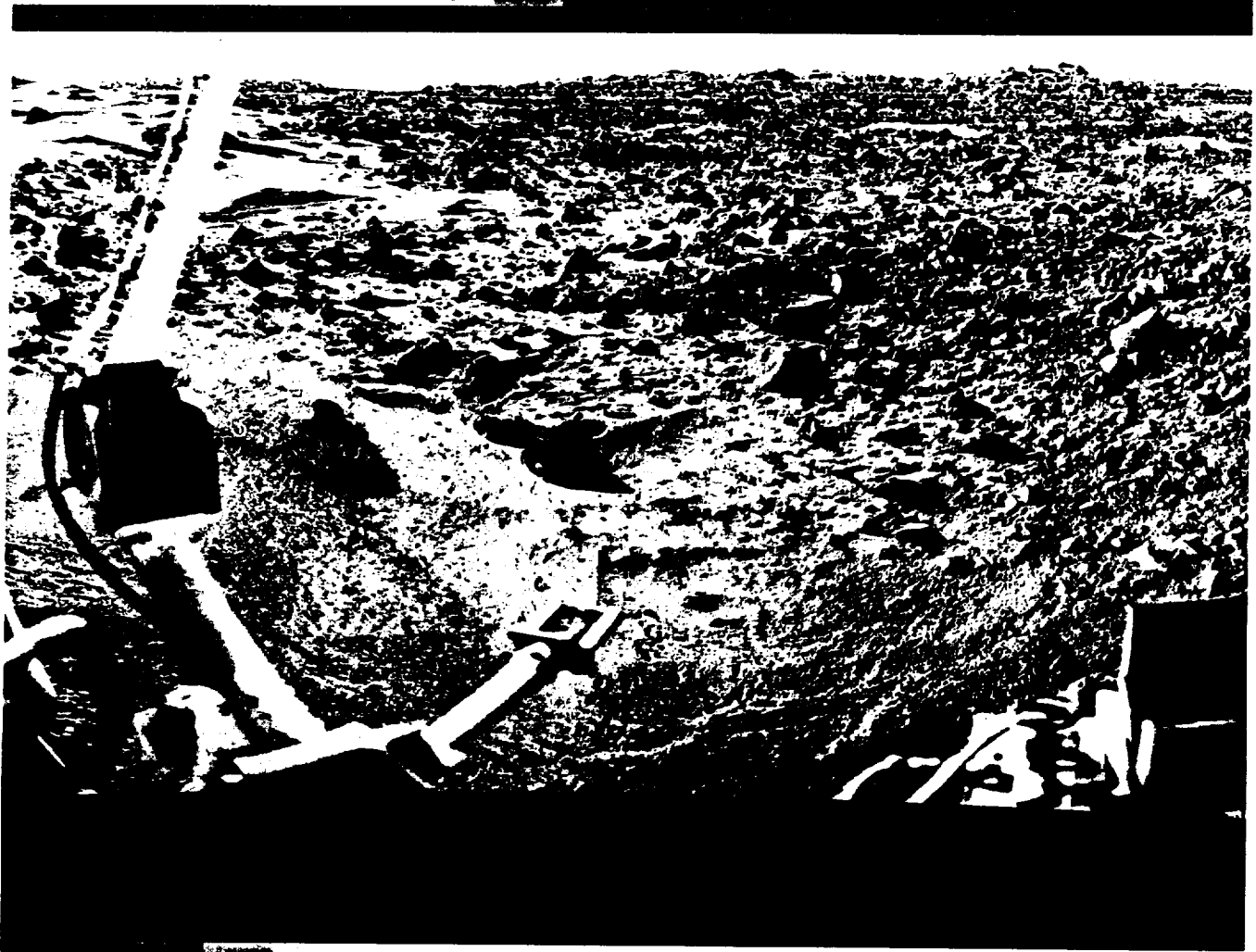


Fig. 3. Elevation Contour Overlay Mosaic: Camera 1 (left), front left quadrant — from IPL PIC ID 79/07/15/100746.

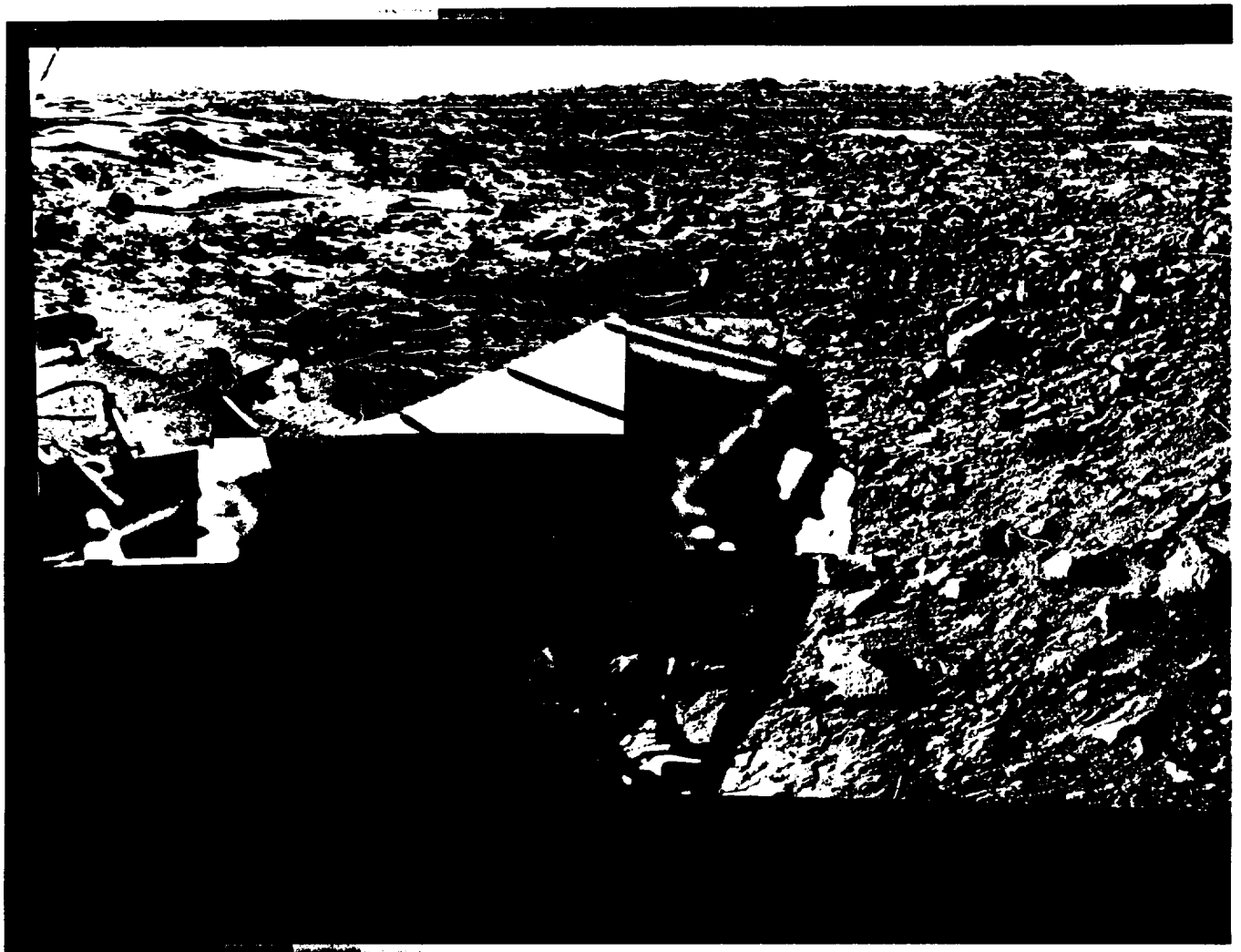


Fig. 4. Elevation Contour Overlay Mosaic: Camera 2 (right), front left quadrant — from IPI PIC ID 79/03/22/035818.

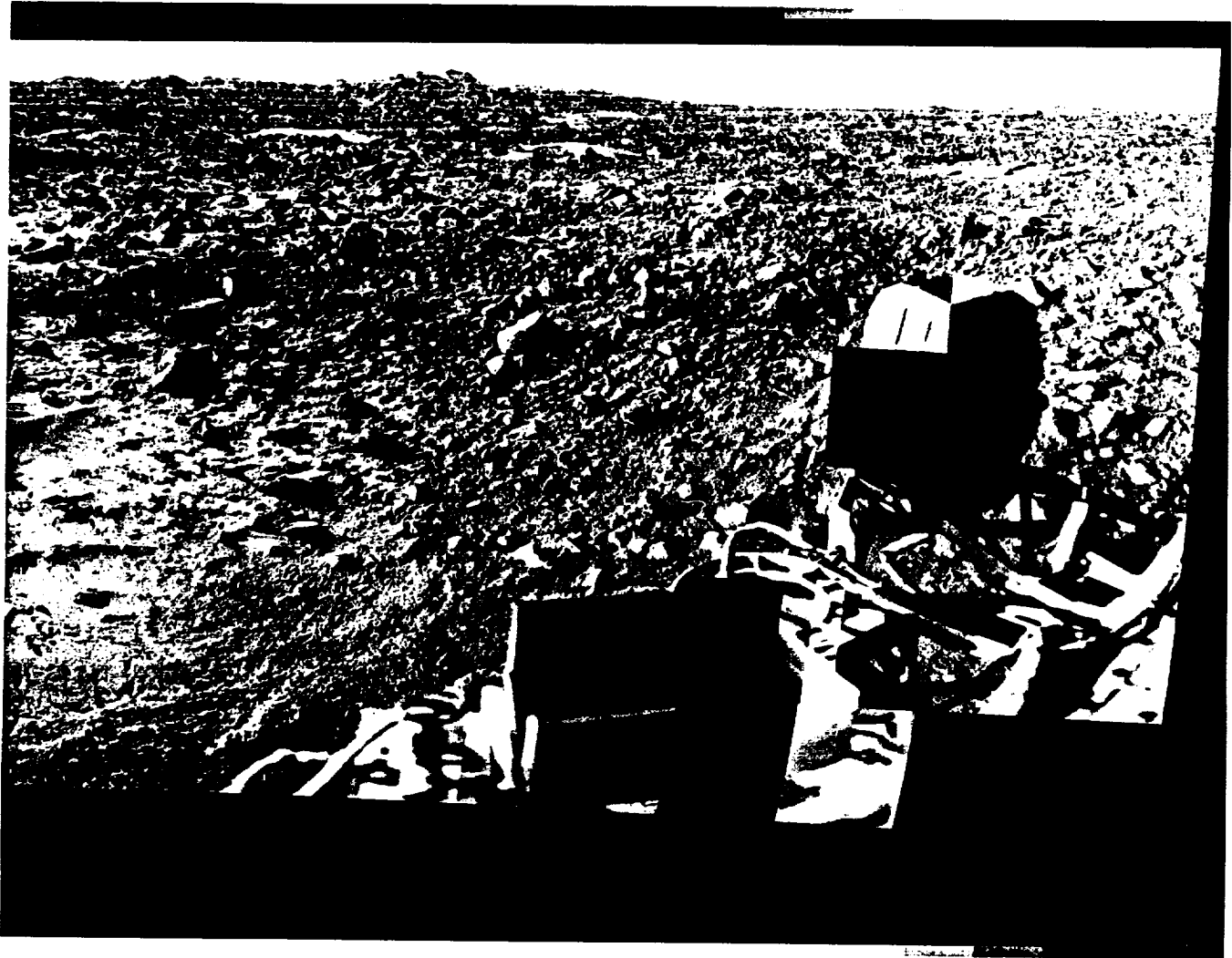


Fig. 5. Elevation Contour Overlay Mosaic: Camera 1 (left), front right quadrant — from IPL PIC ID 79/07/15/100746.



Fig. 6. Elevation Contour Overlay Mosaic: Camera 2 (right), front right quadrant — from IPL PIC ID 79/03/22/035818.

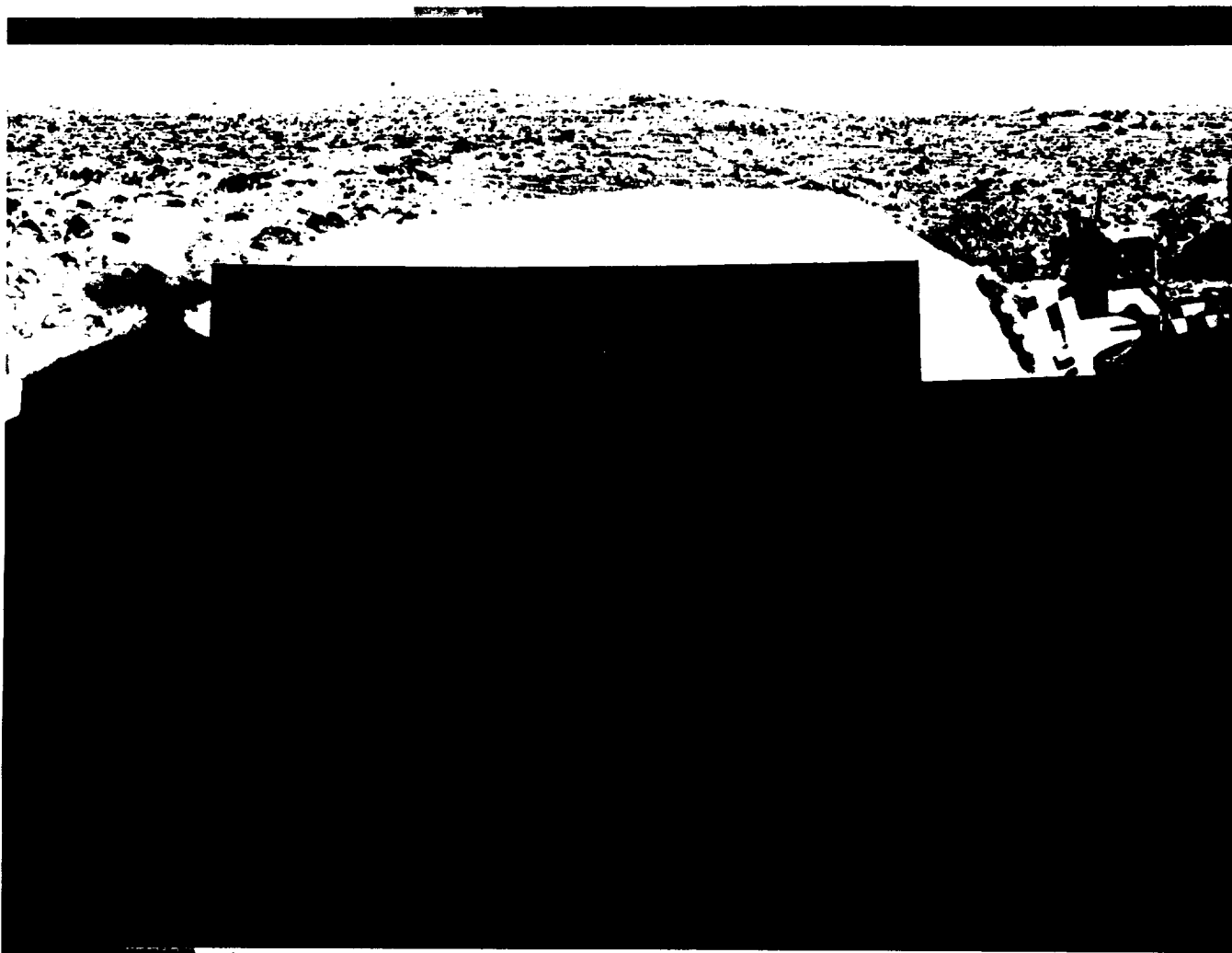


Fig. 7. Elevation Contour Overlay Mosaic: Camera 2 (left), back left quadrant — from IPI PIC ID 79/07/15/112532.

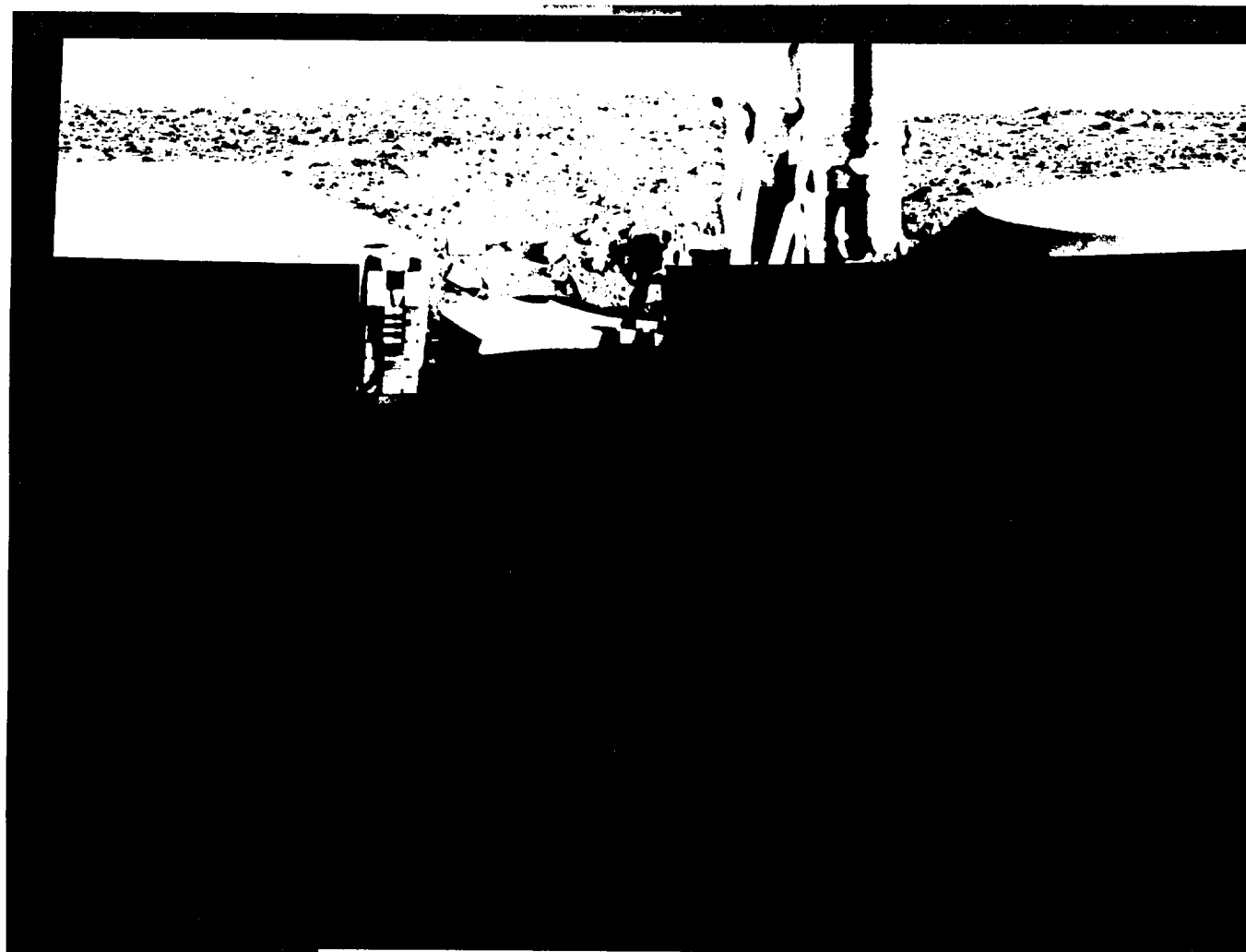


Fig. 8. Elevation Contour Overlay Mosaic: Camera 1 (right), back left quadrant — from IPL PIC ID 79/03/20/034017.

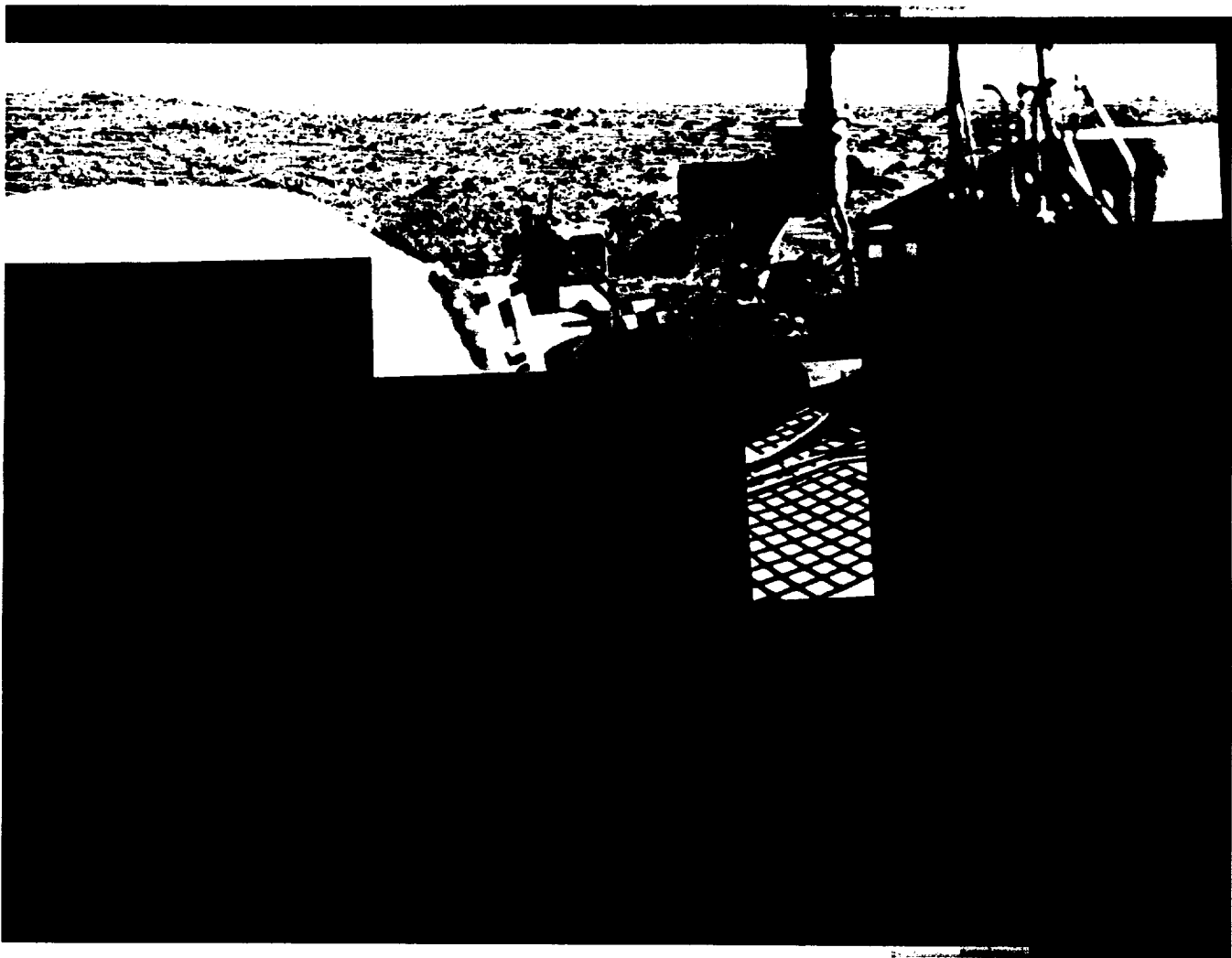


Fig. 9. Elevation Contour Overlay Mosaic: Camera 2 (left), back right quadrant — from IPL PIC ID 79/07/15/112532.

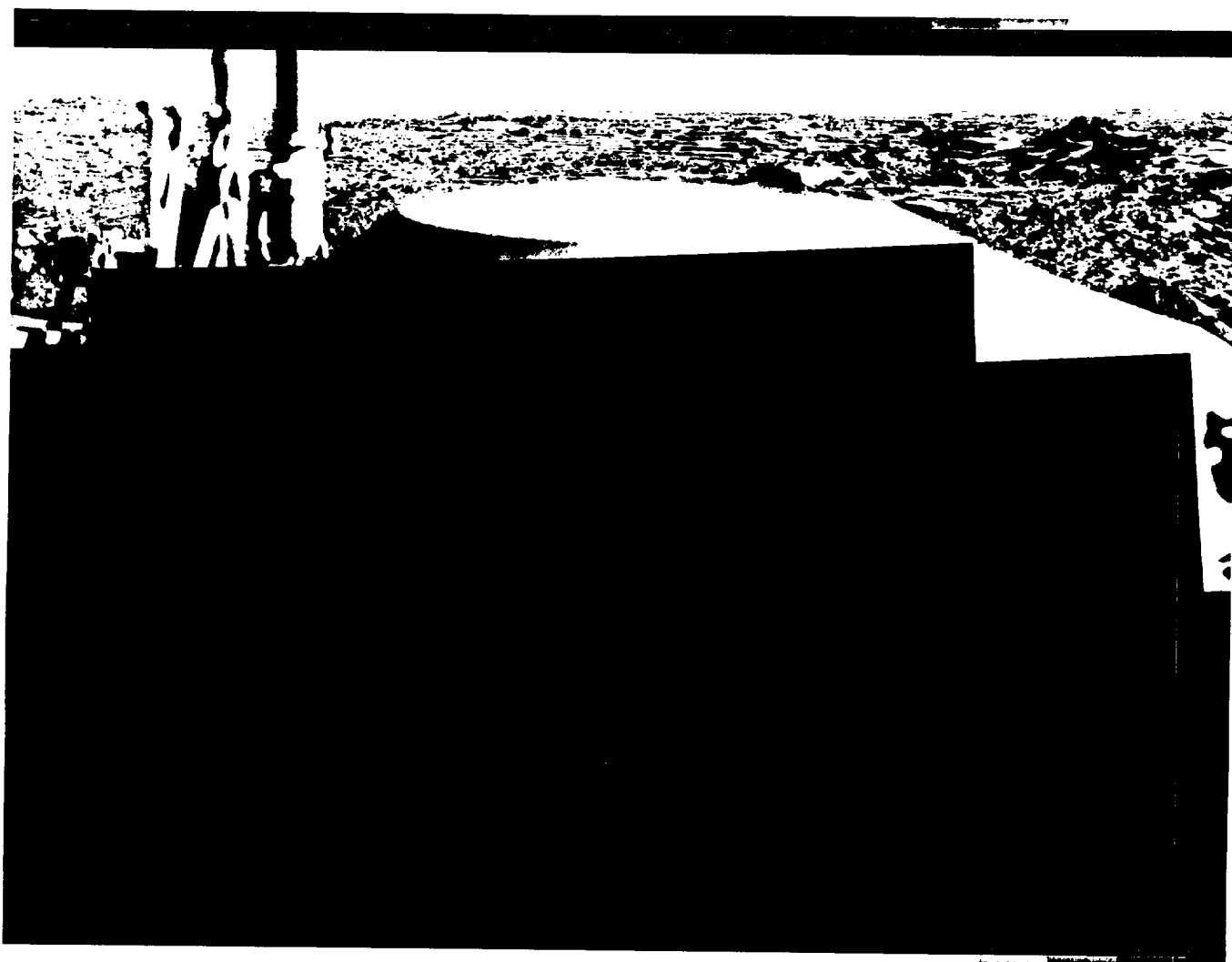


Fig. 10. Elevation Contour Overlay Mosaic: Camera 1 (right), back right quadrant — from IPL PIC ID 79/03/20/034017.

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4.2 Camera Perspective Annotated Elevation Contours — Camera 1

This section contains camera-1 perspective representations of the elevation contour map data. The nature and purpose of these representations has been explained in section 3.2.1. On opposing pages are presented mosaic overlays of the elevation con-

tour lines, and corresponding camera perspective representations of the map data, with the individual lines of the map data annotated with elevation values. The general discussion of section 4.1 applies to the mosaics of this section.



Fig. 11. Elevation Contour Overlay Mosaic: Camera 1 (left), front left quadrant —
from IPL PIC ID 79/07/15/100746.

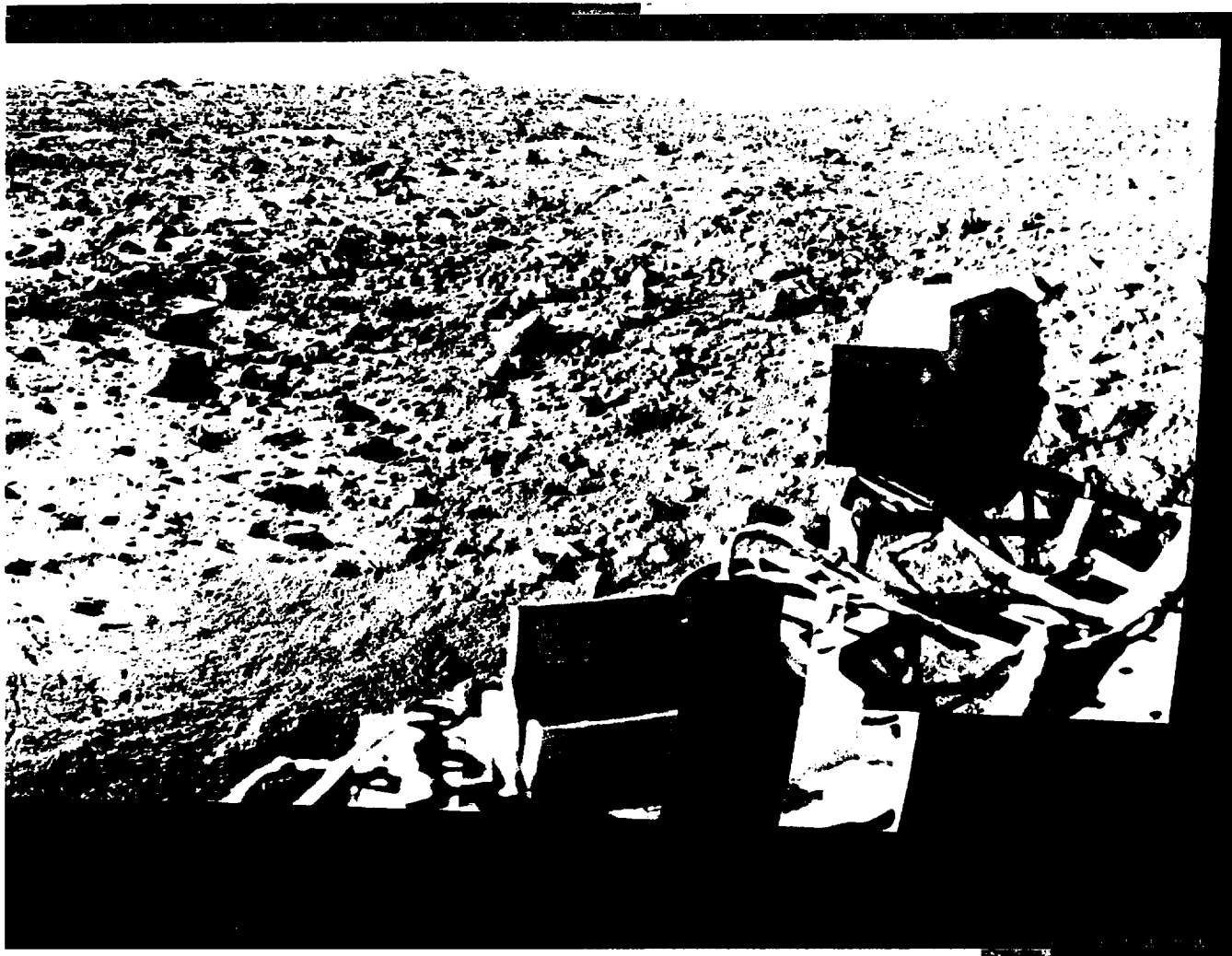


Fig. 13. Elevation Contour Overlay Mosaic: Camera 1 (left), front right quadrant — from IPL PIC ID 79/07/15/100746.

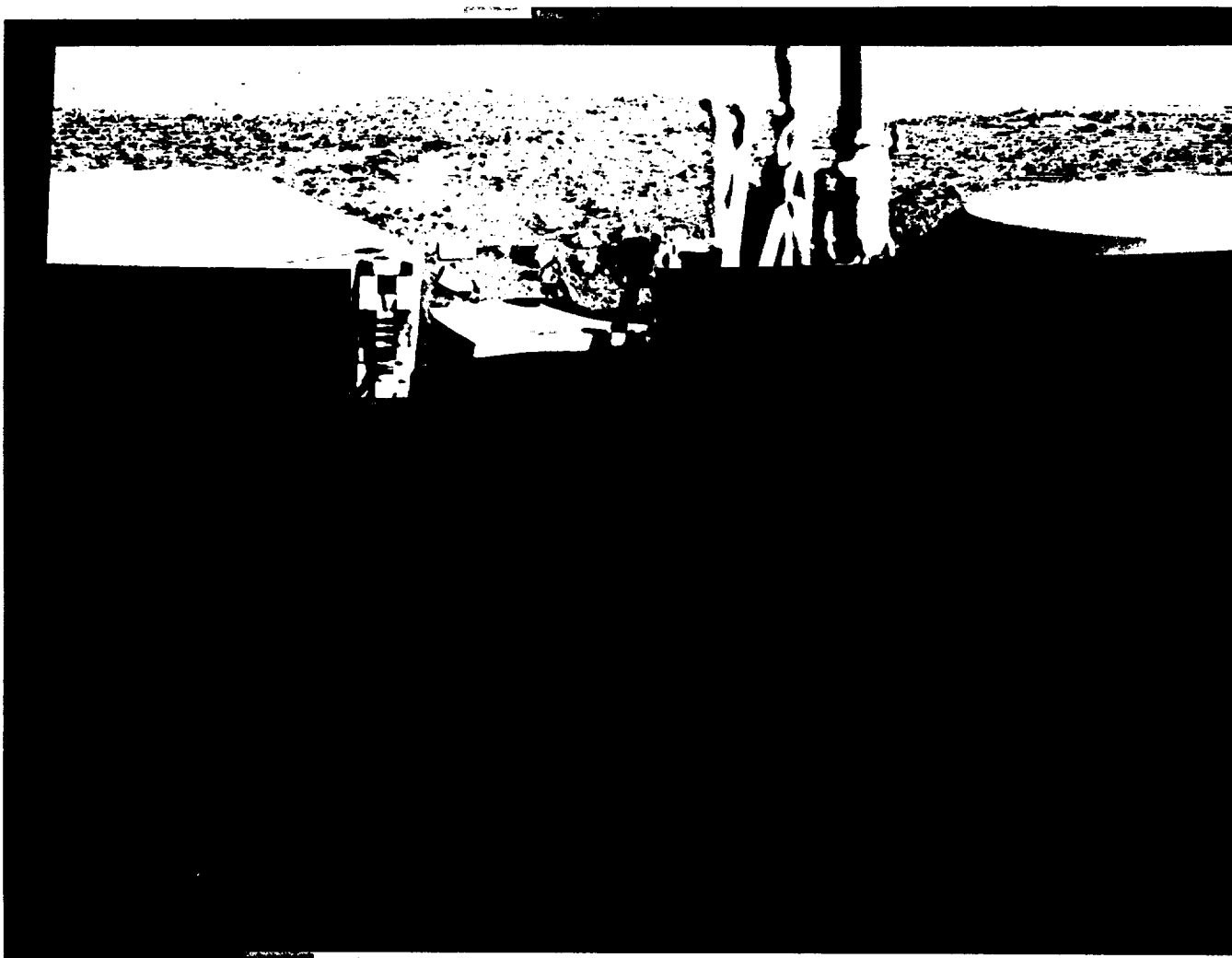


Fig. 15. Elevation Contour Overlay Mosaic: Camera 1 (right), back left quadrant ---
from IPL PIC ID 79/03/20/034017.

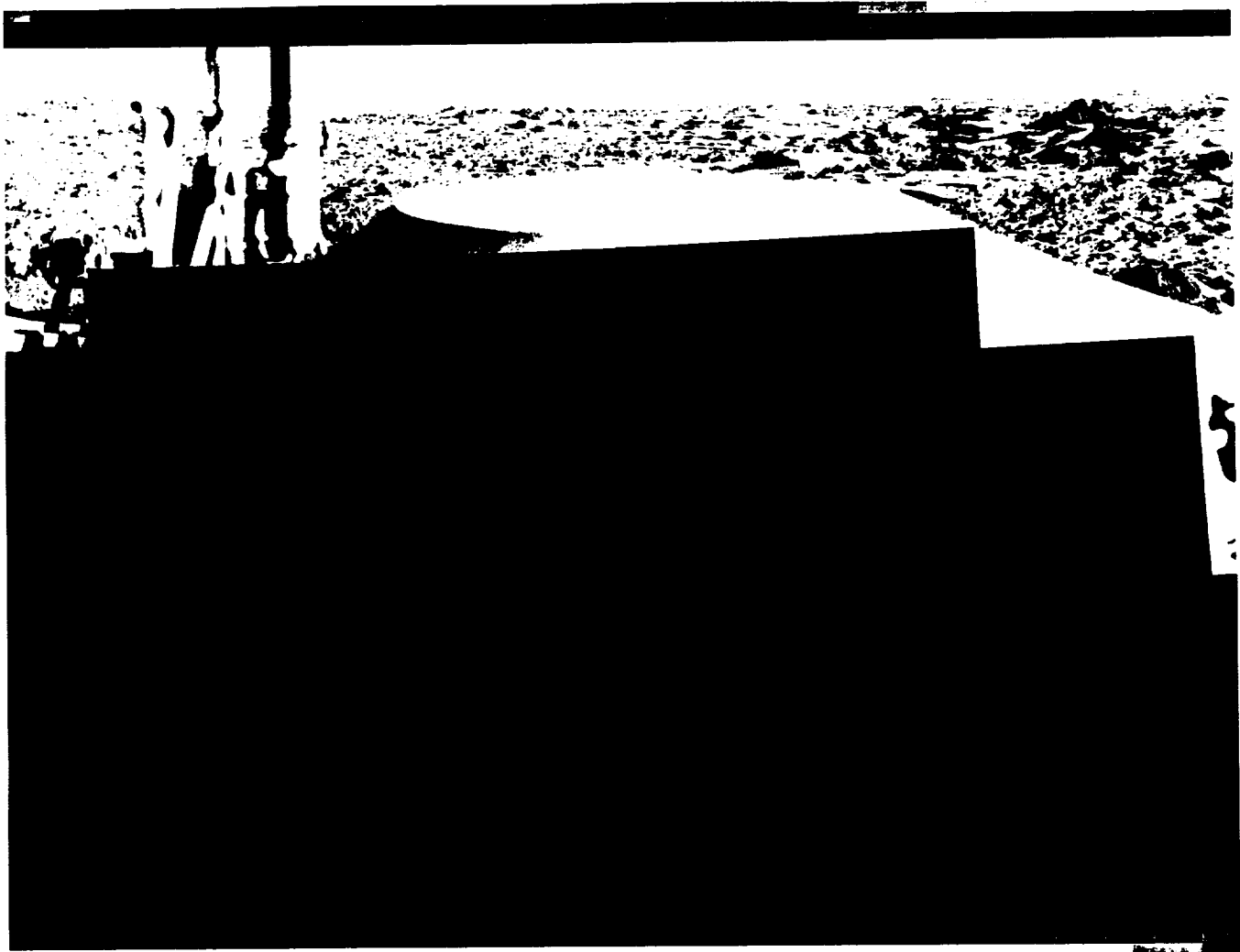


Fig. 17. Elevation Contour Overlay Mosaic: Camera 1 (right), back right quadrant —
from IPL PIC ID 79/03/20/034017.

4.3 Elevation Contour Map Collection

4.3.1 Tabulation of Elevation Contour Map Sheets

The purpose of the following tabulation is to facilitate paging to desired map sheets, as well as to indicate which sheets have been produced. A separate tabular array is presented for

each scale. The rows and columns of the arrays are labeled with the parameters of the sheet array designators (see section 3.3) of the individual map sheets. The column labeling indicates the value of the first parameter of the designator, and the row labeling indicates the second parameter of the designator. The array entries associated with the designators indicate the page number upon which the sheet may be found.

VL1 Contour Sheets		Scale 1:2000						
		-6	-4	-2	+0	+2	+4	+6
+6								
+4								
+2								
+0					31			
-2								
-4								
-6								

VL1 Contour Sheets					Scale 1:100			
	-7	-5	-3	-1	+1	+3	+5	+7
+7								
+5								
+3								
+1				38	39			
-1				40	41			
-3								
-5								
-7								

VI.1 Contour Sheets						Scale 1:5			
	-7	-5	-3	-1	+1	+3	+5	+7	
+7									
+5									
+3			75	76	77	78			
+1			79	80	81	82			
-1									
-3									
-5									
-7									

VL1 Contour Sheets		Scale 1:1000						
		-6	-4	-2	+0	+2	+4	+6
+6								
+4								
+2								
+0					32			
-2								
-4								
-6								

VL1 Contour Sheets					Scale 1:50			
	-7	-5	-3	-1	+1	+3	+5	+7
+7								
+5								
+3								
+1				42	43			
-1				44	45			
-3								
-5								
-7								

VL1 Contour Sheets					Scale 1:2			
	-7	-5	-3	-1	+1	+3	+5	+7
+7								
+5		83	84		85	86		
+3		87	88			89	90	
+1								
-1								
-3								
-5								
-7								

VL1 Contour Sheets		Scale 1:500						
		-6	-4	-2	+0	+2	+4	+6
+6								
+4								
+2								
+0					33			
-2								
-4								
-6								

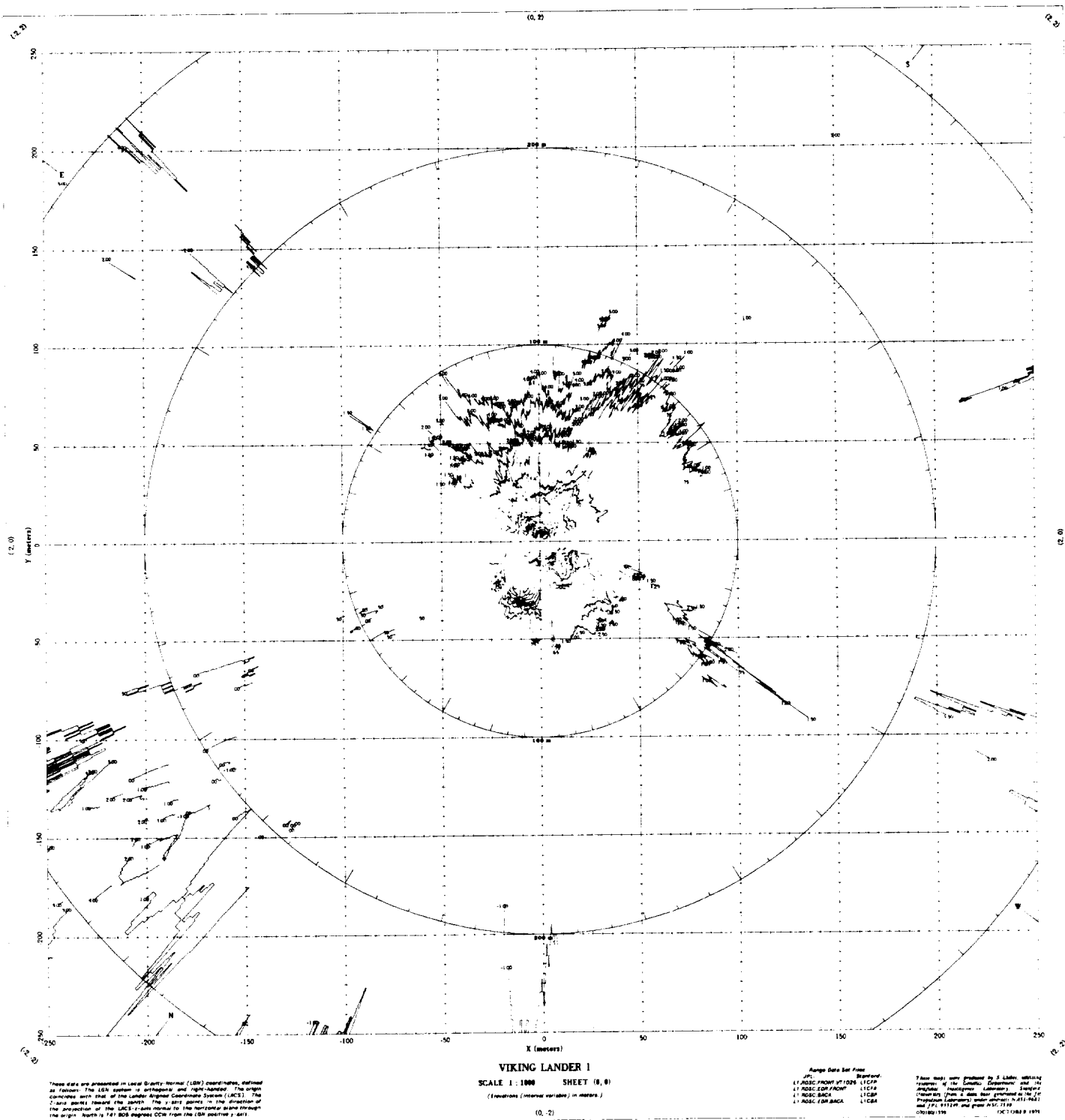
VL1 Contour Sheets					Scale 1:20			
	-7	-5	-3	-1	+1	+3	+5	+7
+7								
+5								
+3			46	47	48	49		
+1			50	51	52	53		
-1			54	55	56	57		
-3			58	59	60	61		
-5								
-7								

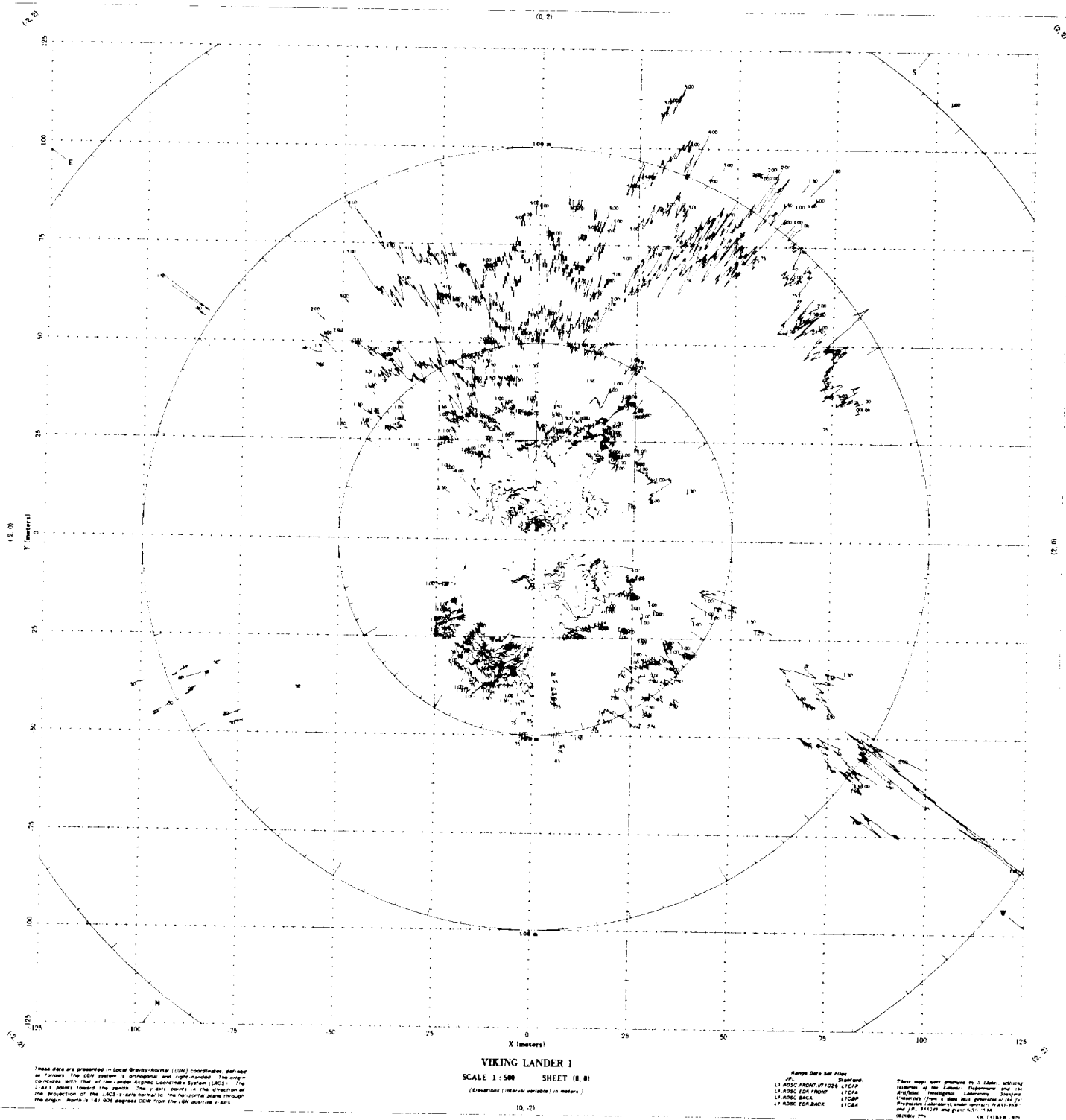
VL1 Contour Sheets						Scale 1:1			
	-7	-5	-3	-1	+1	+3	+5	+7	
+7	91						92	93	
+5								94	
+3									
+1									
-1									
-3									
-5									
-7									

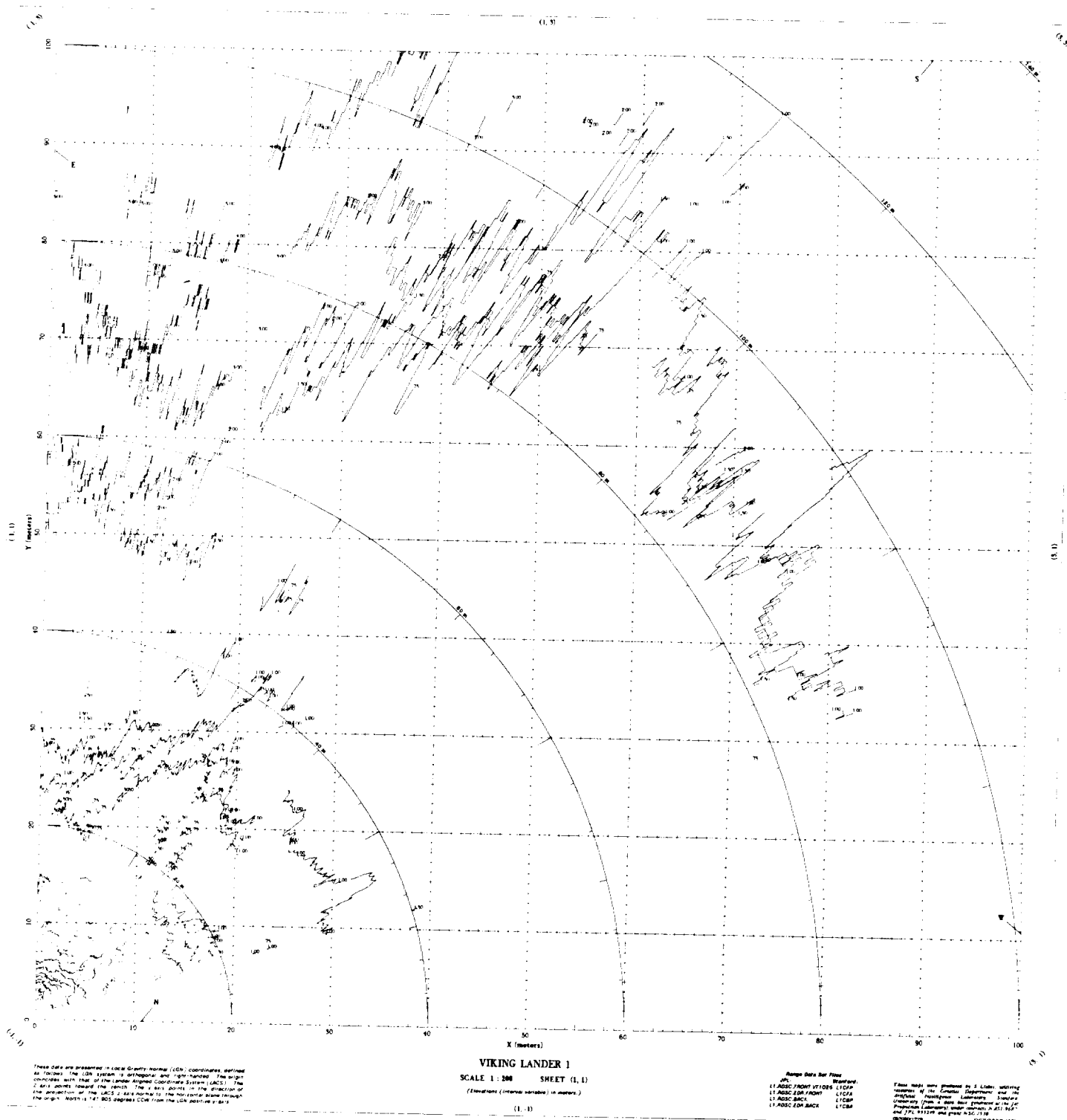
VL1 Contour Sheets					Scale 1:200			
	-7	-5	-3	-1	+1	+3	+5	+7
+7								
+5								
+3								
+1				34	35			
-1				36	37			
-3								
-5								
-7								

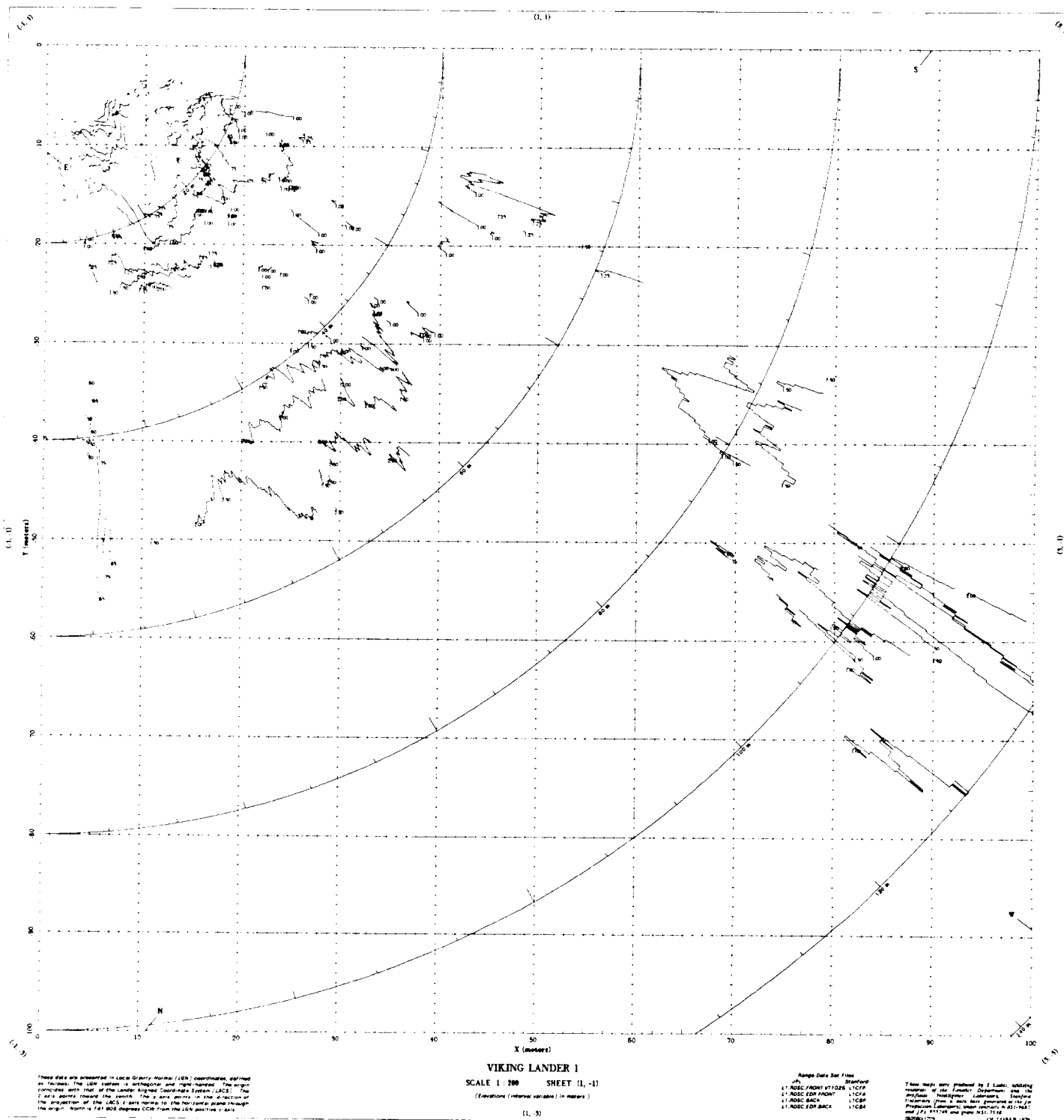
VL1 Contour Sheets					Scale 1:10			
	-7	-5	-3	-1	+1	+3	+5	+7
+7								
+5								
+3			62	63	64	65		
+1			66	67	68	69		
-1						70		
-3			71	72	73	74		
-5								
-7								

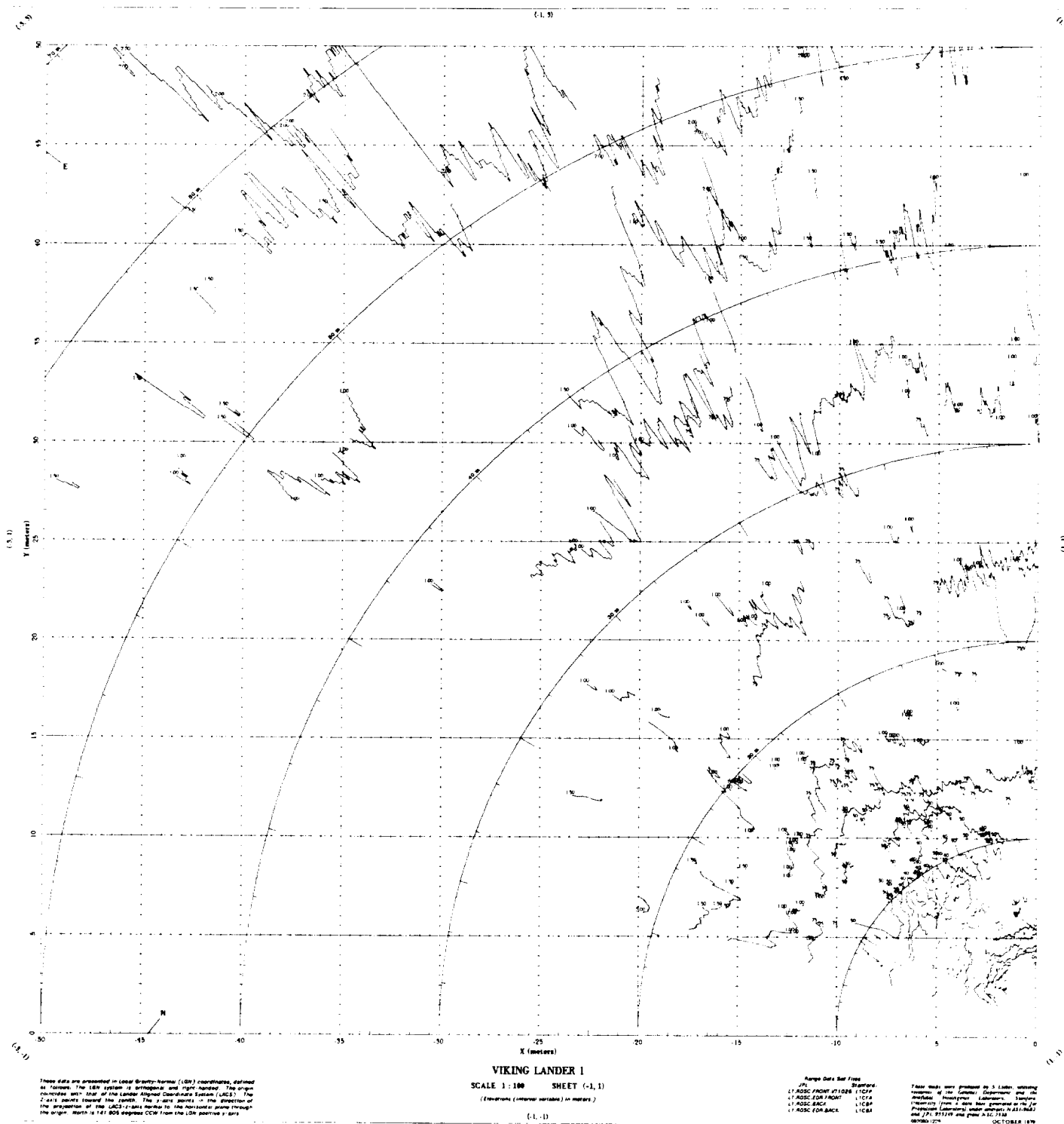
Table 1. Tabulation of Elevation Contour Map Sheets.

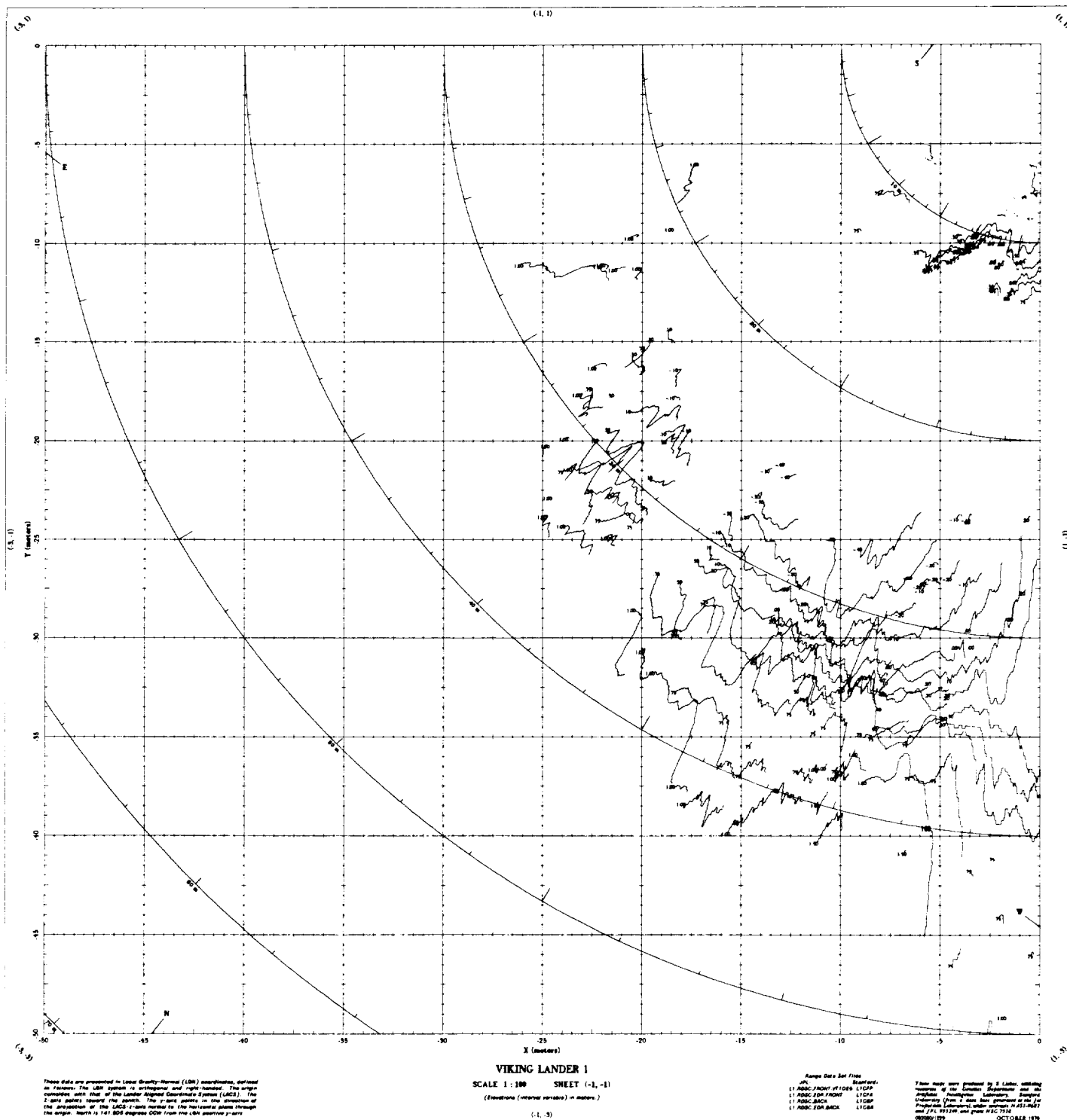


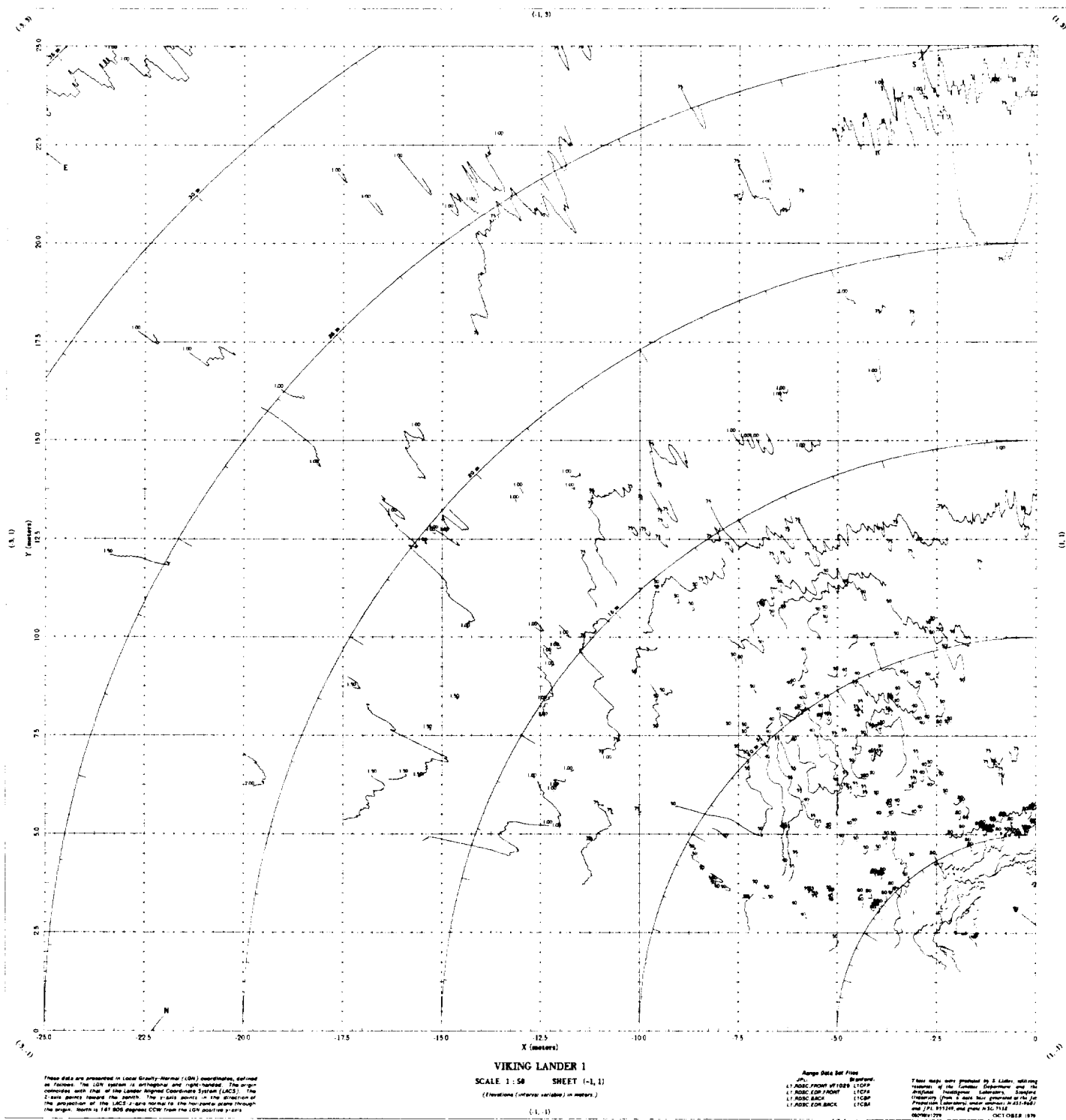


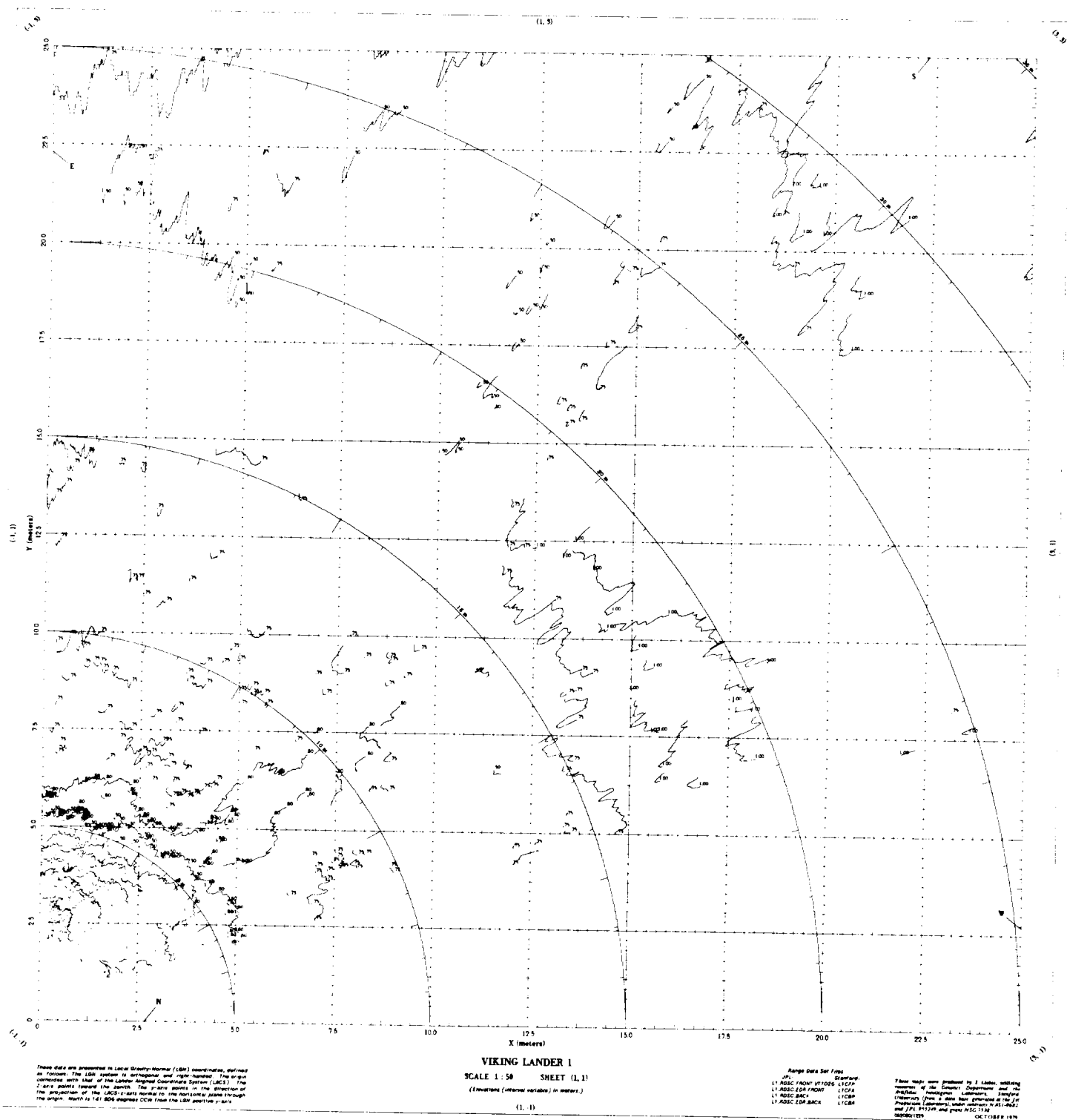


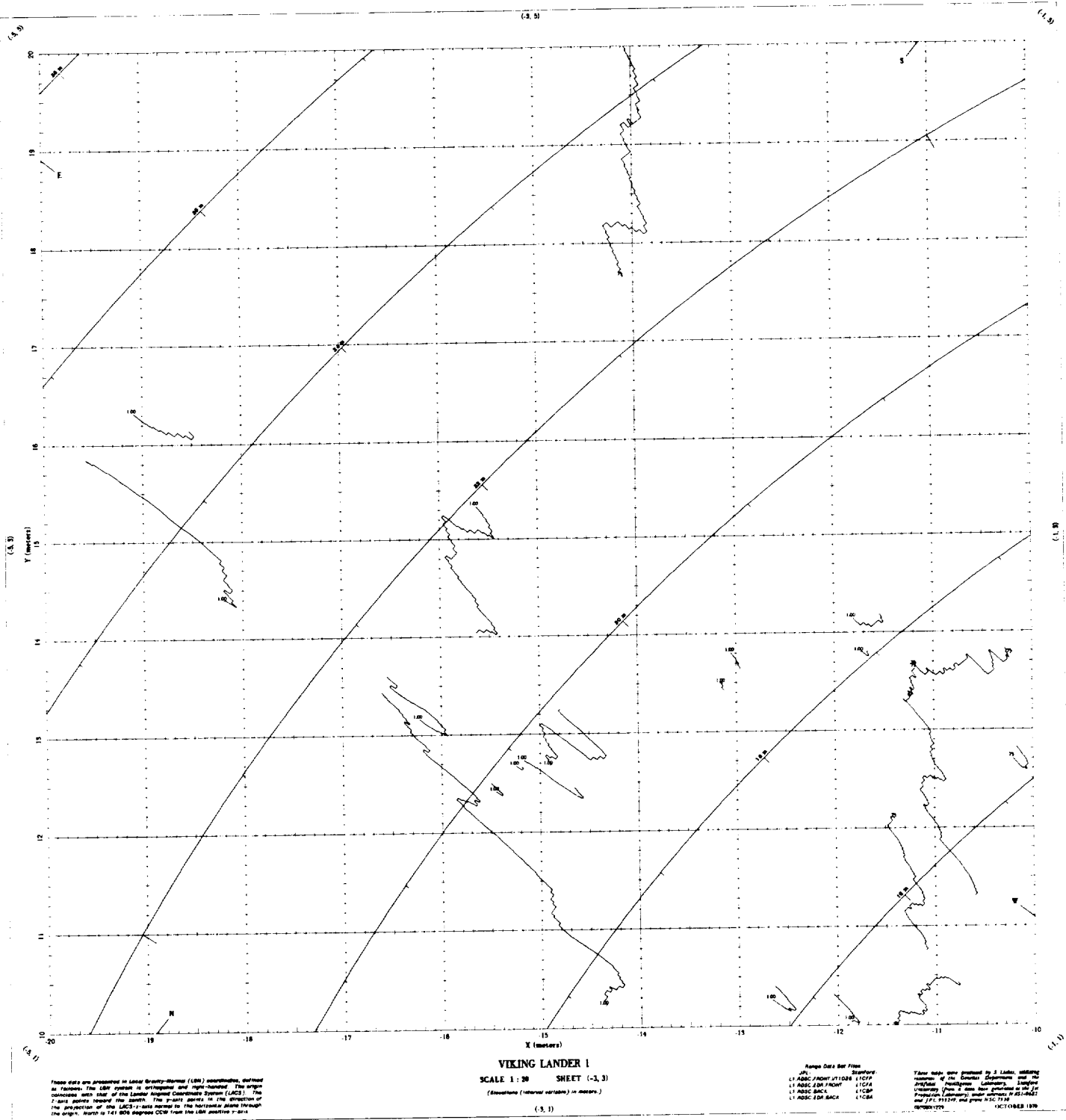


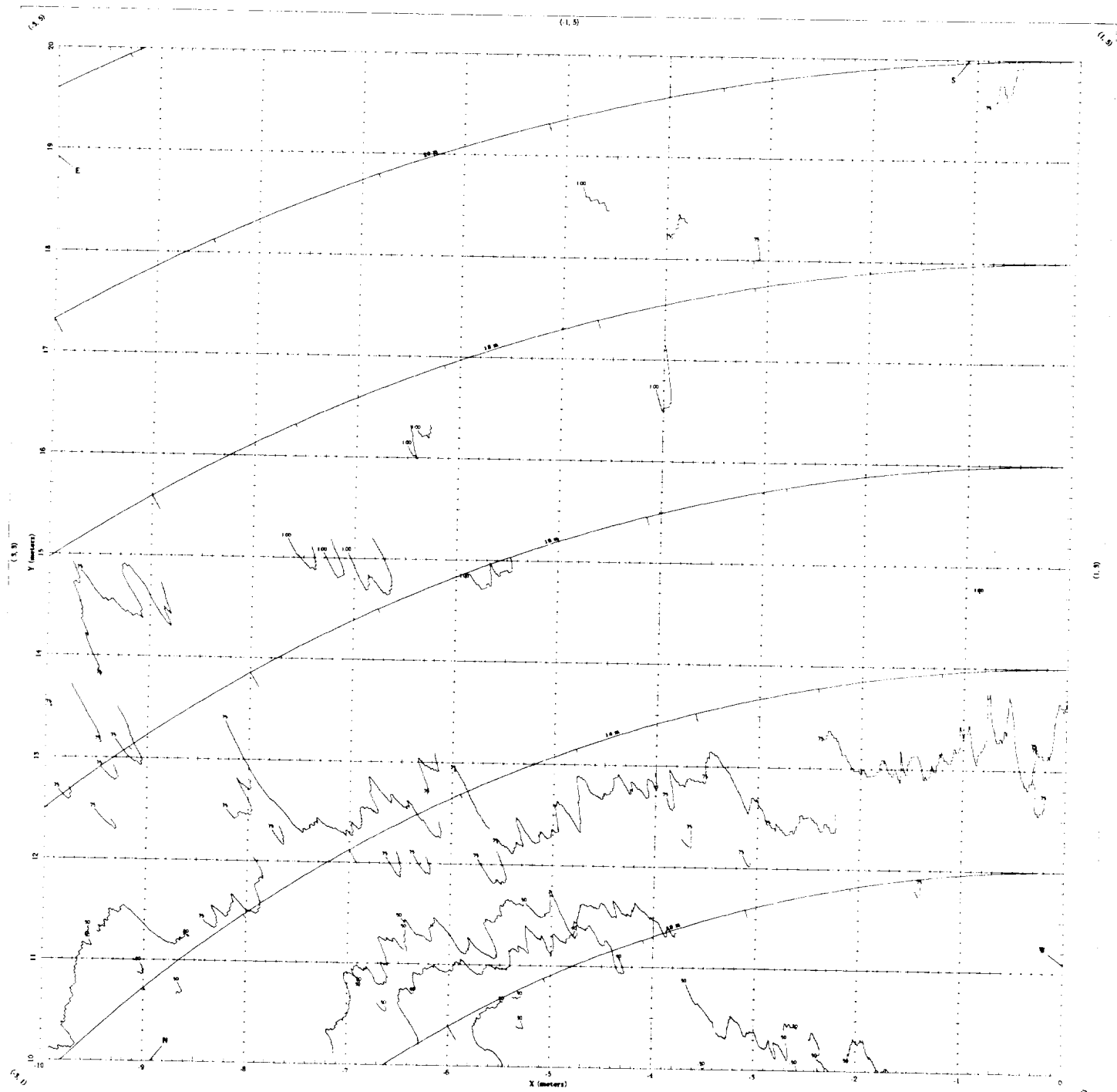










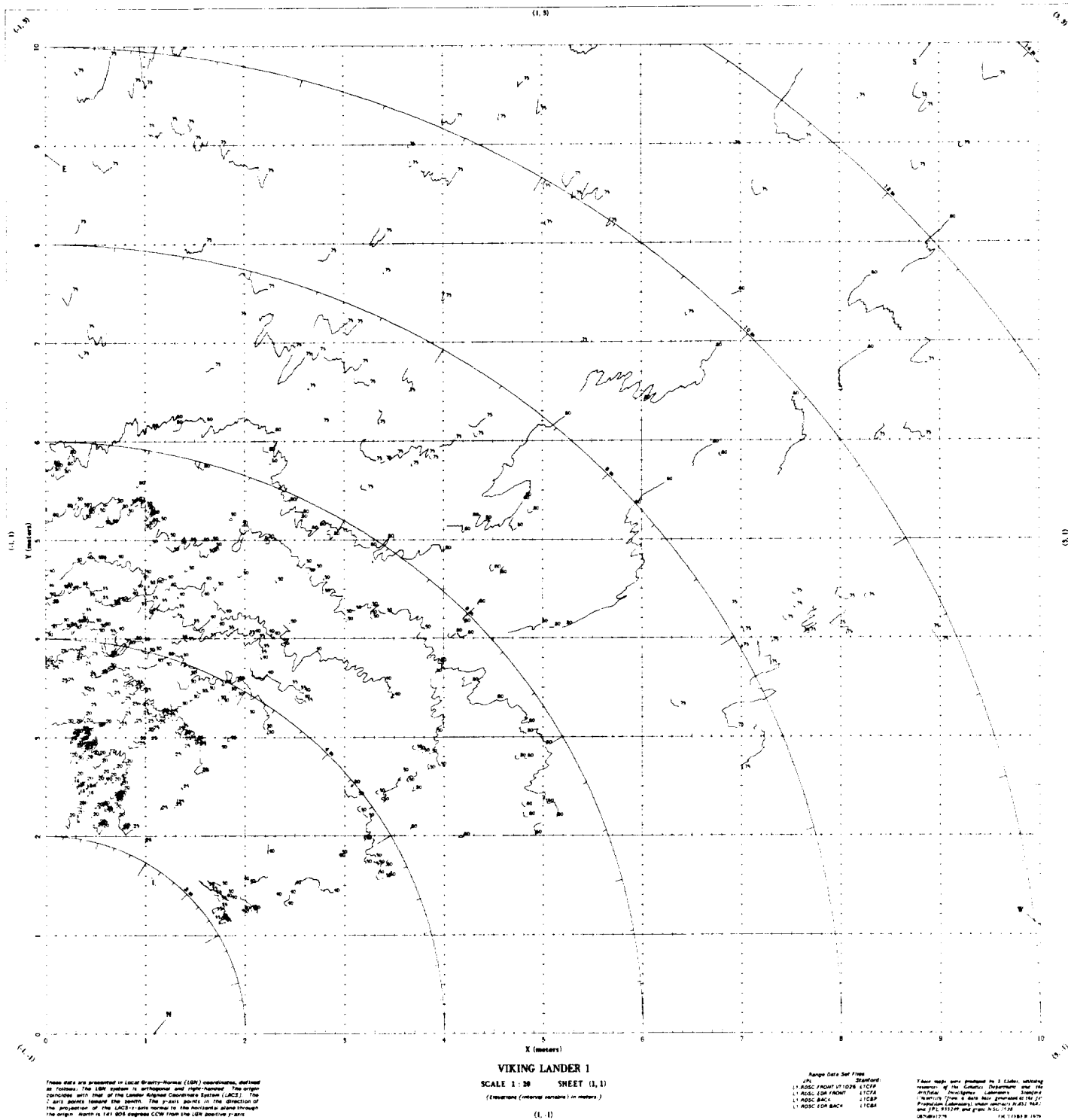


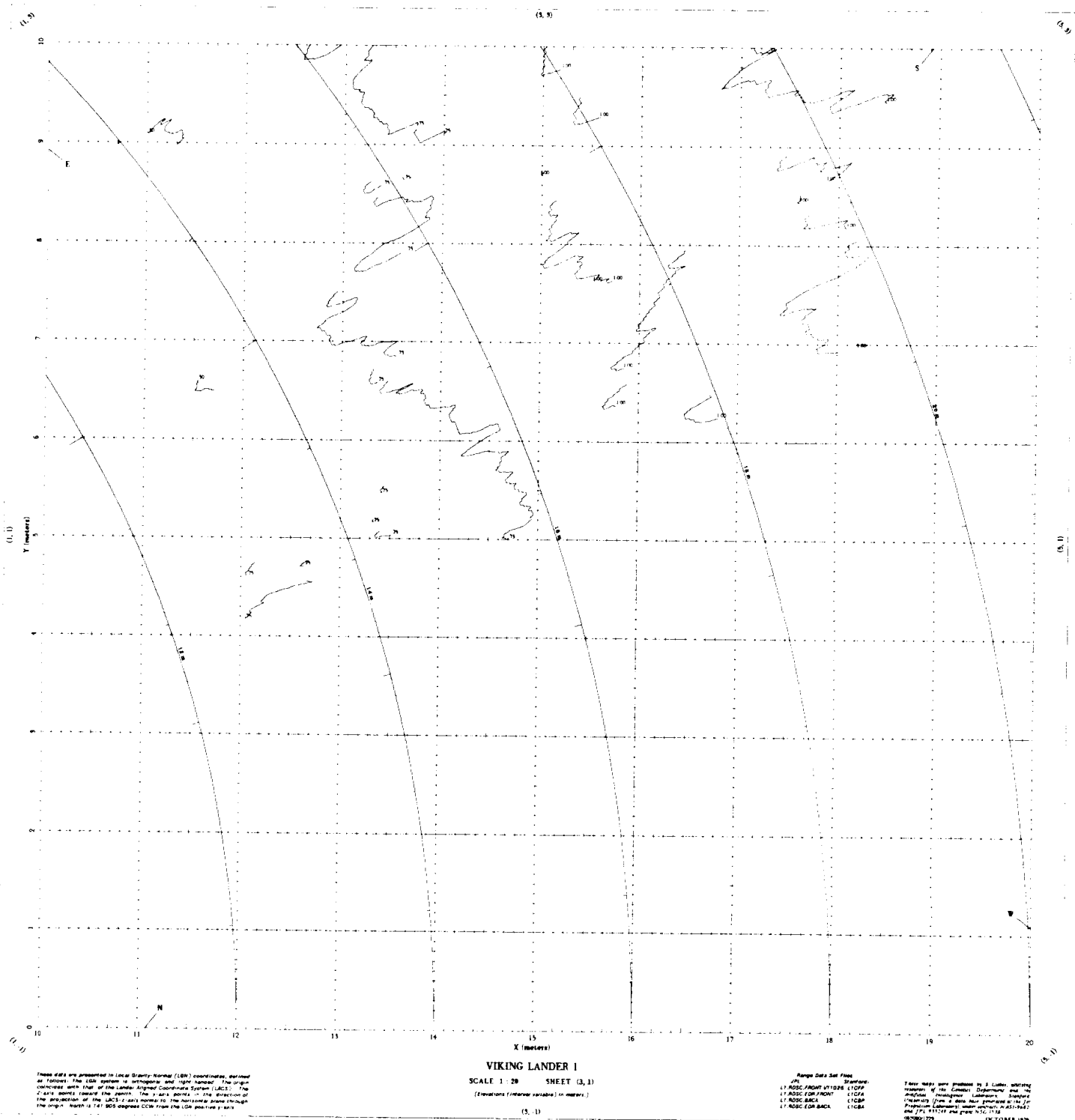
These data are presented in Local Gravity-Normal (LGN) coordinates, defined as follows: The LBR system is orthogonal and right-handed. The origin coincides with that of the Lander Aligned Coordinate System (LACS). The Z-axis points toward the zenith. The Y-axis points in the direction of the intersection of the LACS-Z-axis normal to the horizontal plane through the origin. North is 141.805 degrees clockwise from the LBR positive y-axis.

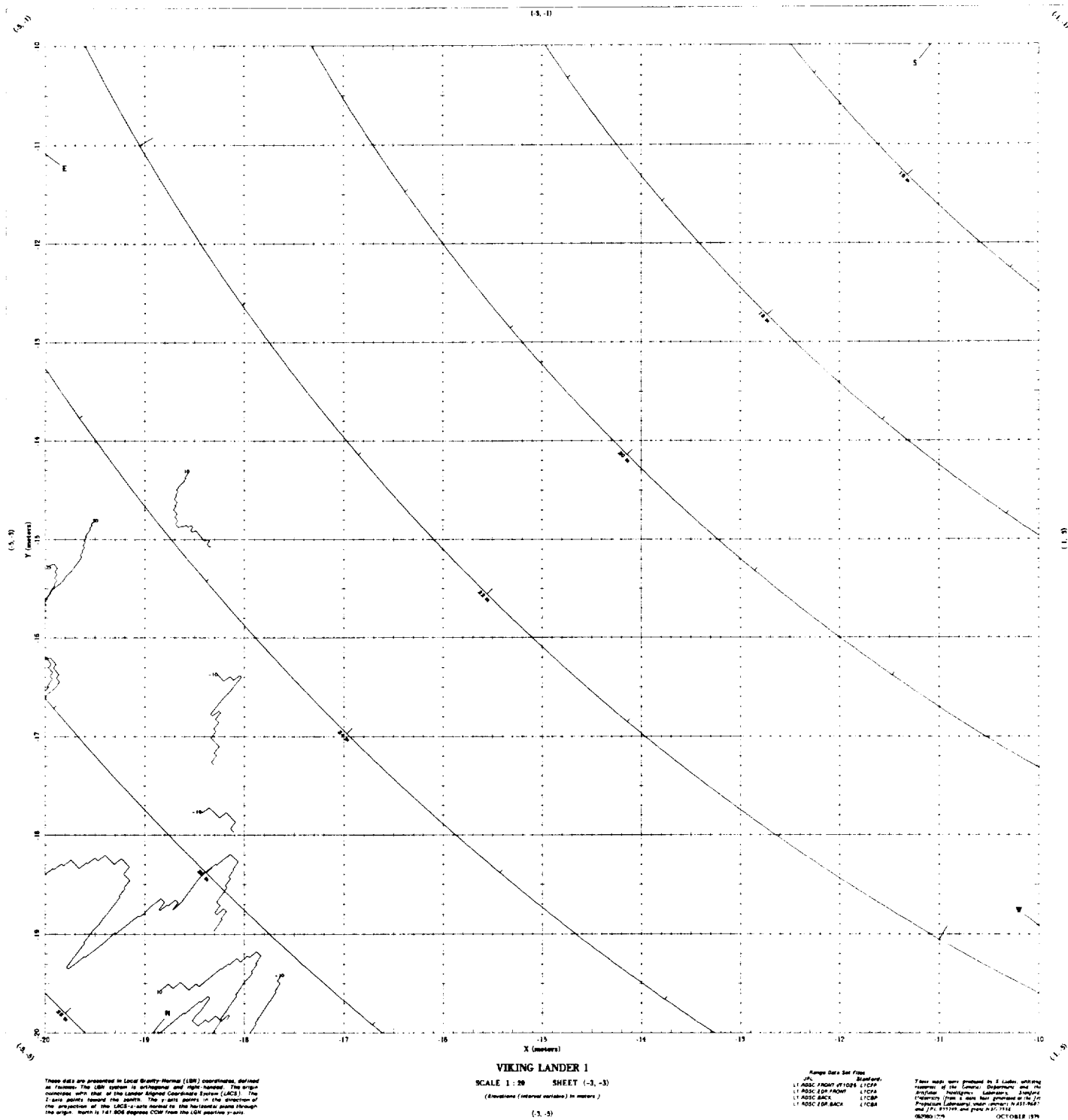
VIKING LANDER 1
SCALE 1:20 SHEET (-1,3)
(Elevations (interpolated) in meters.)
(-1, 1)

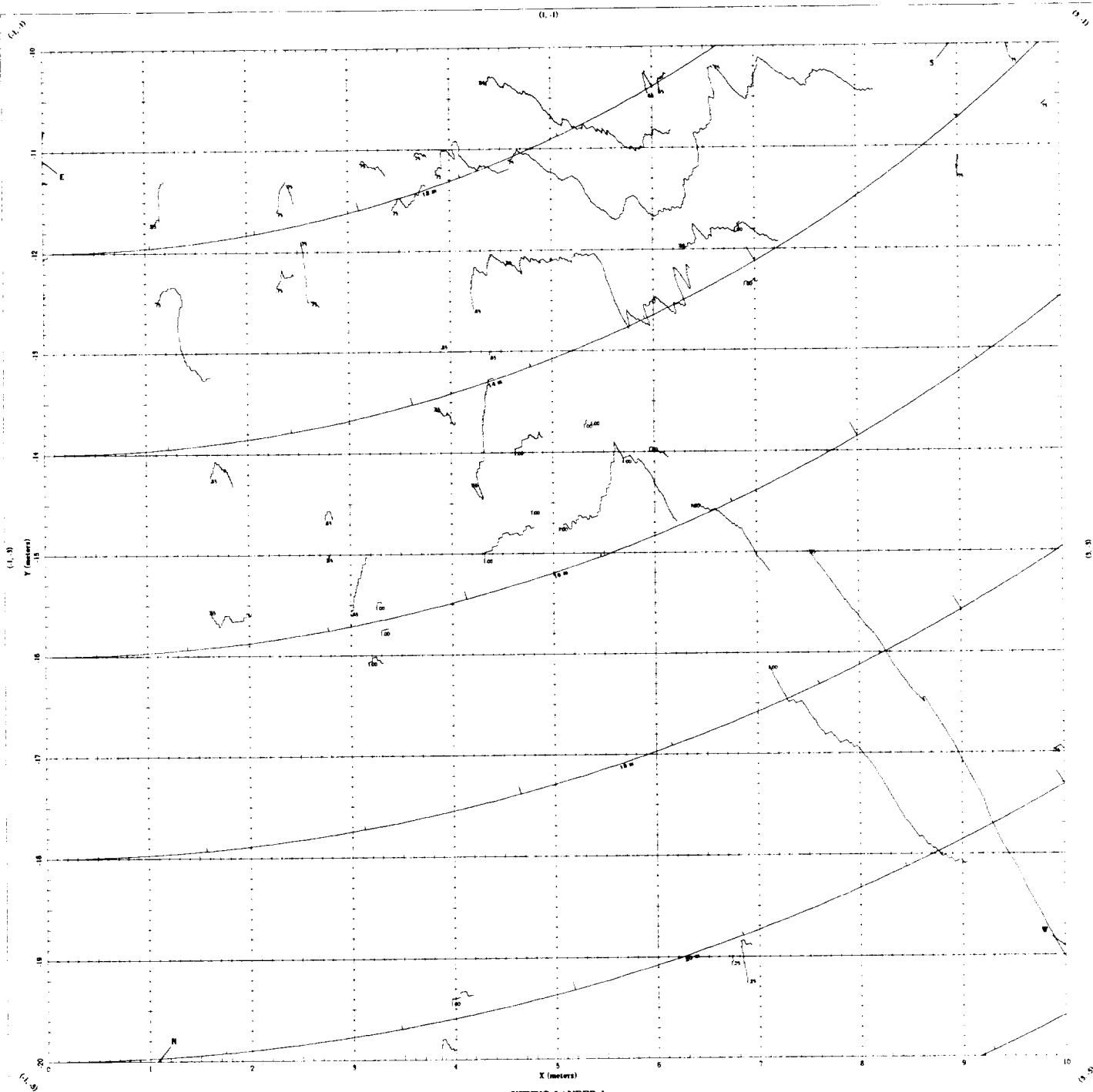
Range Data Set Files
JPL
11.1000 FROM 171025 L1C1P
11.1000 LBR FROM 11024
11.1000 BACK 11C1P
11.1000 LBR BACK 11C1P

These maps were produced by S. Linder, working under the direction of the Center for Space and Terrestrial Sciences, University of California, San Diego, La Jolla, California. The maps were prepared using the JPL 8111-15 and 8111-16 maps N10-713 and N10-714.









These data are presented in Linear Shaded Relief (LSR) representation, defined as follows: The LSR is a two-dimensional map. The origin coincides with that of the Lander Azimuth Coordinate System (LACS). The X-axis points toward the south. The Y-axis points in the direction of the projection of the LACS Z-axis normal to the horizontal plane through the origin. North is 141.808 degrees CCW from the LSR pointing north.

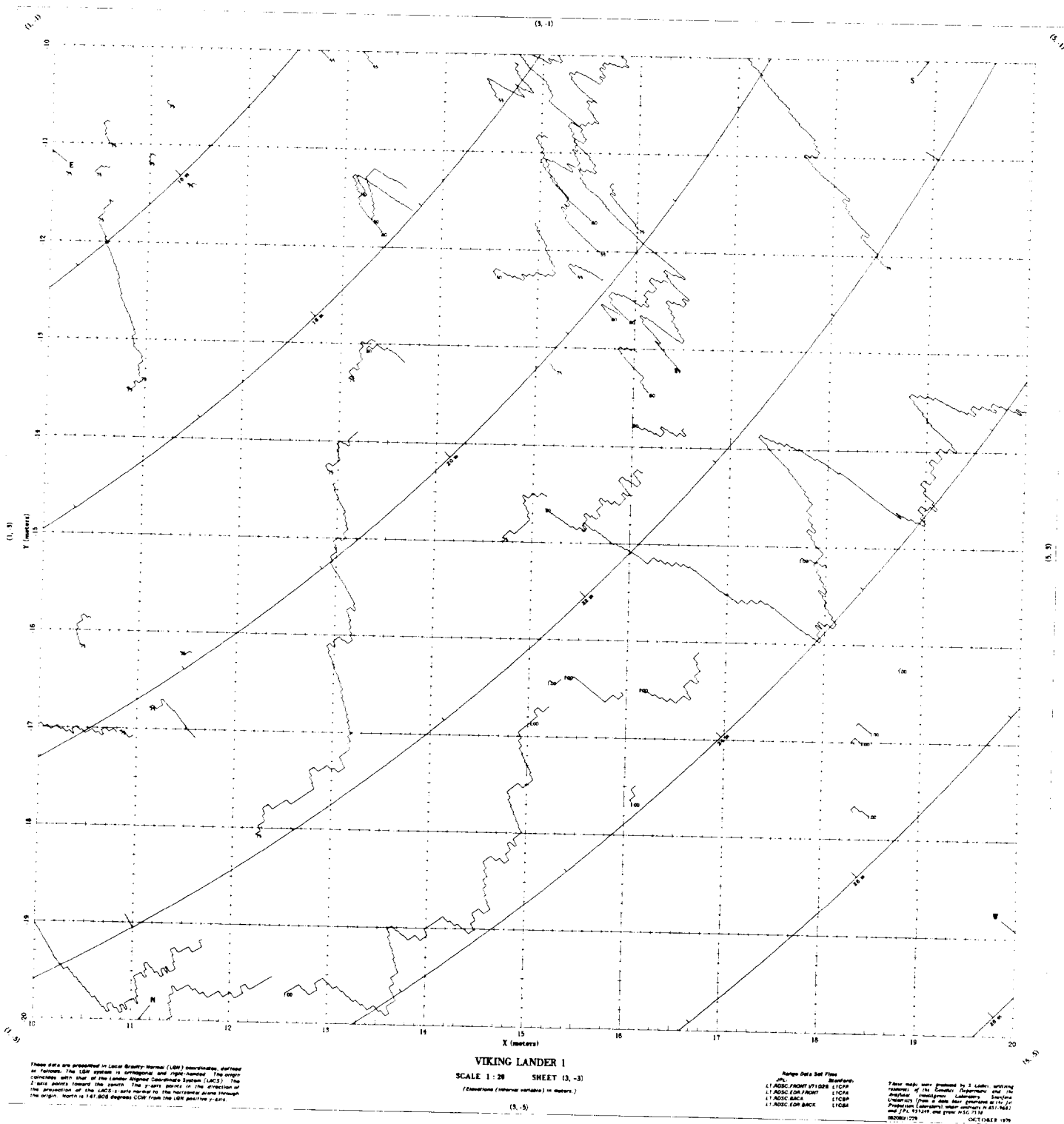
VIKING LANDER 1
SCALE 1:20 SHEET (1, -3)
(Elevations (except variable) in meters)

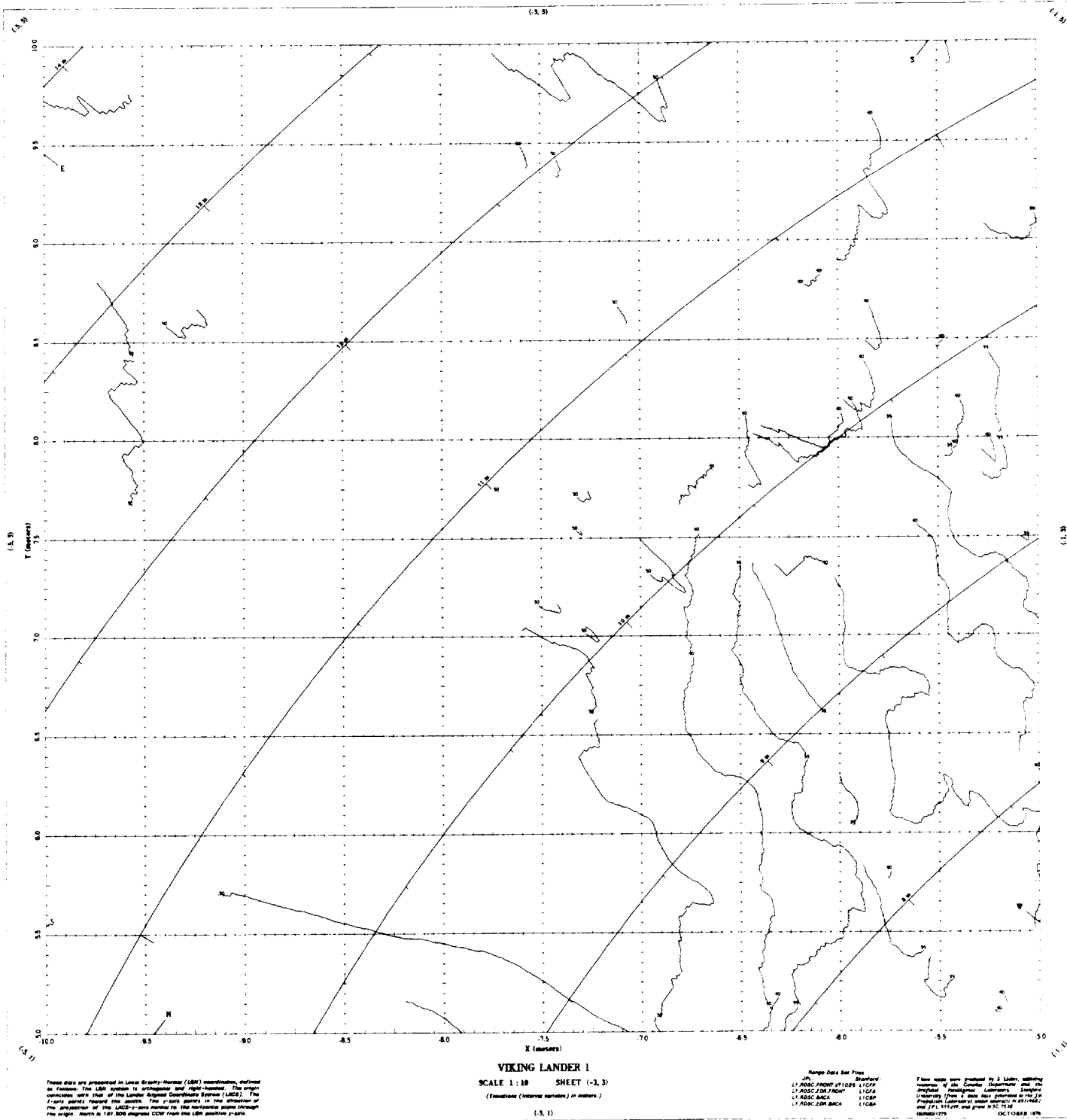
(1, -3)

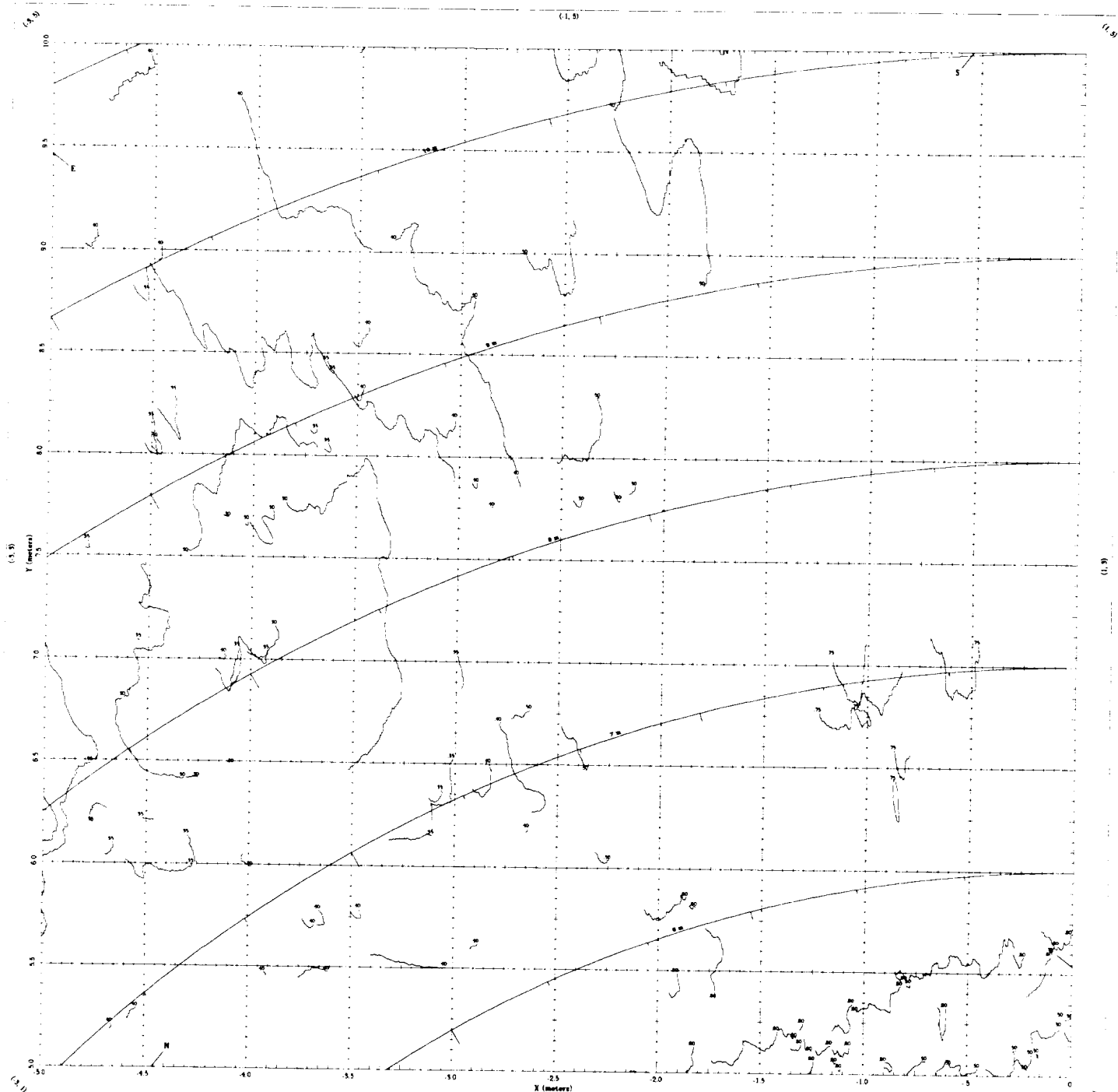
Range Data Set Files
JPL
1) ASOC FRONT VIEWERS (TCPP)
1) ASOC LOR FRONT (TCGA)
1) ASOC BACK (TCBP)
1) ASOC LOR BACK (TCBA)

These maps were produced by 3 Lander Imaging
Systems: the Lander Imaging System (LIS),
the Lander Imaging System (LIS), and the
Lander Imaging System (LIS). The maps were
produced by 3 Lander Imaging Systems: the
Lander Imaging System (LIS), the Lander
Imaging System (LIS), and the Lander Imaging
System (LIS). The maps were produced by 3
Lander Imaging Systems: the Lander Imaging
System (LIS), the Lander Imaging System (LIS),
and the Lander Imaging System (LIS). The maps
were produced by 3 Lander Imaging Systems: the
Lander Imaging System (LIS), the Lander Imaging
System (LIS), and the Lander Imaging System (LIS).

OCTOBER 1976







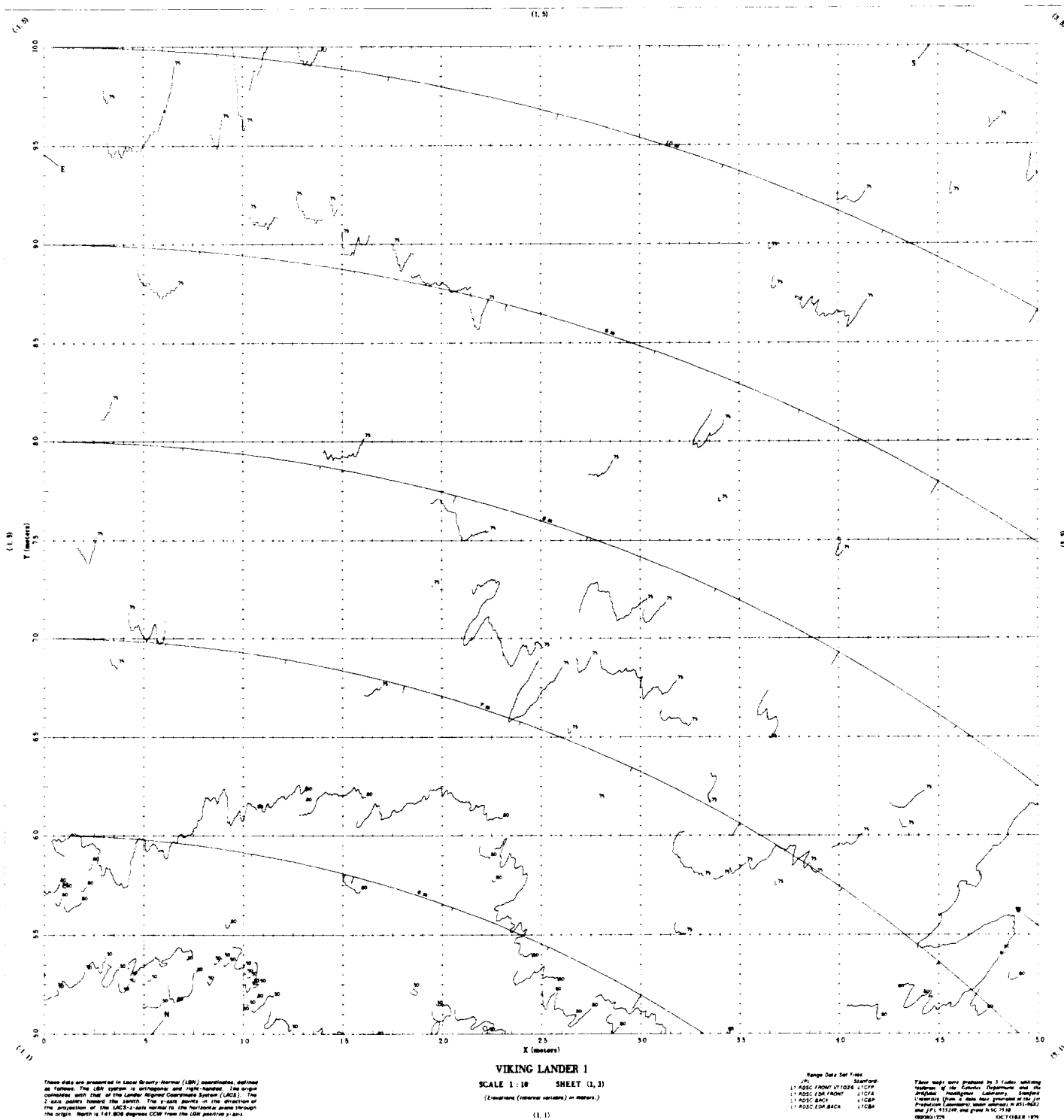
These data are presented in Local Binary-Number (LBN) coordinates, derived as follows. The LBN system is orthogonal and right-handed. The origin coincides with that of the Lander Right Coordinate System (LACS). The Z-axis points toward the zenith. The Y-axis points in the direction of the projection of the LACS Y-axis toward the horizontal plane through the origin. North is 144.808 degrees CCW from the LBN positive Y-axis.

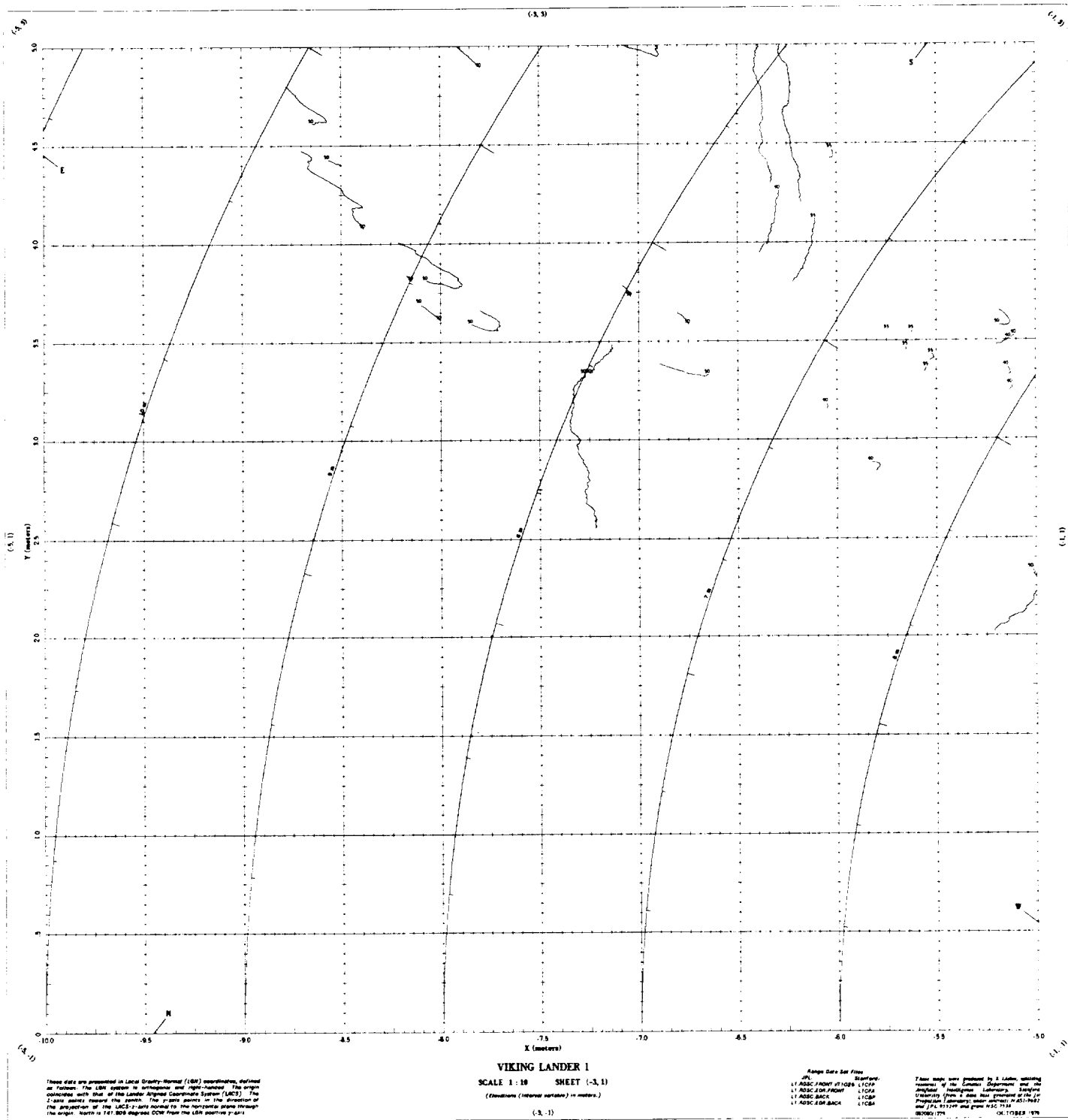
VIKING LANDER 1
SCALE 1:10 SHEET (-1,3)
(Elevations (vertical variables) in meters.)

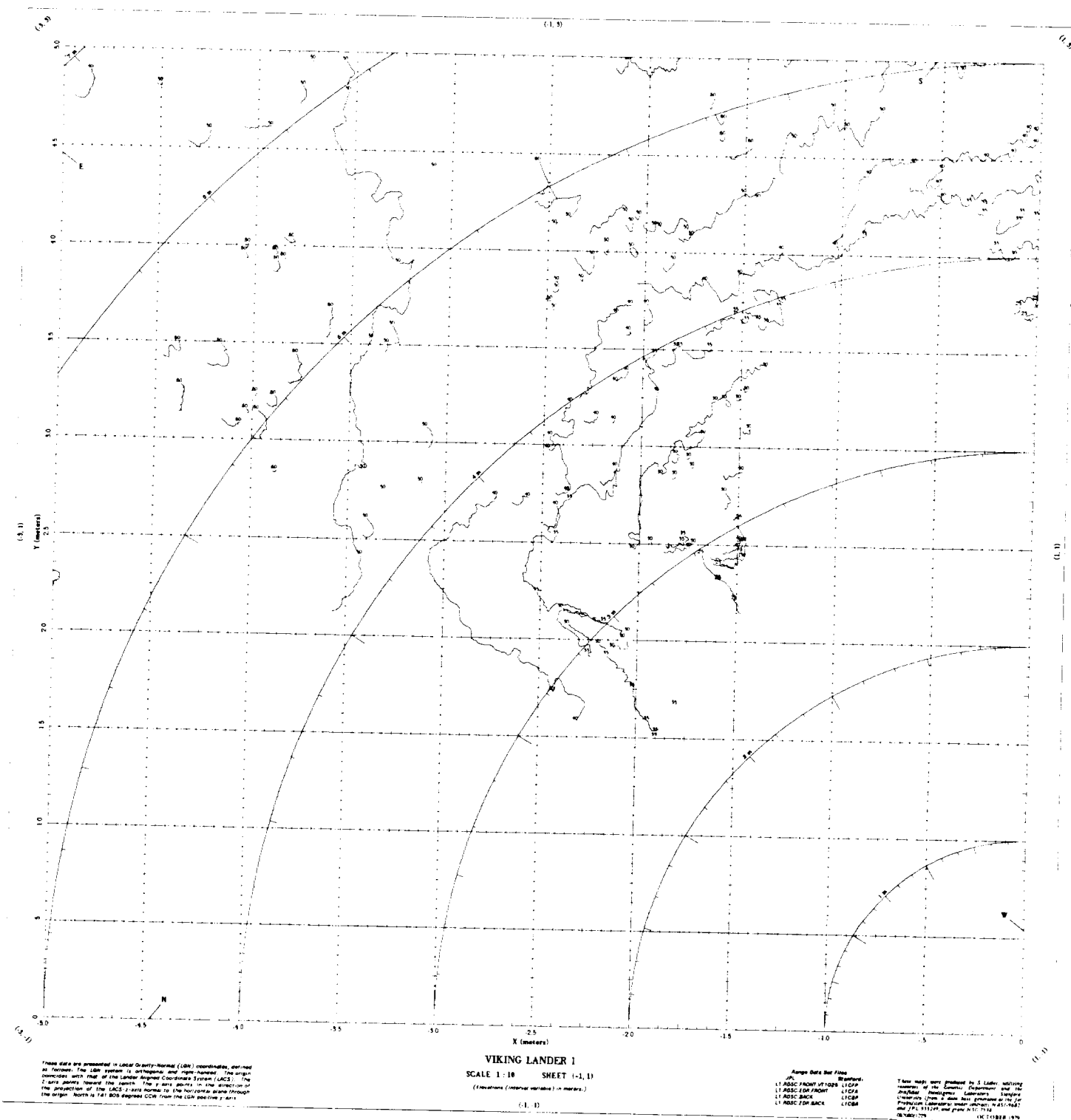
(-1, 3)

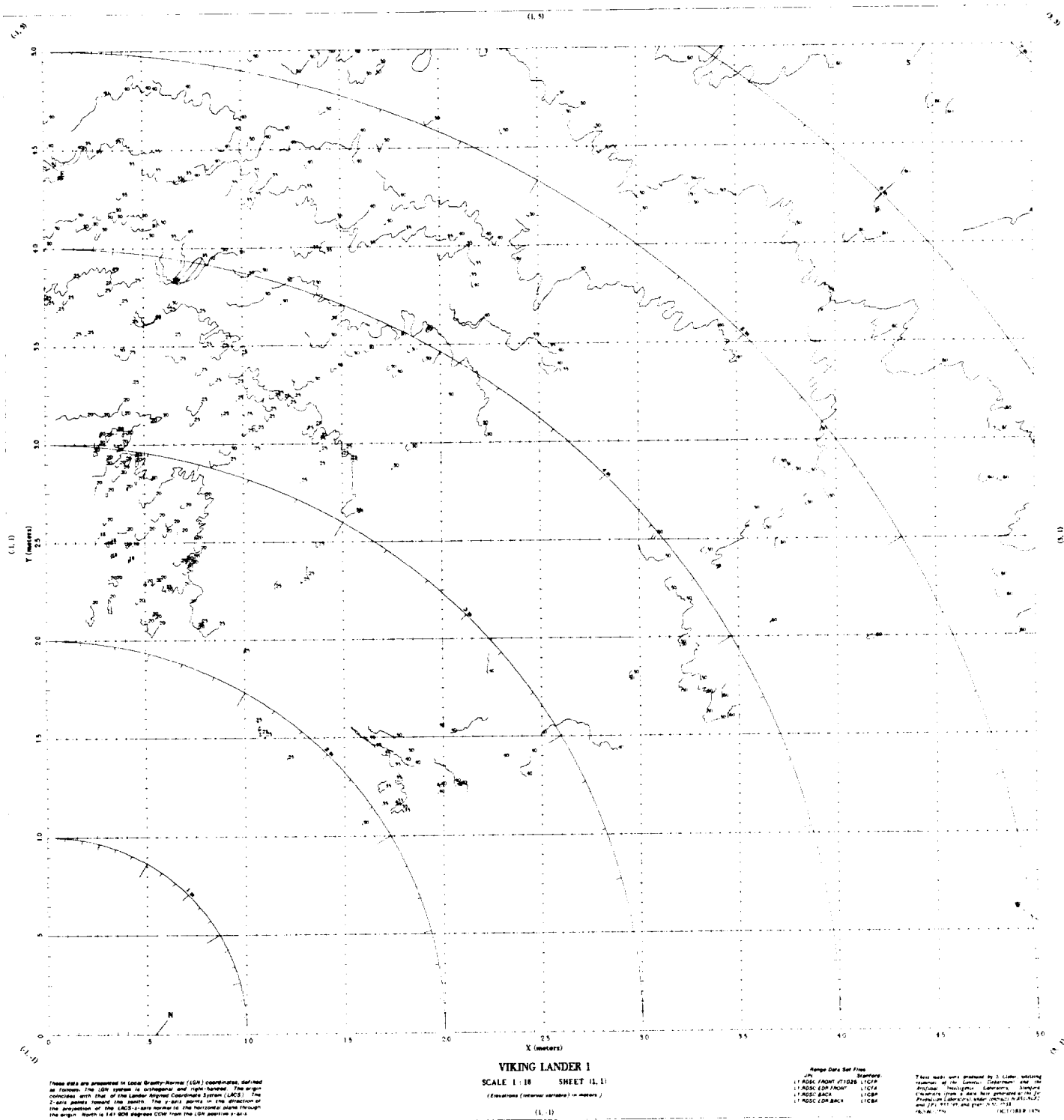
Range Data Set Files
JPL Standard:
1) ROSE FRONT PT1028 (TCFA)
1) ROSE LON FRONT (TCFA)
1) ROSE BACK (TCFA)
1) ROSE LON BACK (TCFA)

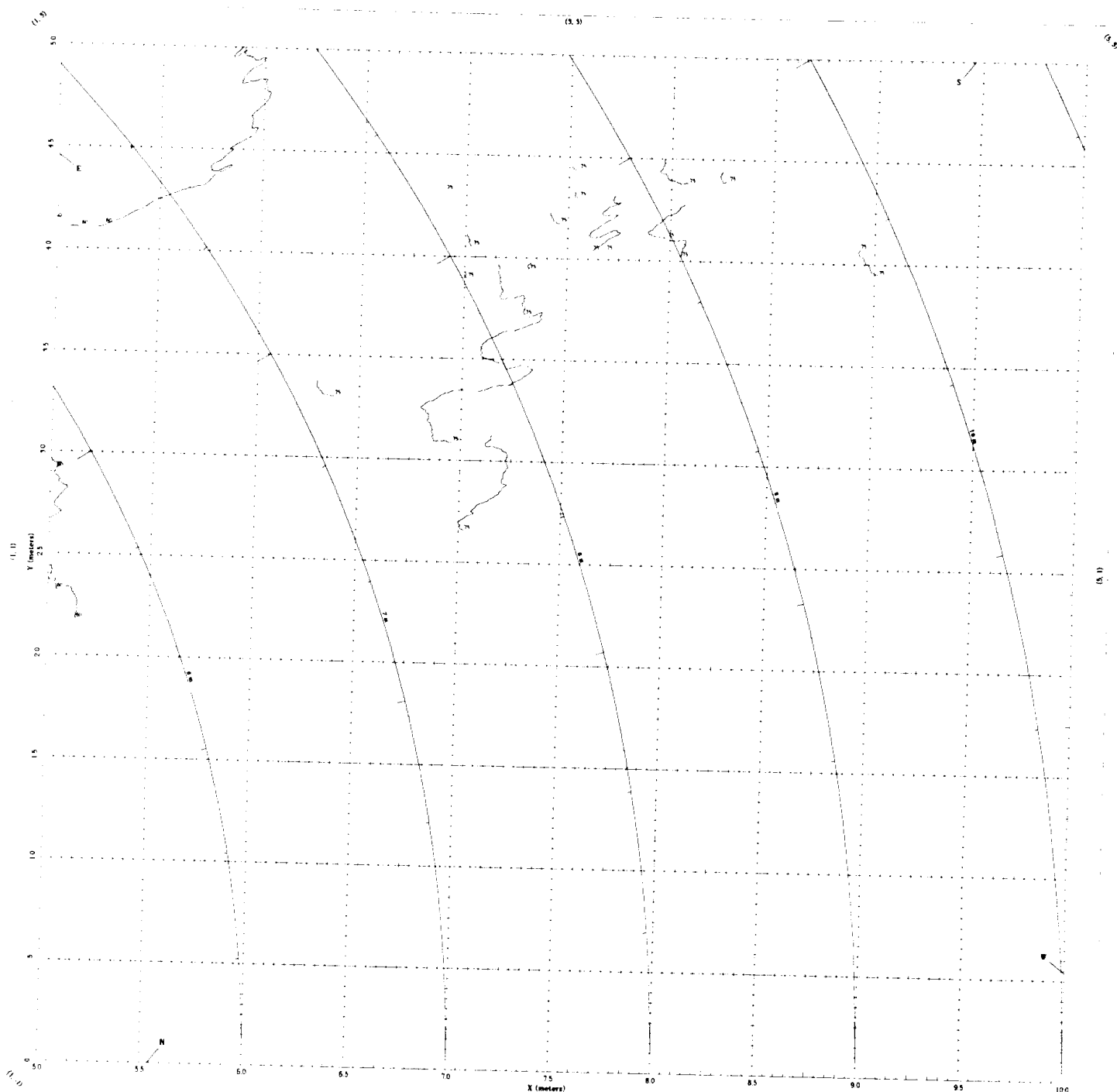
Files made were produced by S. Lander, including
images of the Camera, Department of the
Astronomy, University of California, Berkeley
University of California, Berkeley, Department of
Physics, University of California, Berkeley, and
JPL 931199 and from NAC 7138
OCTOBER 1979











These data are presented in Local Orthographic Reference (LOR) coordinates, derived at Phoenix. The LOR system is orthogonal and right-handed. The origin coincides with that of the Mars Global Coordinate System (MGS). The X axis points toward the south. The Y axis points in the direction of the projection of the MGS Z axis onto the Mars surface plane through the origin. North is 141.905 degrees CGR from the LOR positive Y axis.

VIKING LANDER 1
SCALE 1:10 SHEET (1,1)
(Elevations (vertical scale) in meters)

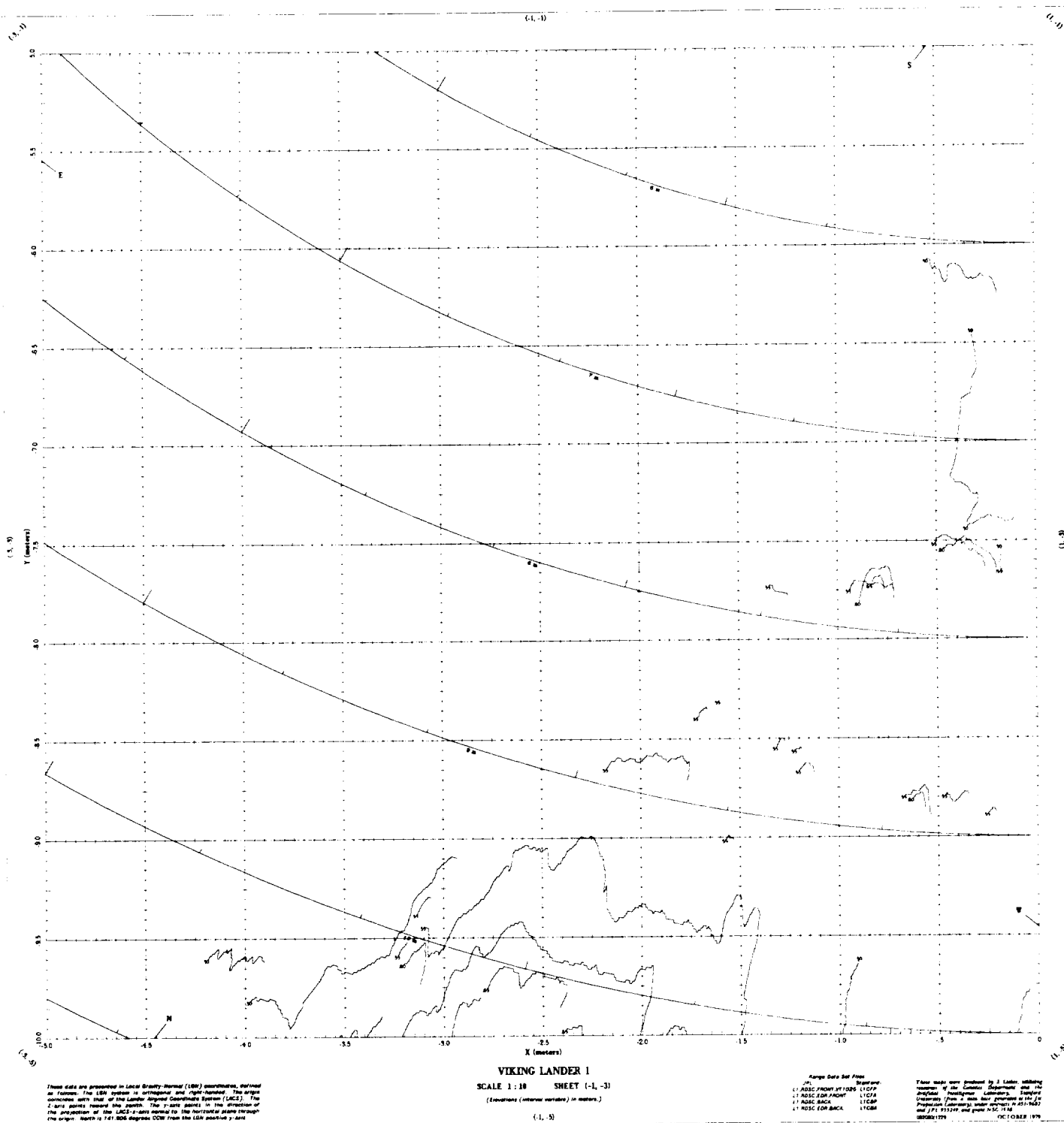
(5, 1)

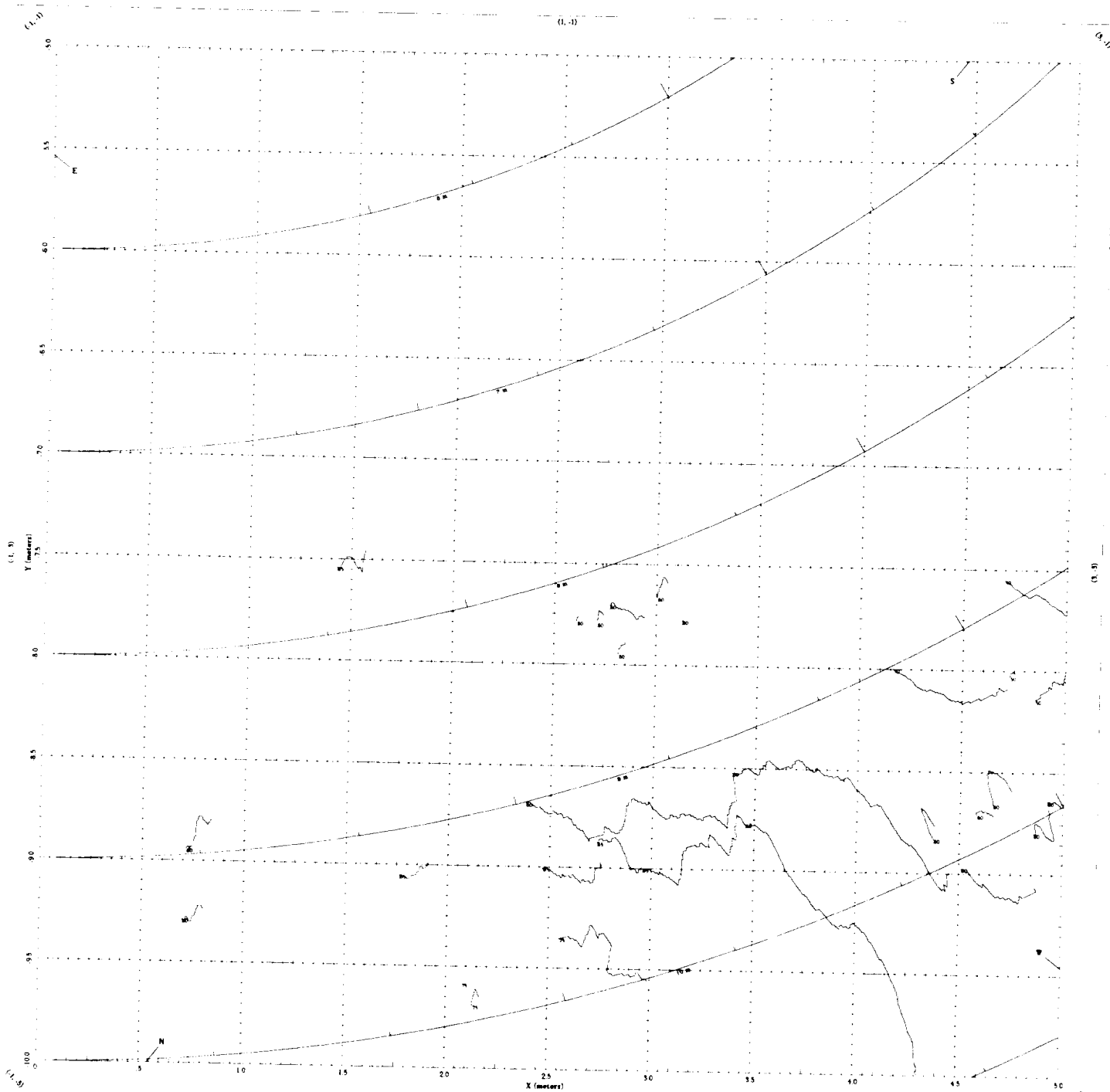
Range Data Set Files
L1 RSC FRONT ST1088 L1C1A
L1 RSC EDN FRONT L1C1A
L1 RSC BACK L1C1A
L1 RSC EDN BACK L1C1A

These data were produced by S. Linder, utilizing
software at the University of Arizona, Tucson, AZ.
The data were processed by the
University of Arizona, Tucson, AZ.
The data were processed by the
University of Arizona, Tucson, AZ.
The data were processed by the
University of Arizona, Tucson, AZ.

000000-1000

OCTOBER 1978





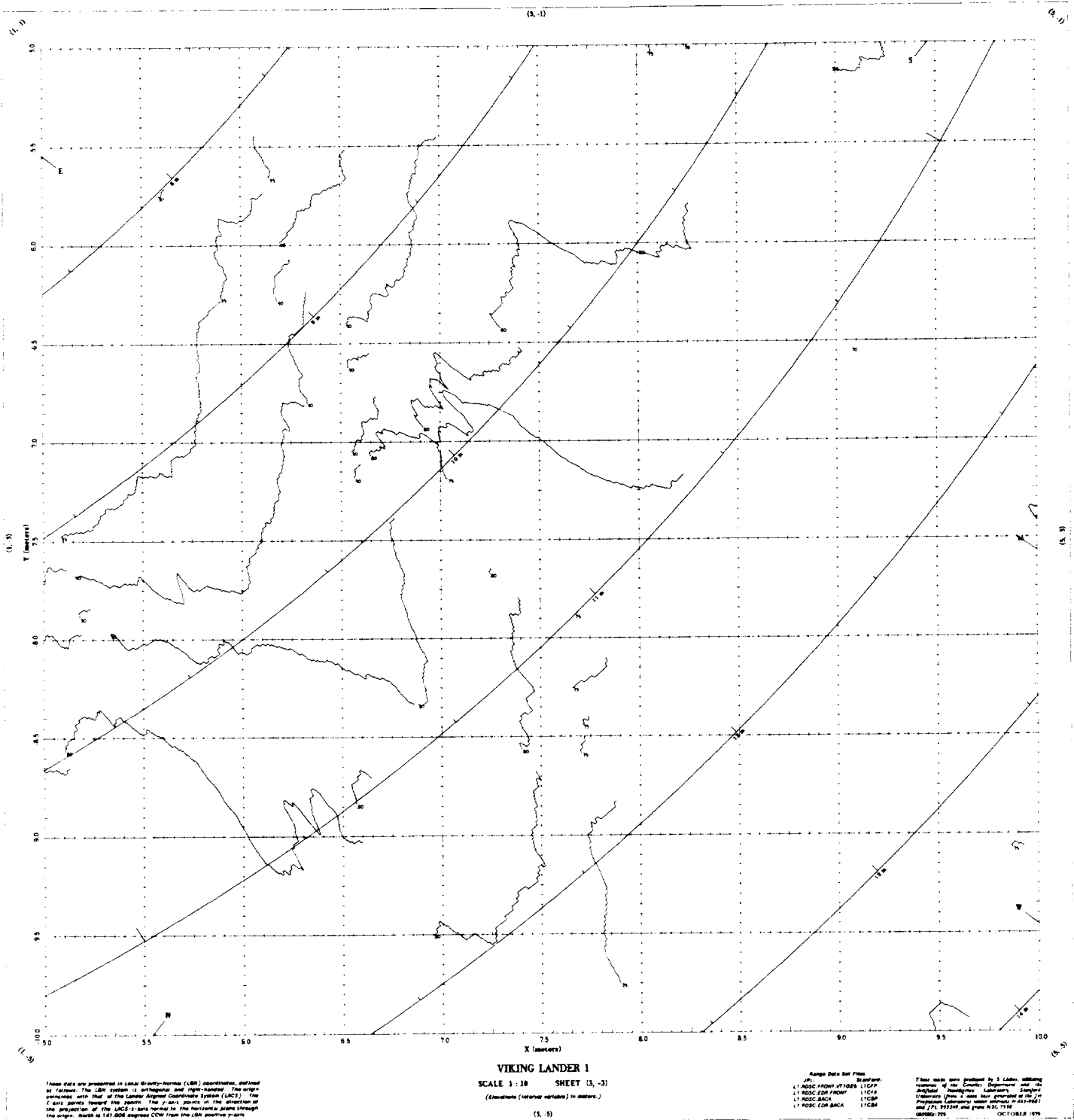
These data are presented in Local Gravity Normal (LGN) coordinates, defined as follows. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lander Inertial Coordinate System (LICS). The x-axis points toward the south. The y-axis points in the direction of the projection of the LICS z-axis north to the horizontal plane through the origin. North is 147.806 degrees CCW from the LGN positive y-axis.

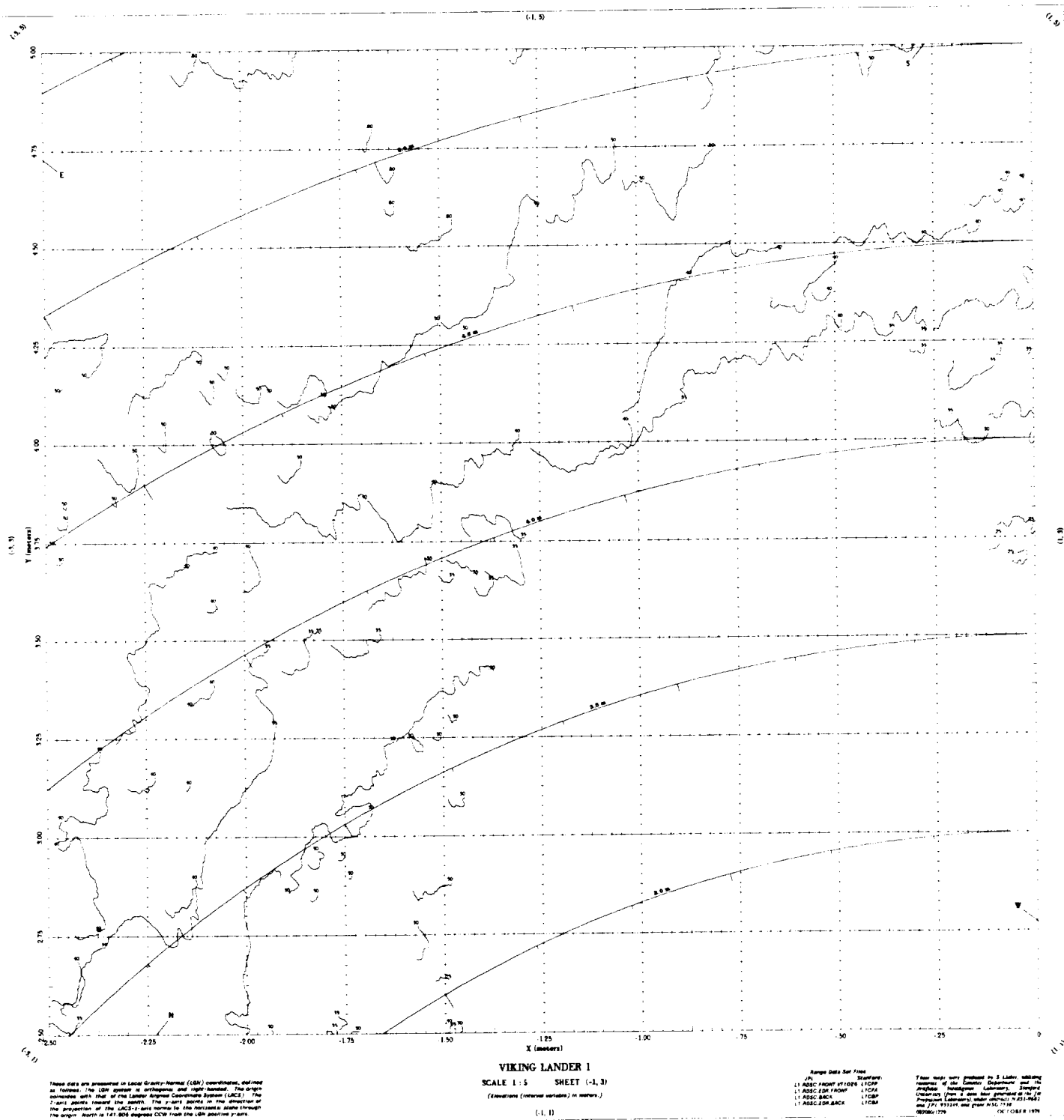
VIKING LANDER 1
SCALE 1:10 SHEET (1, -3)
(Elevations (vertical variable) in meters.)

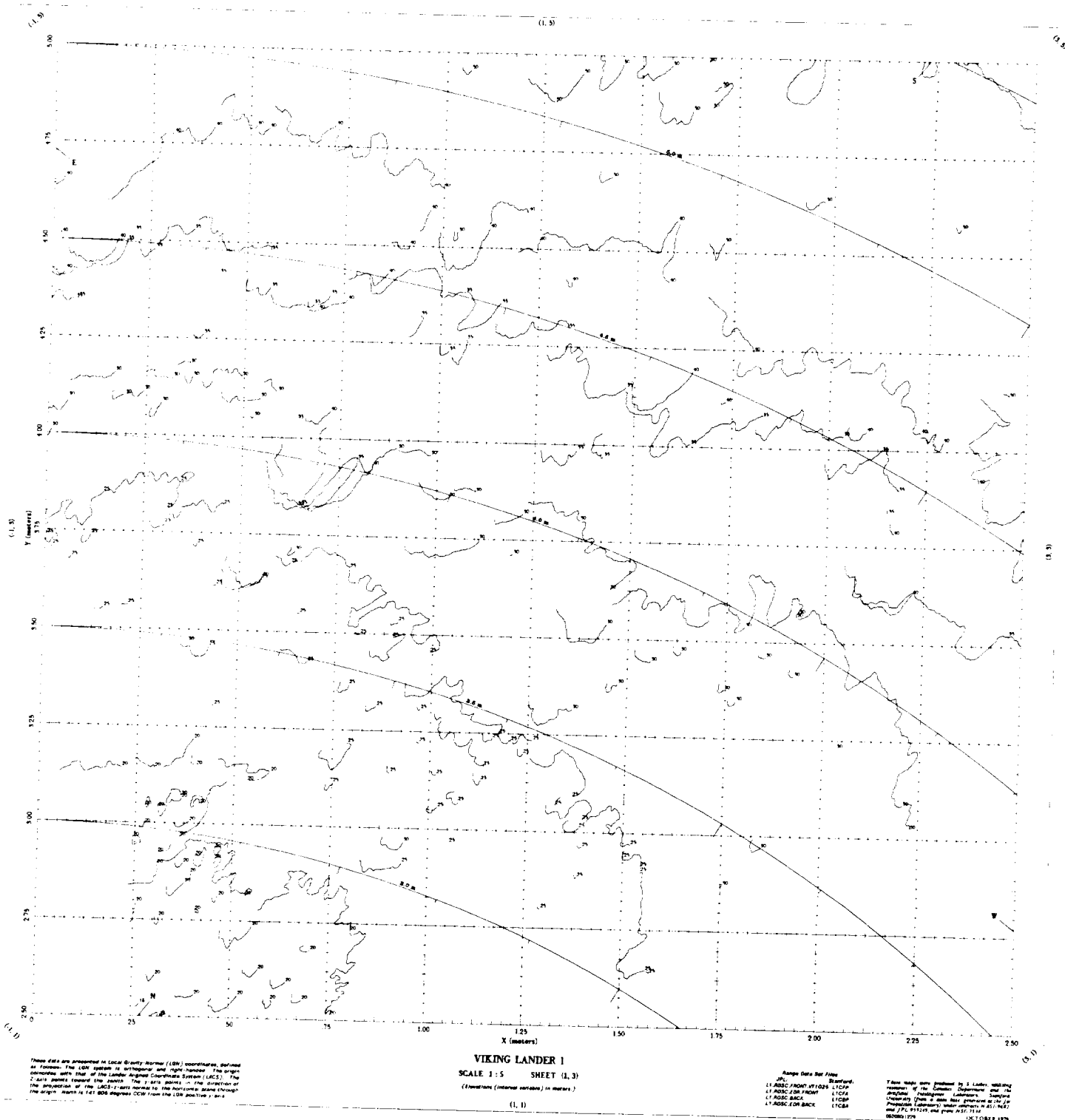
(1, 5)

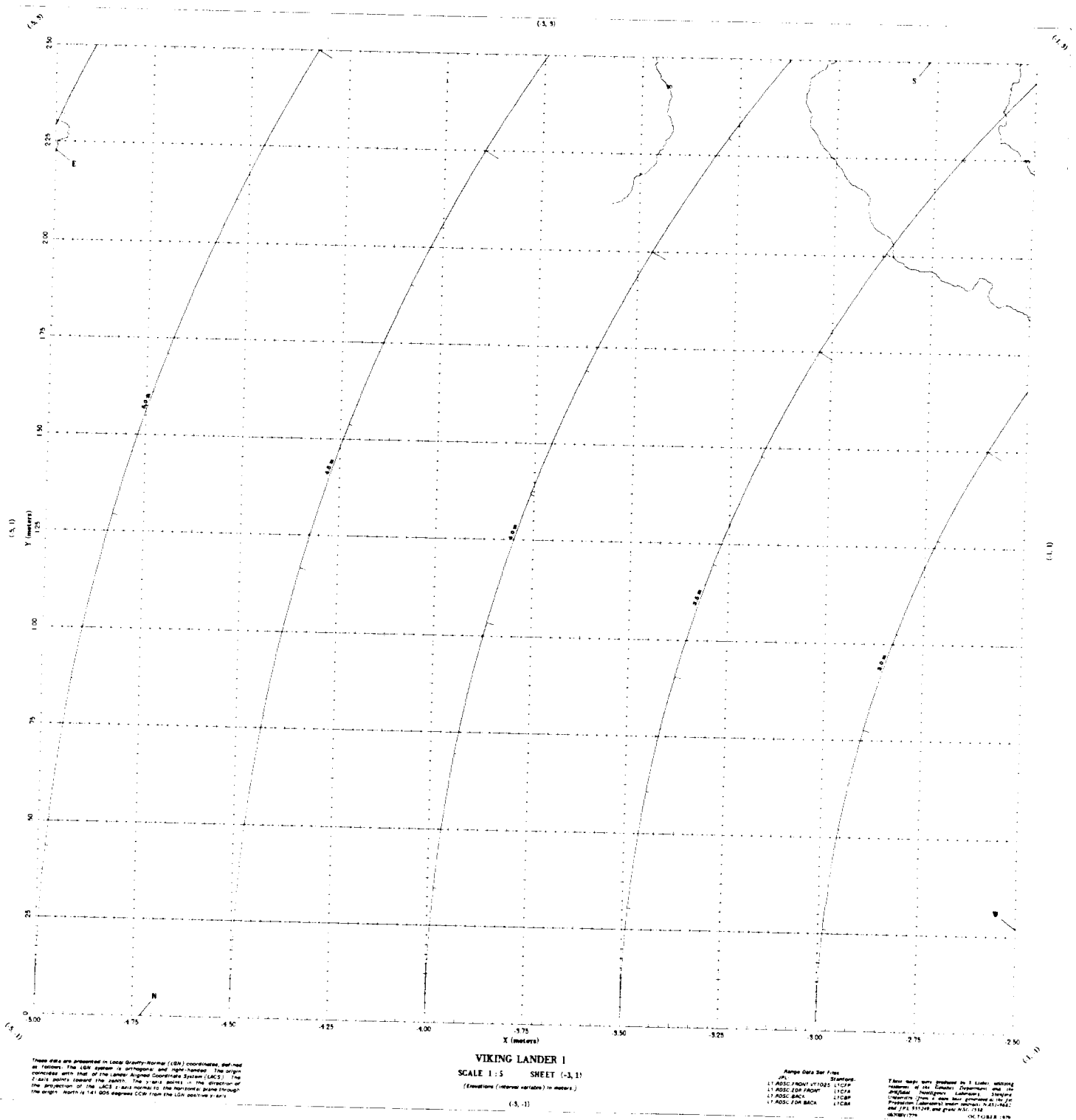
Range Data Set Flag
JPL Standard
1. NSC FROM V1028 L1CFF
1. NSC ON ROM L1CFF
1. NSC BACK L1CFF
1. NSC FOR BACK L1CFF

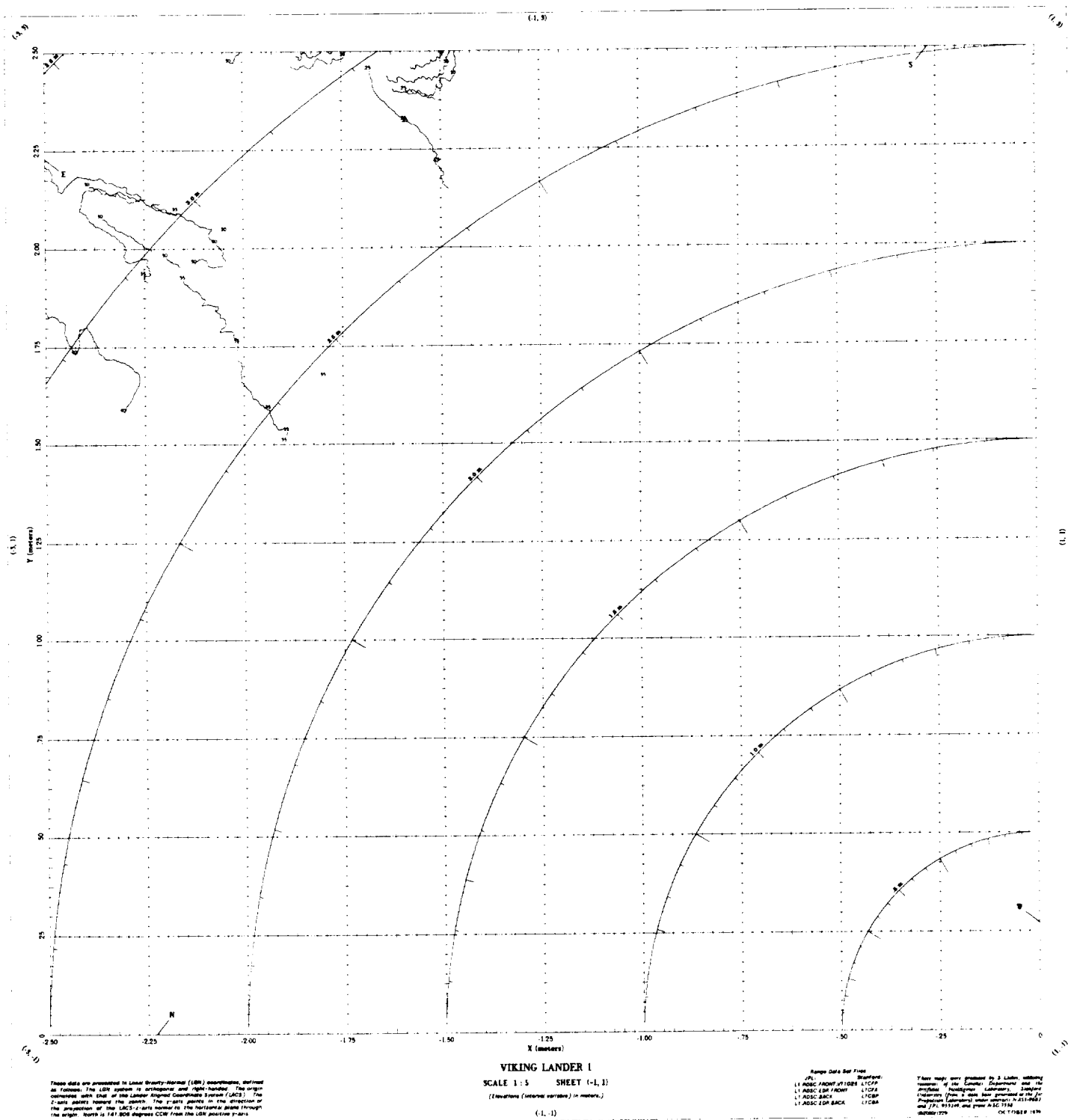
These maps were produced by S. L. Lander, utilizing resources of the Lander Department and the JPL Navigation Laboratory. The data were generated by the JPL Navigation Laboratory under contract N-001-60-2 and JPL 111249 and given NAC 1112.

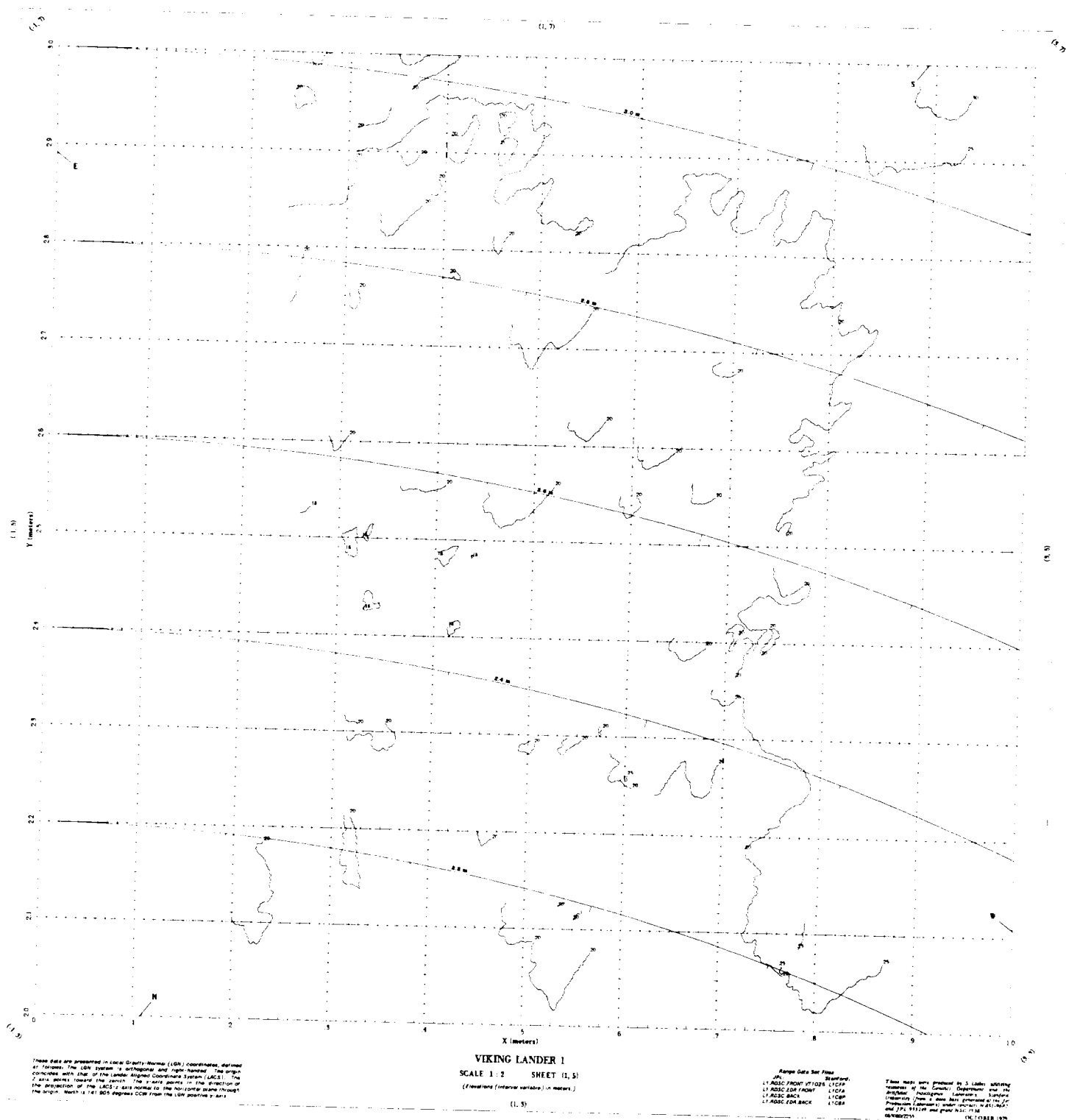


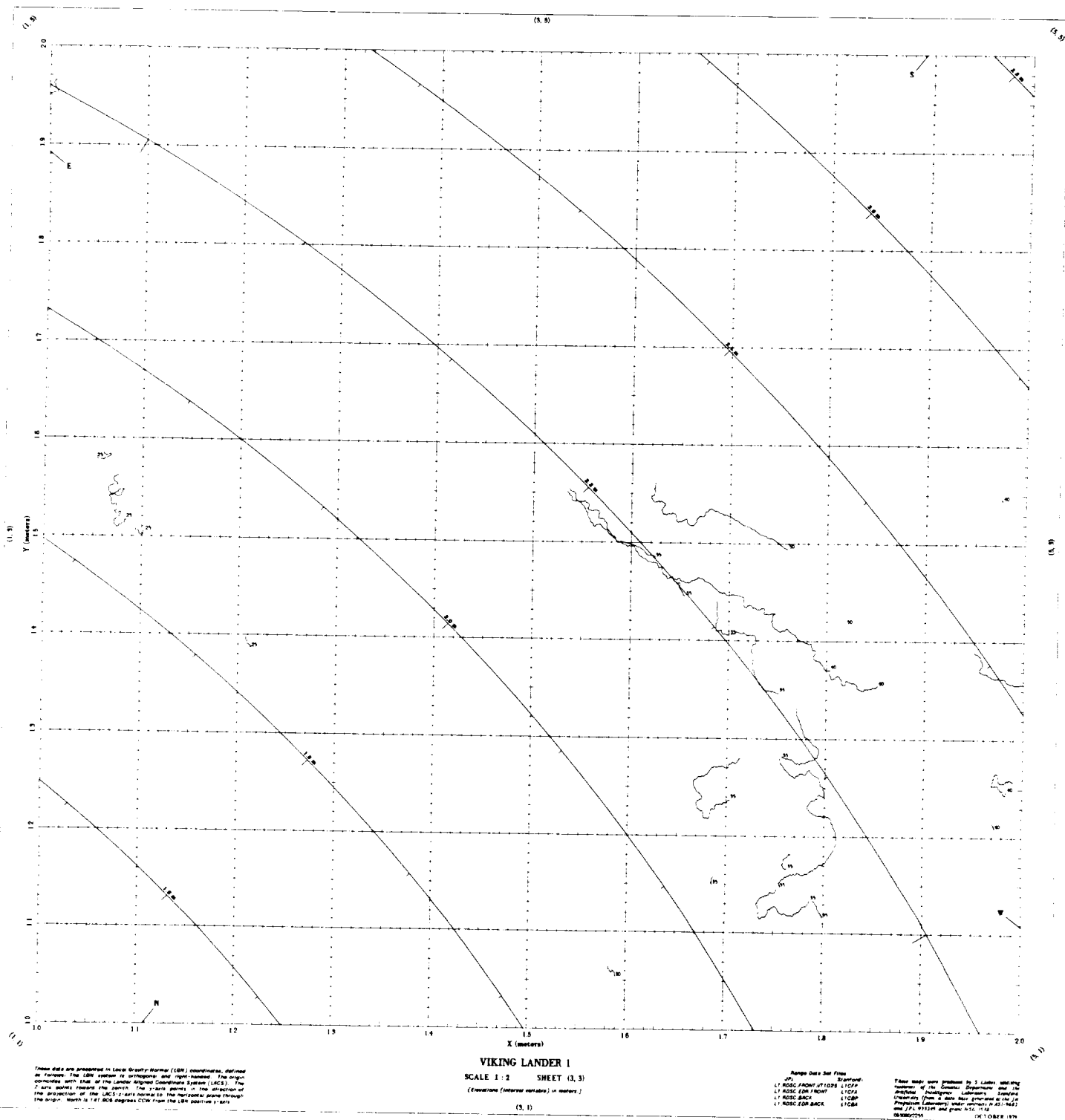


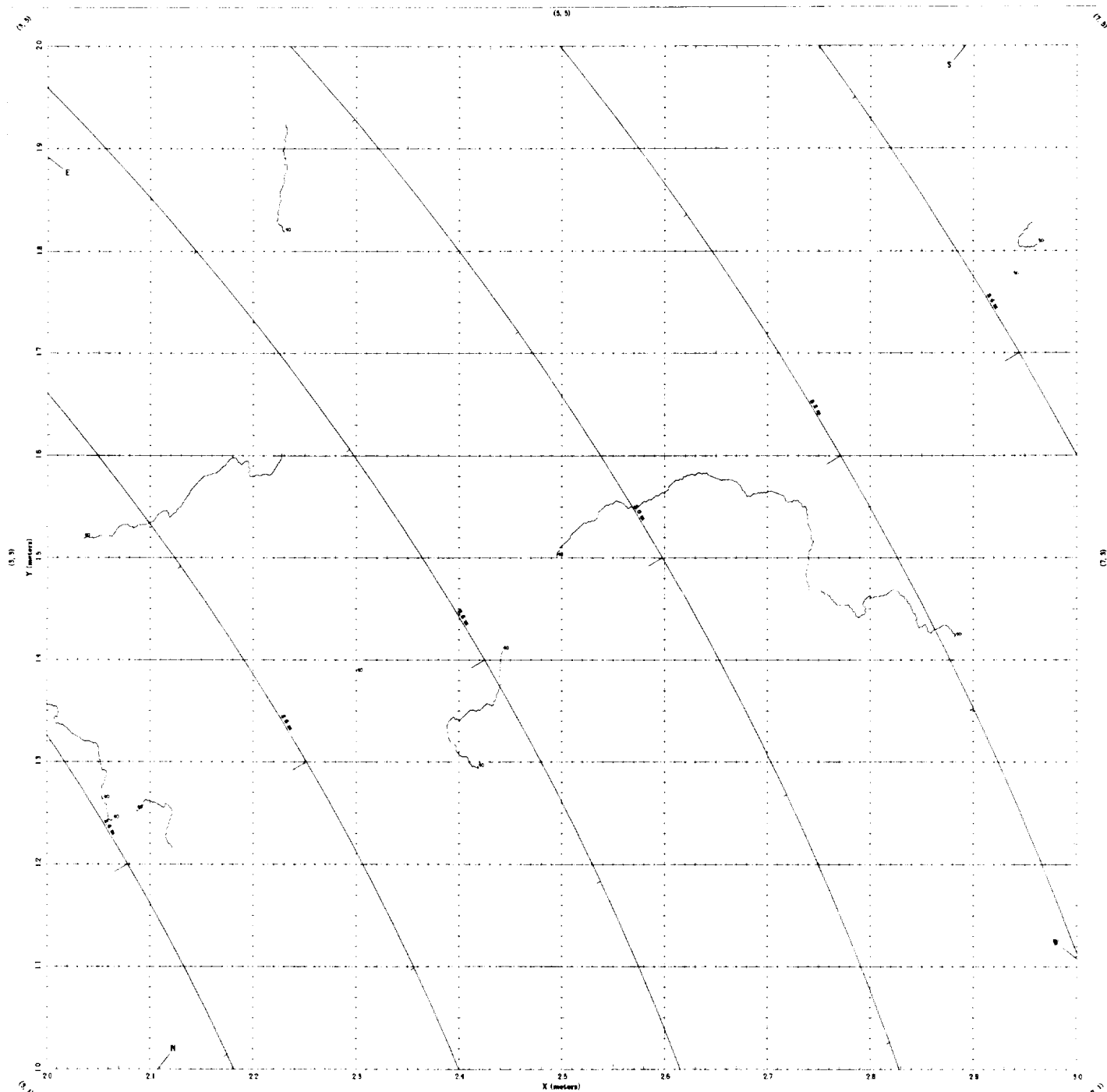












VIKING LANDER 1

SCALE 1:1 SHEET (5, 3)

(Elevations (meters) are in meters.)

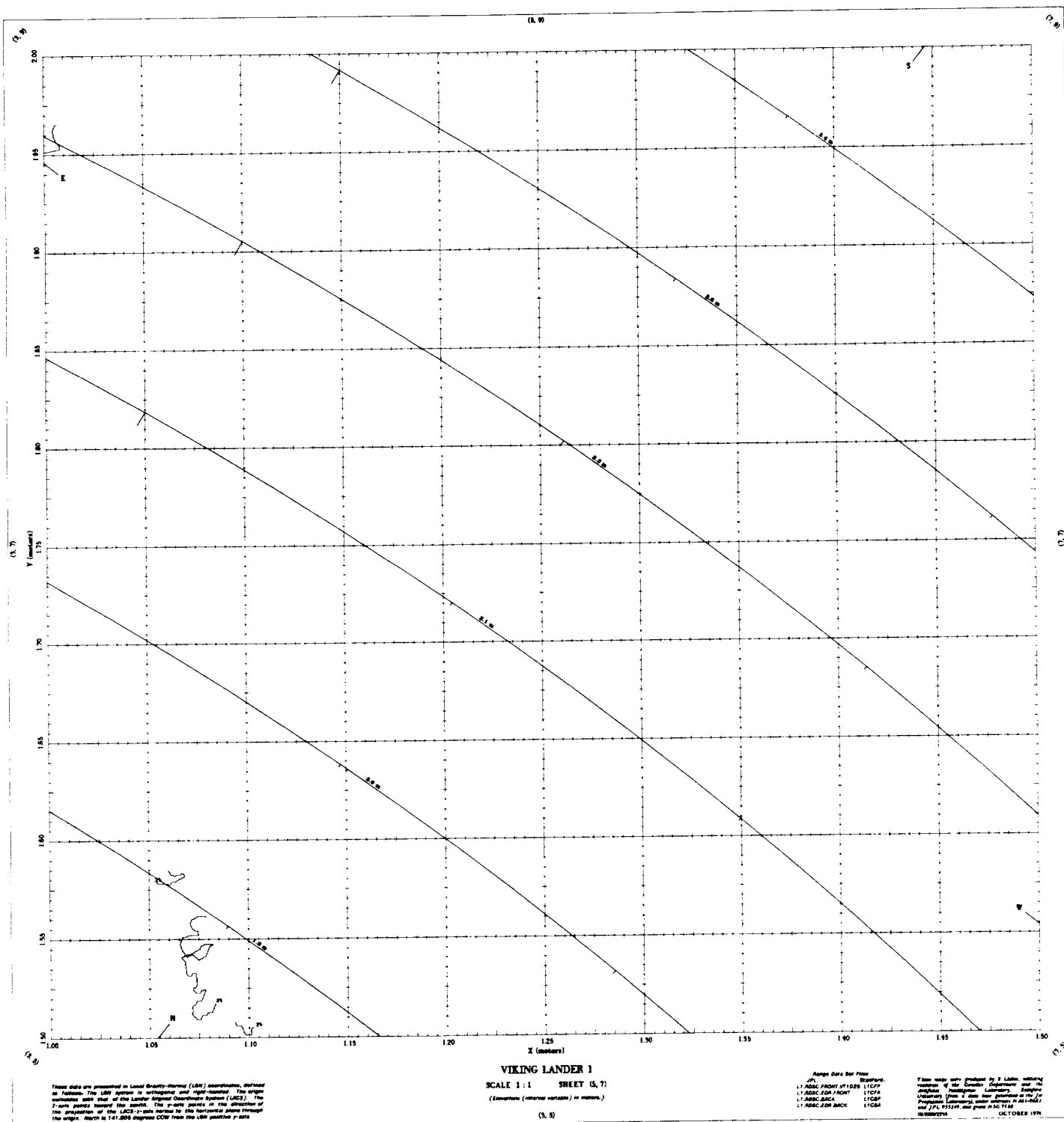
(5, 1)

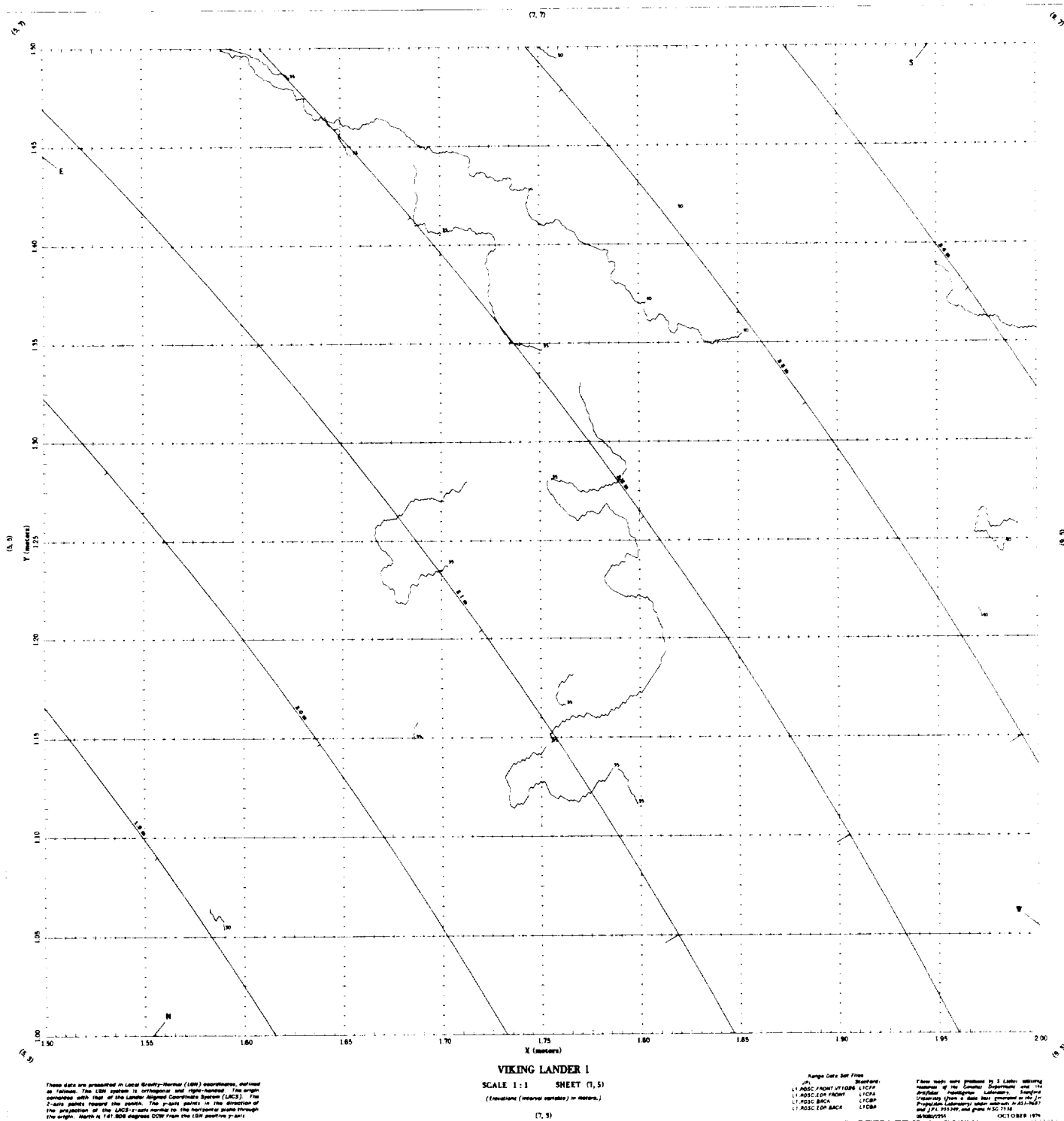
Ramp Data Set Files

JPL:
 1. ASOC FRONT V1000 L1000
 1. ASOC EDM FRONT L1000
 1. ASOC BACK L1000
 1. ASOC EDM BACK L1000

These data were provided by J. Lander, visiting scientist, of the Center for Space and Terrestrial Sciences, University of California, San Diego. The data were provided to the JPL Viking Lander Project by J. Lander, visiting scientist, of the Center for Space and Terrestrial Sciences, University of California, San Diego. The data were provided to the JPL Viking Lander Project by J. Lander, visiting scientist, of the Center for Space and Terrestrial Sciences, University of California, San Diego.

UNCLASSIFIED
 14 DEC 1979





5. VERTICAL PROFILES

5.1 Vertical Profile Mosaic Overlay Stereo Pairs

This section contains mosaic stereo pairs of images into which have been overlayed the systematic vertical profiles. The general remarks of section 4.1 apply.



Fig. 19. Vertical Profile Overlay Mosaic: Camera 1 (left), front left quadrant —
from IPL PIC ID 79/07/15/102049.



Fig. 20. Vertical Profile Overlay Mosaic: Camera 2 (right), front left quadrant —
from IPL PIC ID 79/06/06/021223.

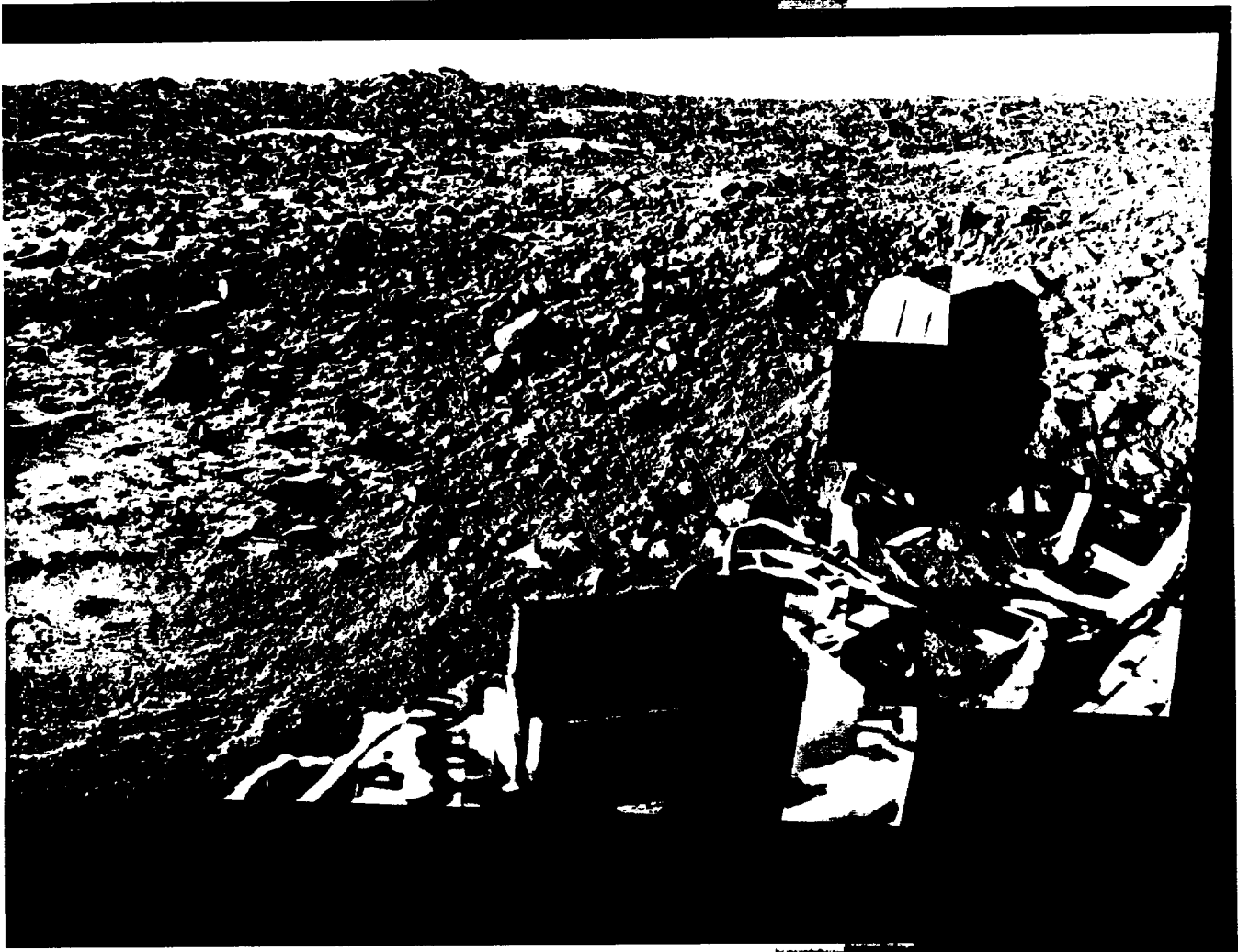


Fig. 21. Vertical Profile Overlay Mosaic: Camera 1 (left), front right quadrant —
from IPL PIC ID 79/07/15/102049.



Fig. 22. Vertical Profile Overlay Mosaic: Camera 2 (right), front right quadrant —
from IPL PIC ID 79/06/06/021223.

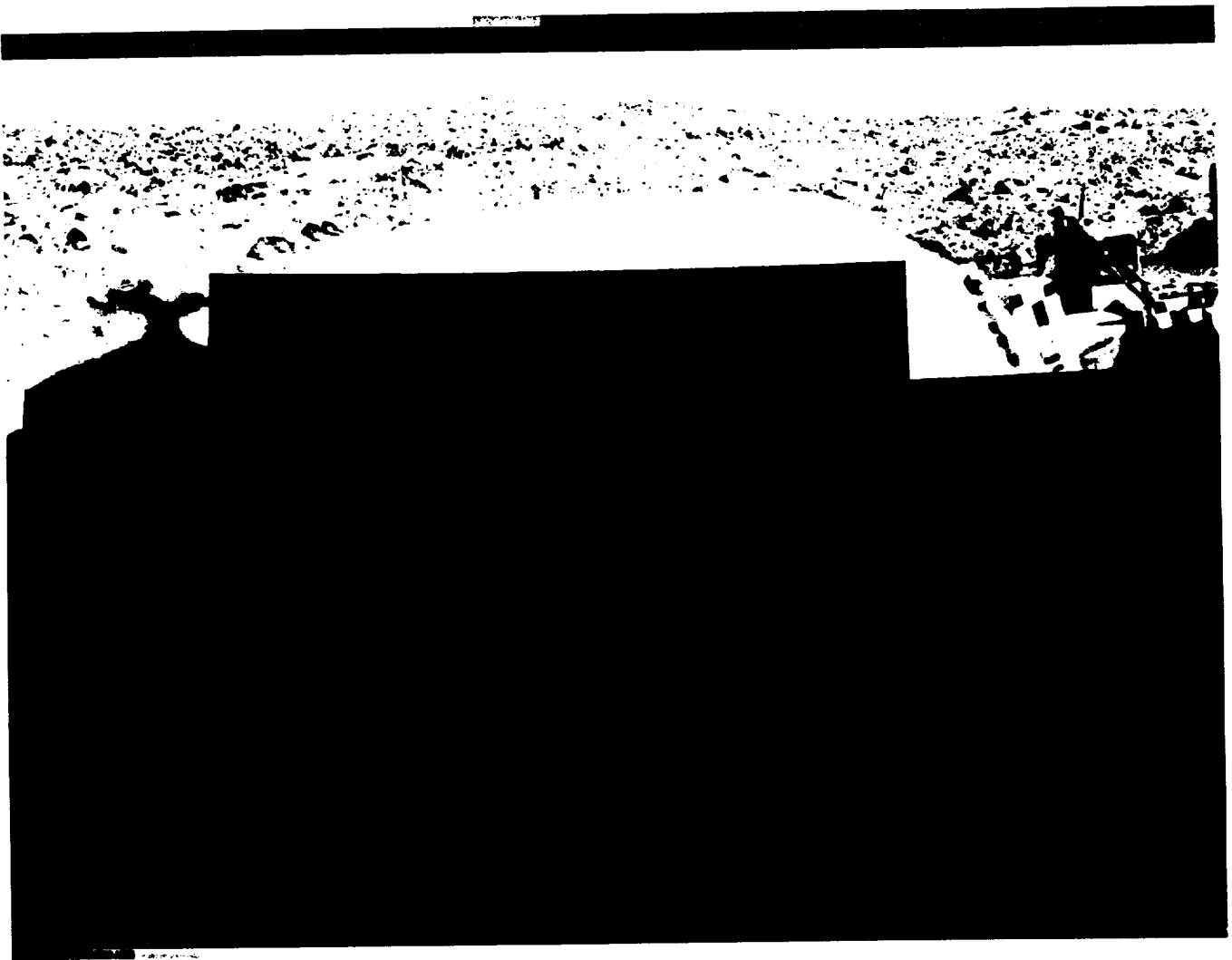


Fig. 23. Vertical Profile Overlay Mosaic: Camera 2 (left), back left quadrant -
from IPL PIC ID 79/07/15/113640.

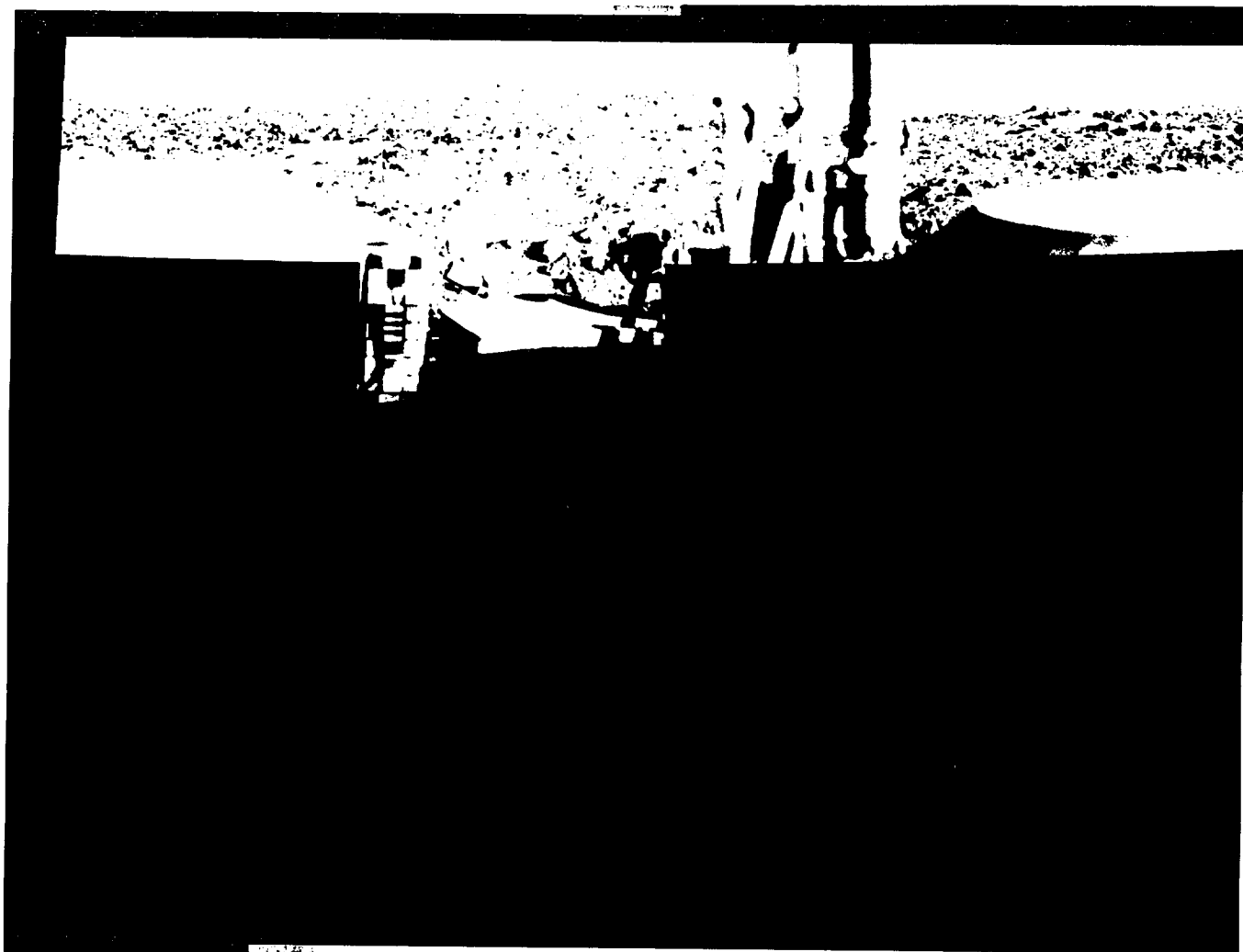


Fig. 24. Vertical Profile Overlay Mosaic: Camera 1 (right), back left quadrant --
from IPL PIC ID 79/03/20/044430.

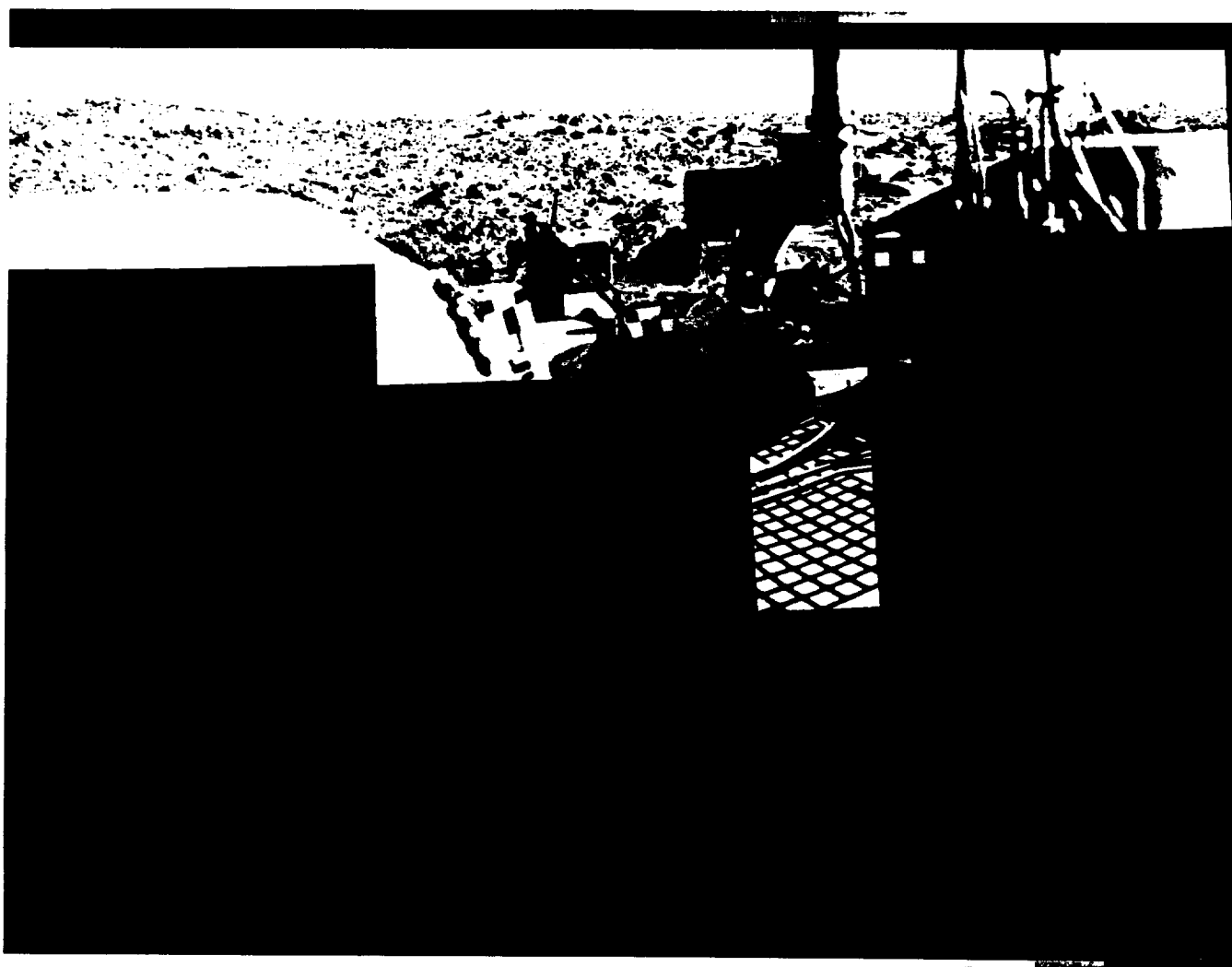


Fig. 25. Vertical Profile Overlay Mosaic: Camera 2 (left), back right quadrant —
from IPL PIC ID 79/07/15/113640.

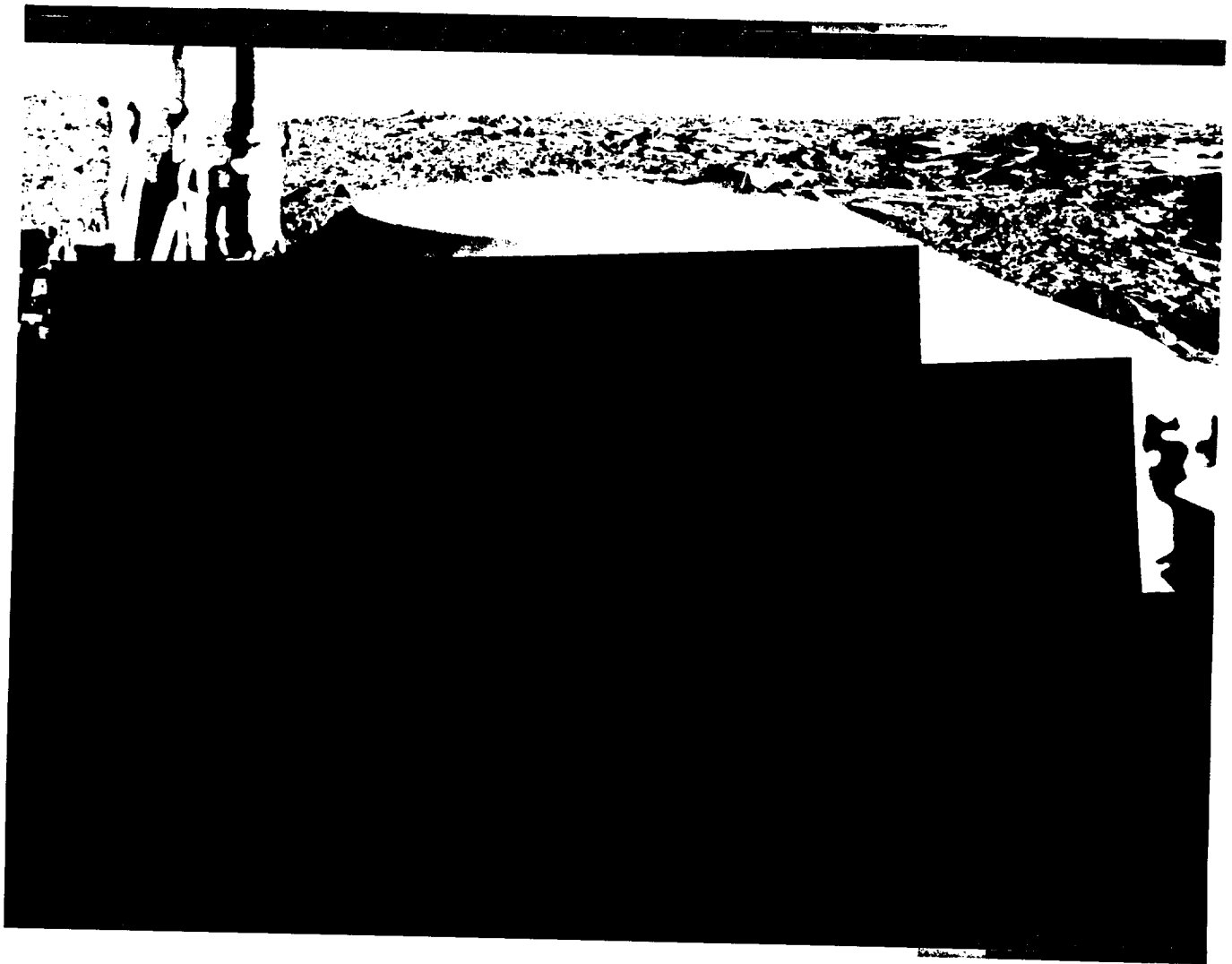


Fig. 26. Vertical Profile Overlay Mosaic: Camera 1 (right), back right quadrant —
from IPL PIC ID 79/03/20/014430.

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5.2 Camera Perspective Annotated Vertical Profiles — Camera 1

This section contains camera-1 perspective representations of the elevation contour map data. The nature and purpose of these representations has been explained in section 3.2.1. On opposing pages are presented mosaic overlays of the elevation con-

tour lines, and corresponding camera perspective representation of the map data, with the individual lines of the map data annotated with elevation values. The general discussion of section 4.1 applies to the mosaics of this section.

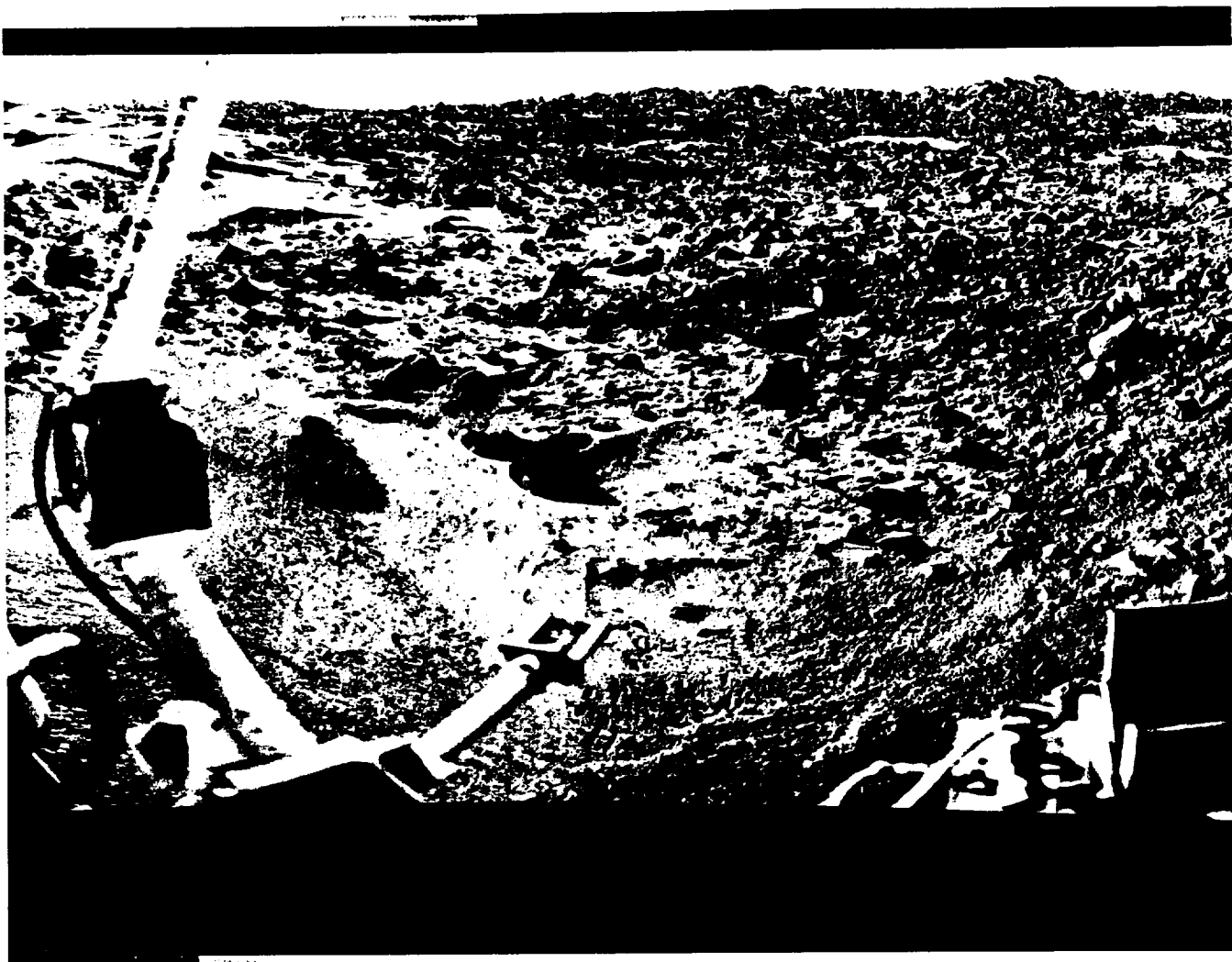


Fig. 27. Vertical Profile Overlay Mosaic: Camera 1 (left), front left quadrant —
from IPL, PIC ID 79/07/15/102049.

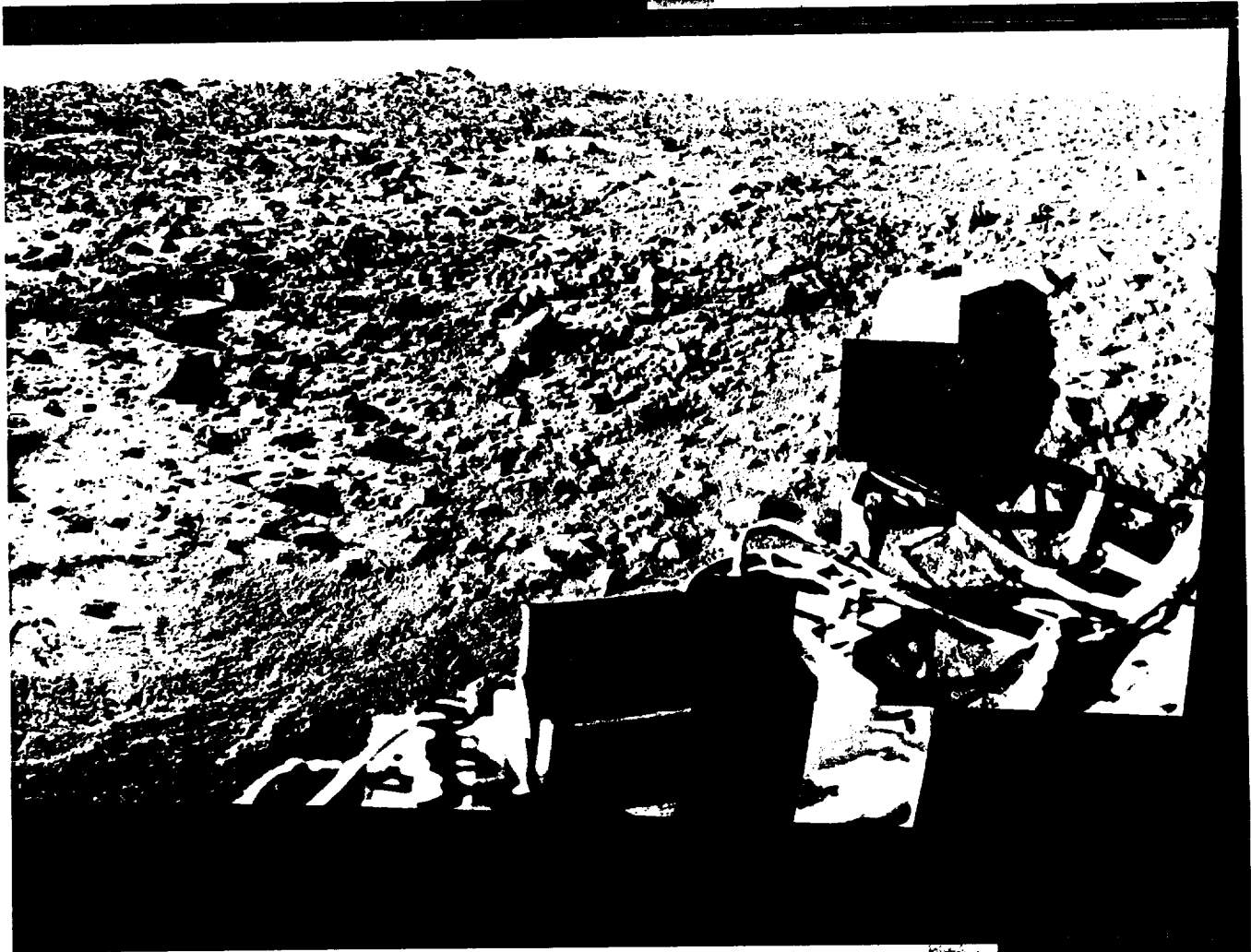


Fig. 29. Vertical Profile Overlay Mosaic: Camera 1 (left), front right quadrant — from IPI, PIC ID 79/07/15/102049.

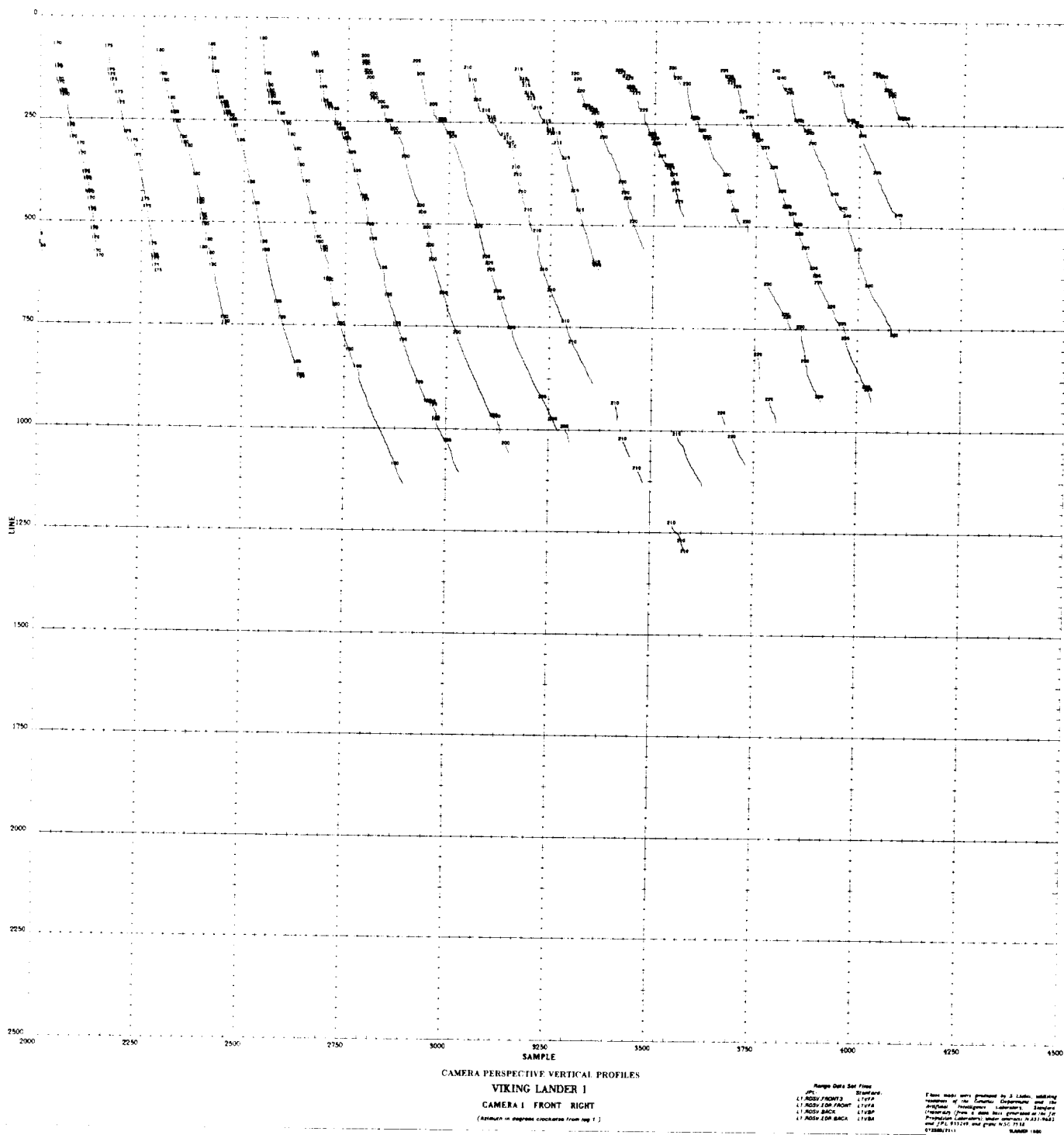


Fig. 30. Camera Perspective Annotated Elevation Vertical Profile Lines:
Camera 1 (left), front right quadrant.

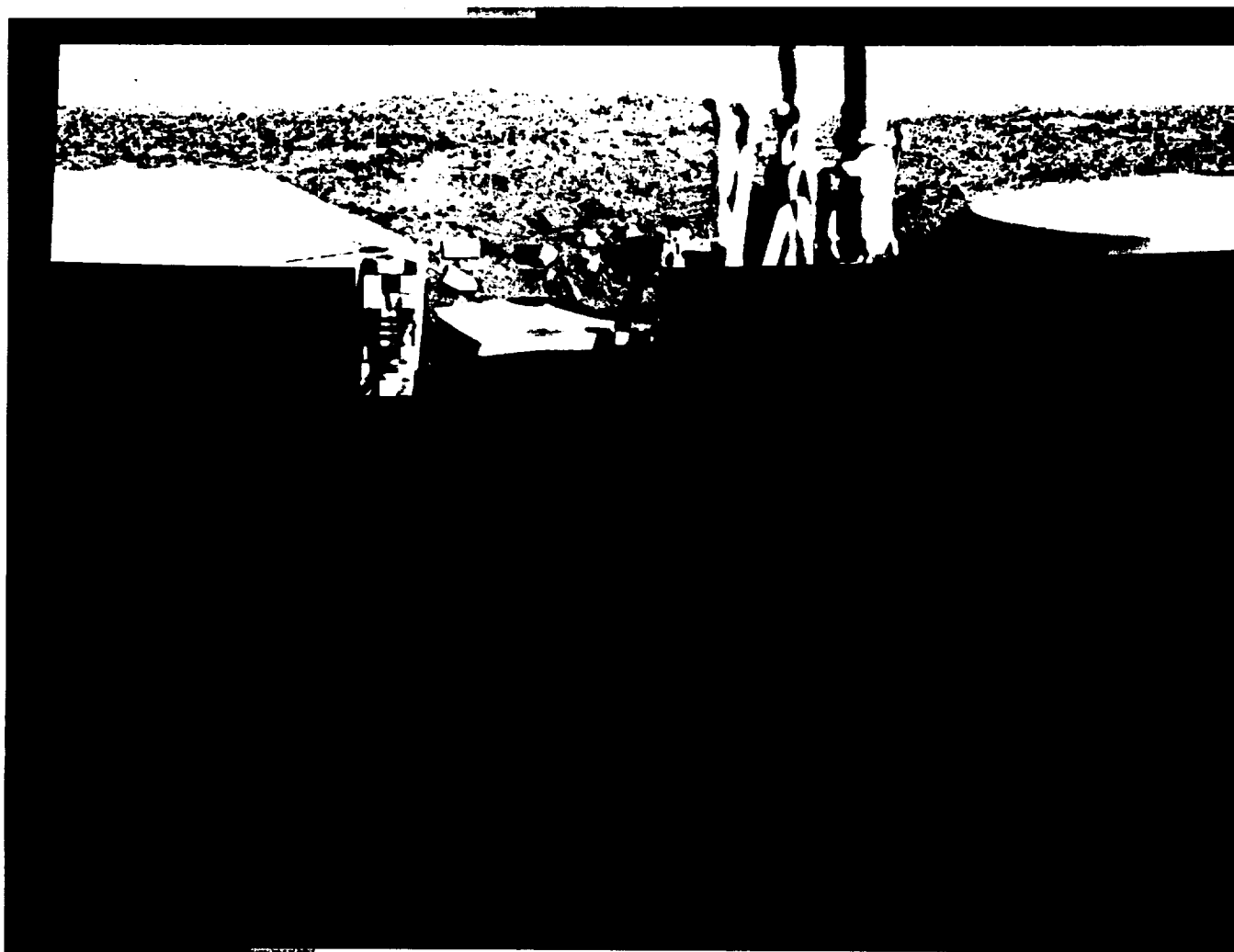


Fig. 31. Vertical Profile Overlay Mosaic: Camera 1 (right), back left quadrant —
from IPI PIC ID 79/03/20/044430.

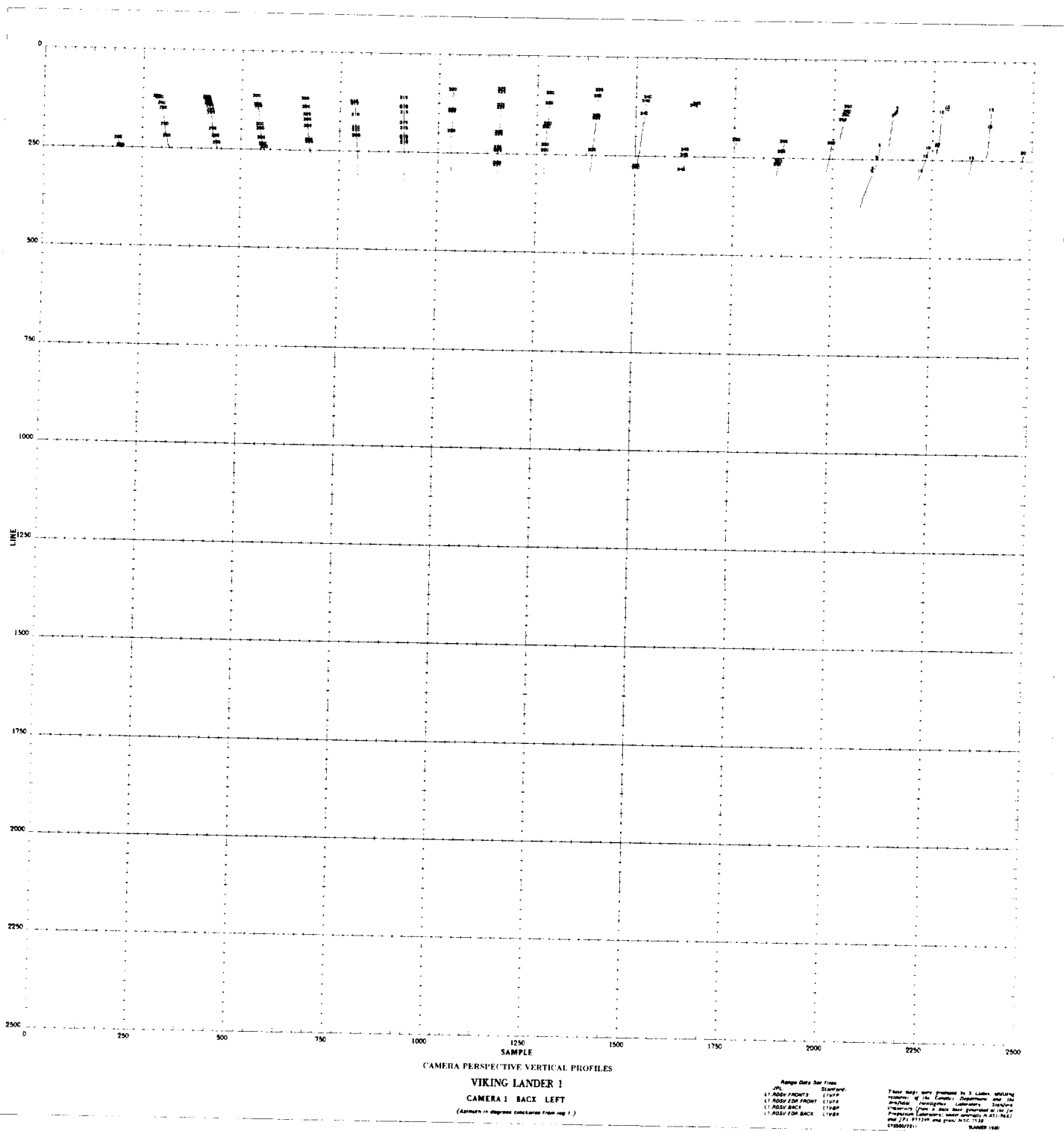


Fig. 32. Camera Perspective Annotated Elevation Vertical Profile Lines:
Camera 1 (left), back left quadrant.

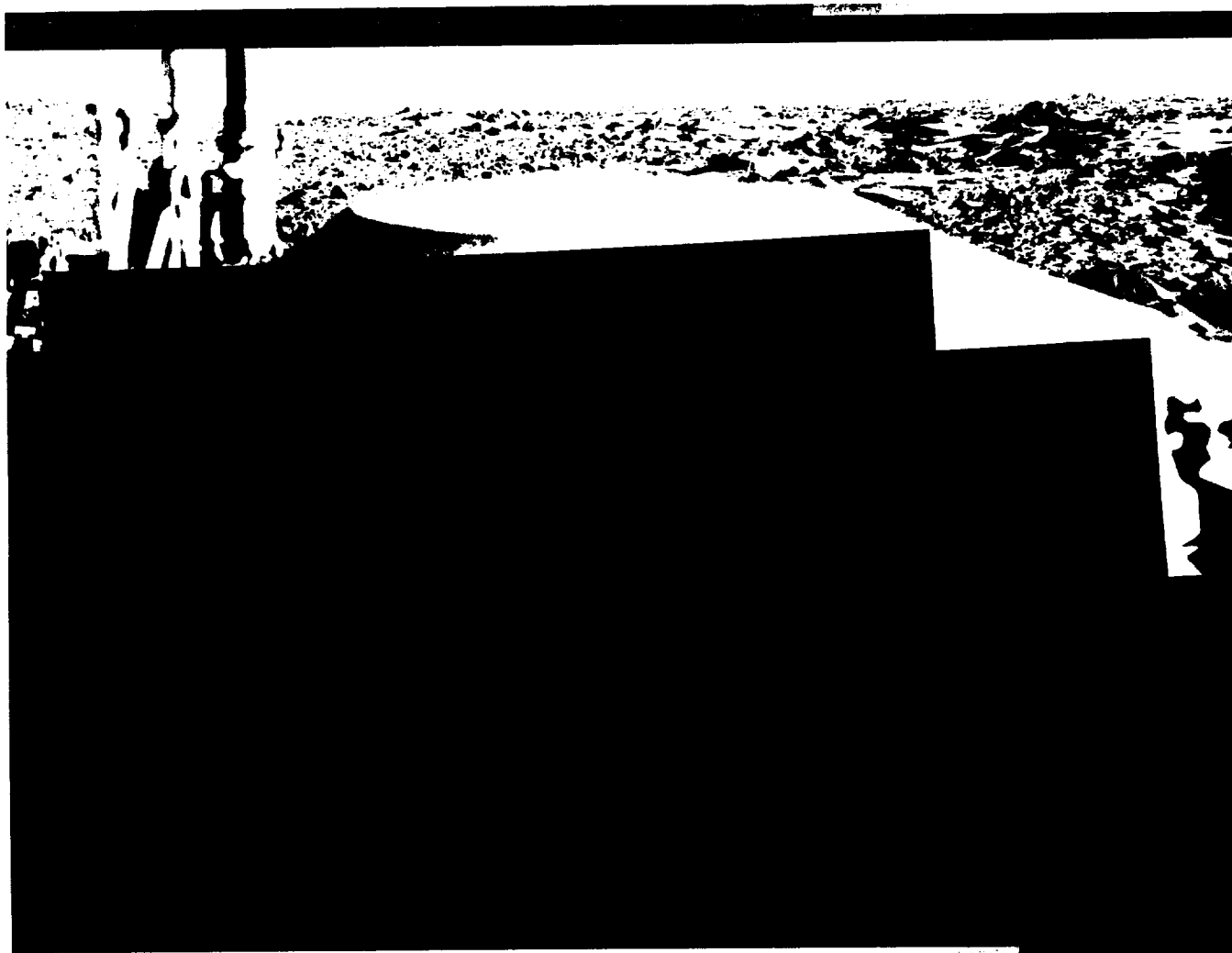
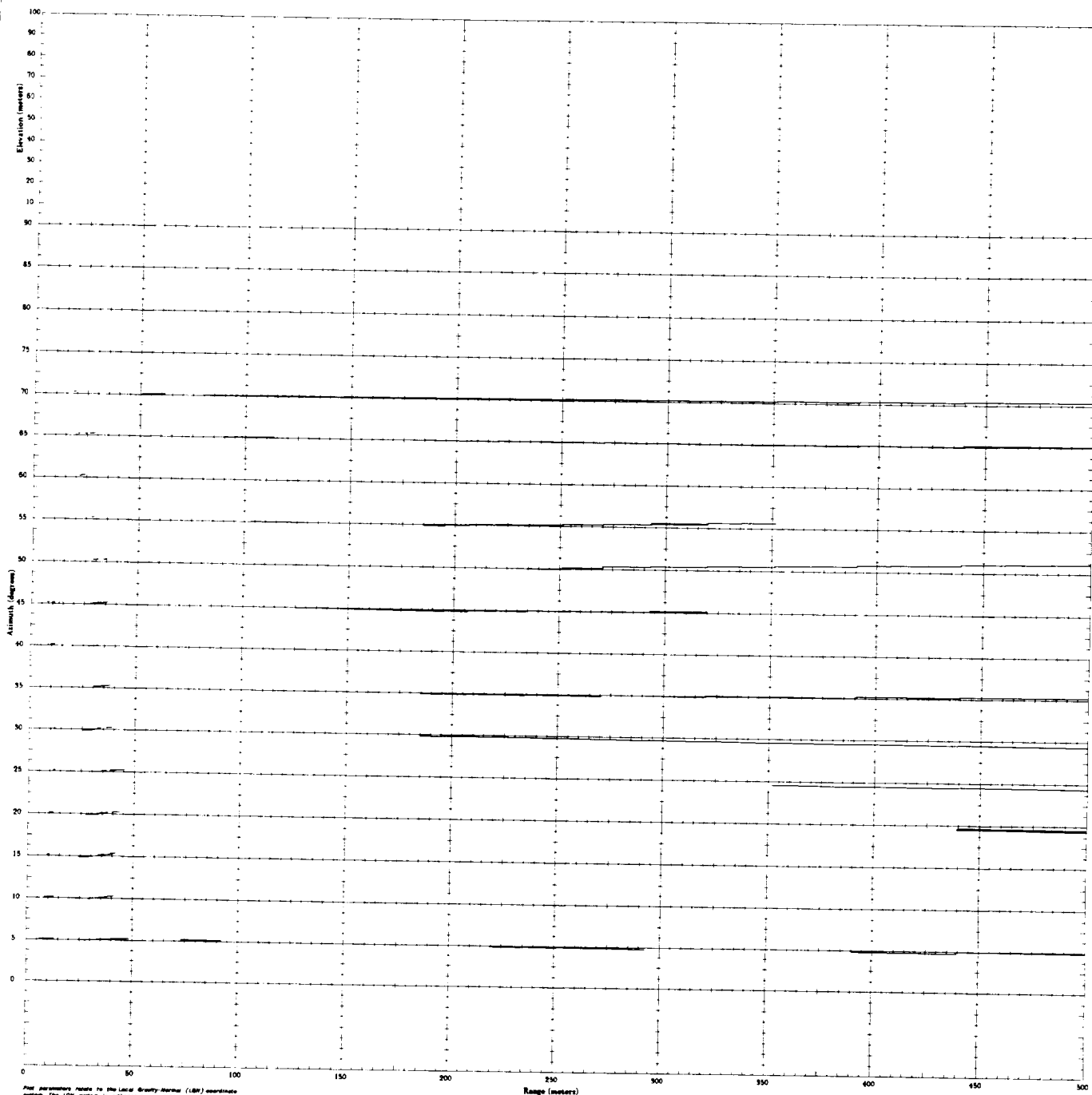


Fig. 33. Vertical Profile Overlay Mosaic: Camera 1 (right), back right quadrant —
from IPL PIC ID 79/03/20/044430.

5.3 Vertical Profile Sheet Collection

This section contains the vertical profile sheet collection. No tabulation of the kind preceding the elevation contour maps has been prepared. The sheets are ordered as follows. First, they are grouped by quadrant, in clockwise order starting from the

back left quadrant (0 - 90°), then within each quadrant by order of decreasing scale number, and within each scale by order of increasing range.

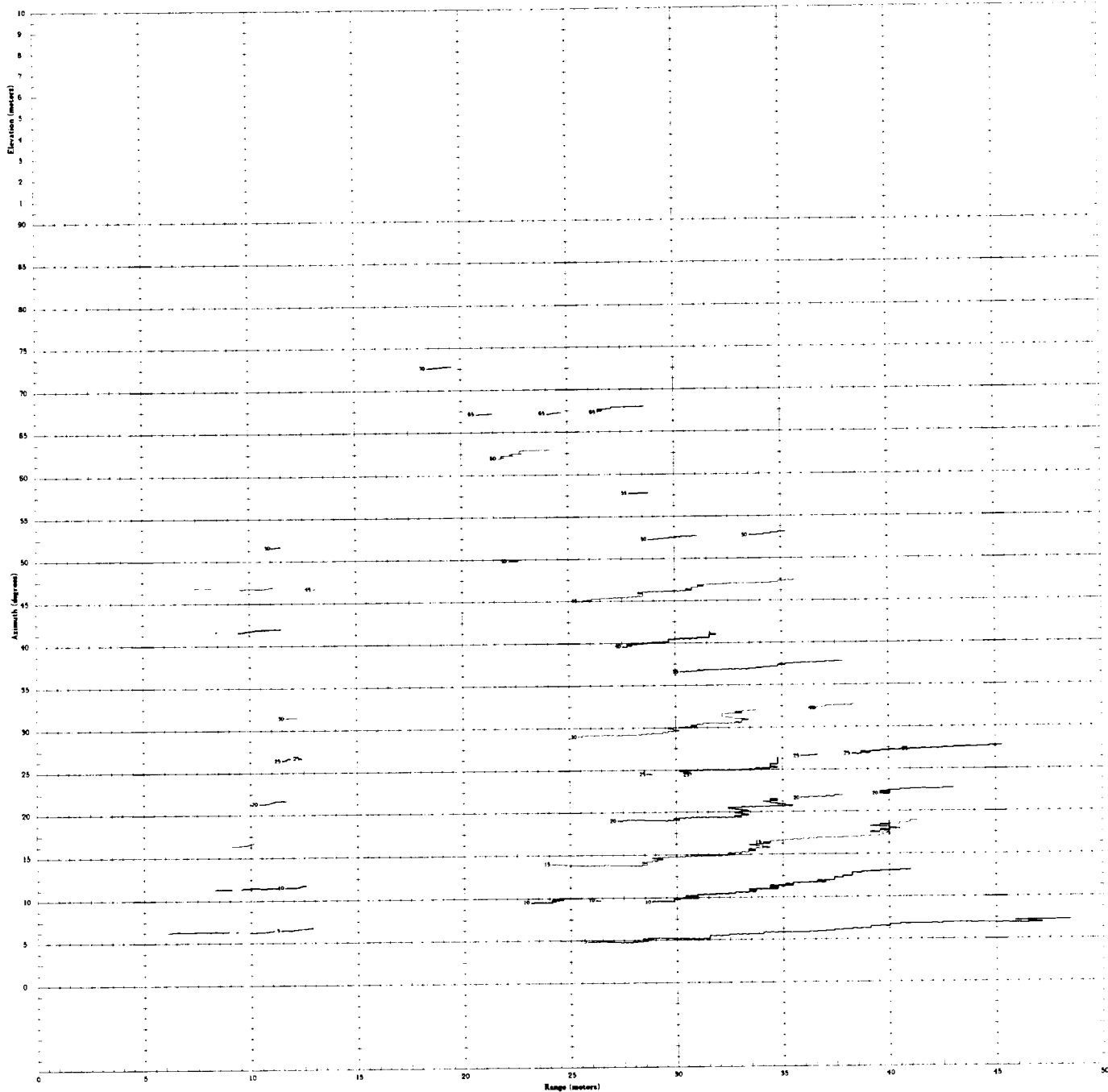


Plot parameters relate to the Local Gravity Reference (LGR) coordinate system. The LGR system is orthogonal and right-handed. The origin coincides with that of the Lander Reference Coordinate System (LRCS). The 2-axis points toward the zenith. The 1-axis points in the direction of the projection of the LRCS 2-axis toward the horizon plane through the origin. The plot distance (horizontal range) is the distance from the LGR 2-axis. The ordinate (elevation) is the value of the LGR 1-coordinate. The area of elevation for each profile is covered by the associated profile azimuth range. Azimuth is measured (measured) from above, resulting in a series of reference along the LGR negative 1-axis.

VIKING LANDER 1 SYSTEMATIC VERTICAL PROFILES SCALE 1:1000 RANGE 0 TO 500 METERS 0 TO 90 DEGREES

Range Data Set Files
 1. ADD: EDI: FRONT 1195A
 1. ADD: EDI: FRONT 1195B
 1. ADD: EDI: BACK 1195C
 1. ADD: EDI: BACK 1195D

These maps were produced by J. Lander, utilizing
 resources of the Computer Department and the
 Physics Department, University of Southern
 California. They are based on data generated at the Jet
 Propulsion Laboratory, under contract NAs1-6657
 and JPL 751101, and from JPL 751101.
 JPL 751101

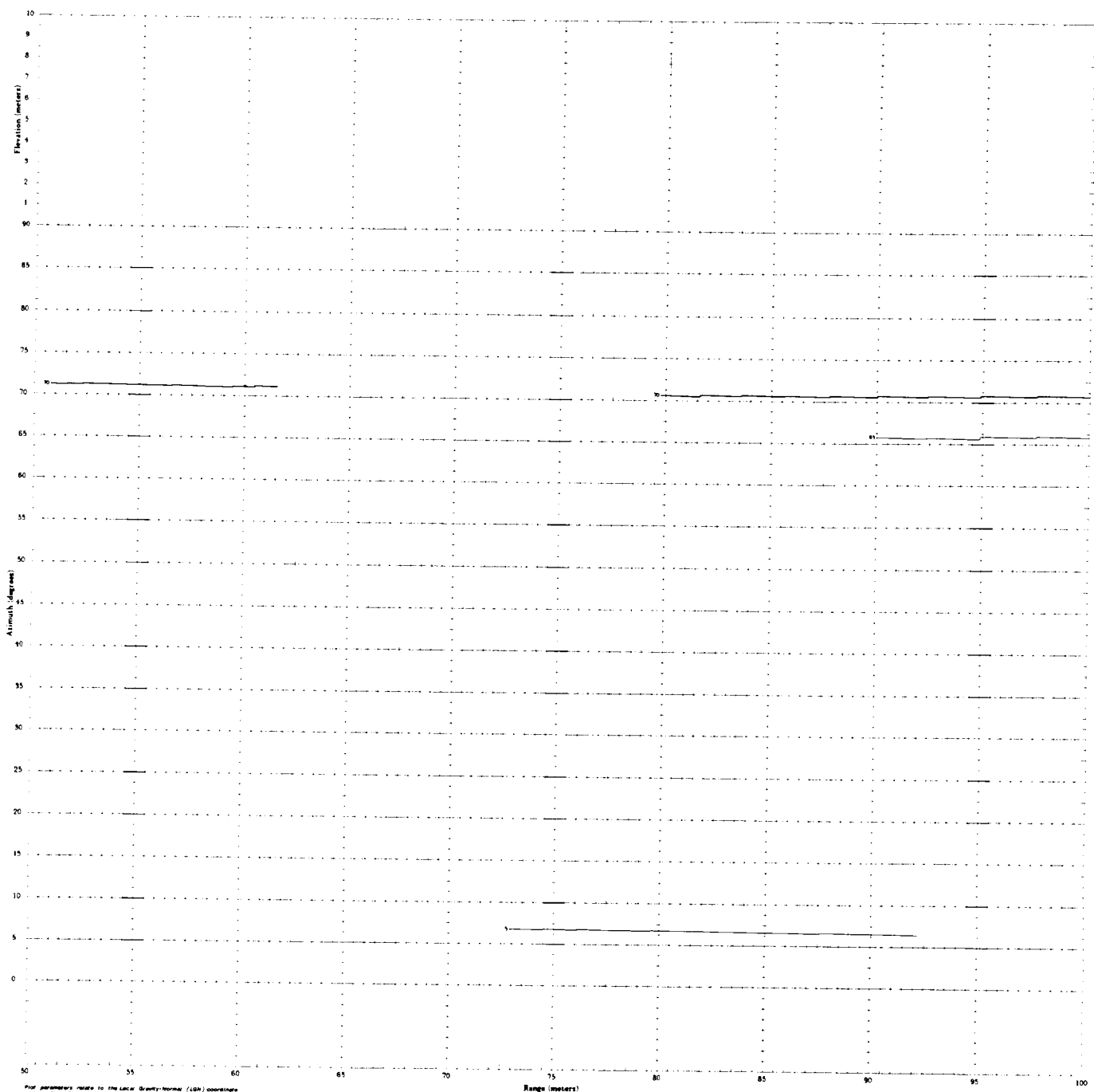


Four parameters relate to the Local Gravity-Inertial (LGI) coordinate system. The LGI system is orthogonal and right-handed. The origin coincides with that of the Lander-Airframe Coordinate System (LACS). The Z-axis points toward the planet. The X-axis points in the direction of the perpendicular of the LACS Z-axis normal to the horizontal plane through the origin. The Y-axis points in the direction of the LACS Z-axis. The distance (range) is the value of the LGI Z-coordinate. The slope of elevation for each profile is obtained by the associated profile screen page. Elevation is measured clockwise from level, relative to a zero of reference along the LGI negative Z-axis.

VIKING LANDER 1 SYSTEMATIC VERTICAL PROFILES SCALE 1:100 RANGE 0 TO 50 METERS 0 TO 90 DEGREES

Range Data Set File
L1 ADDS FOR FRONT L197A
L1 ADDS FOR BACK L197B
L1 ADDS FOR BACK L197C
L1 ADDS FOR BACK L197D

These maps were produced by 3 Lander Mission
teams of the Lander Operations and the
Biological Investigation Laboratory. Images
obtained from a scan line generated in the
Lander Laboratory were processed by the
Lander Mission and processed by the
Lander Mission.

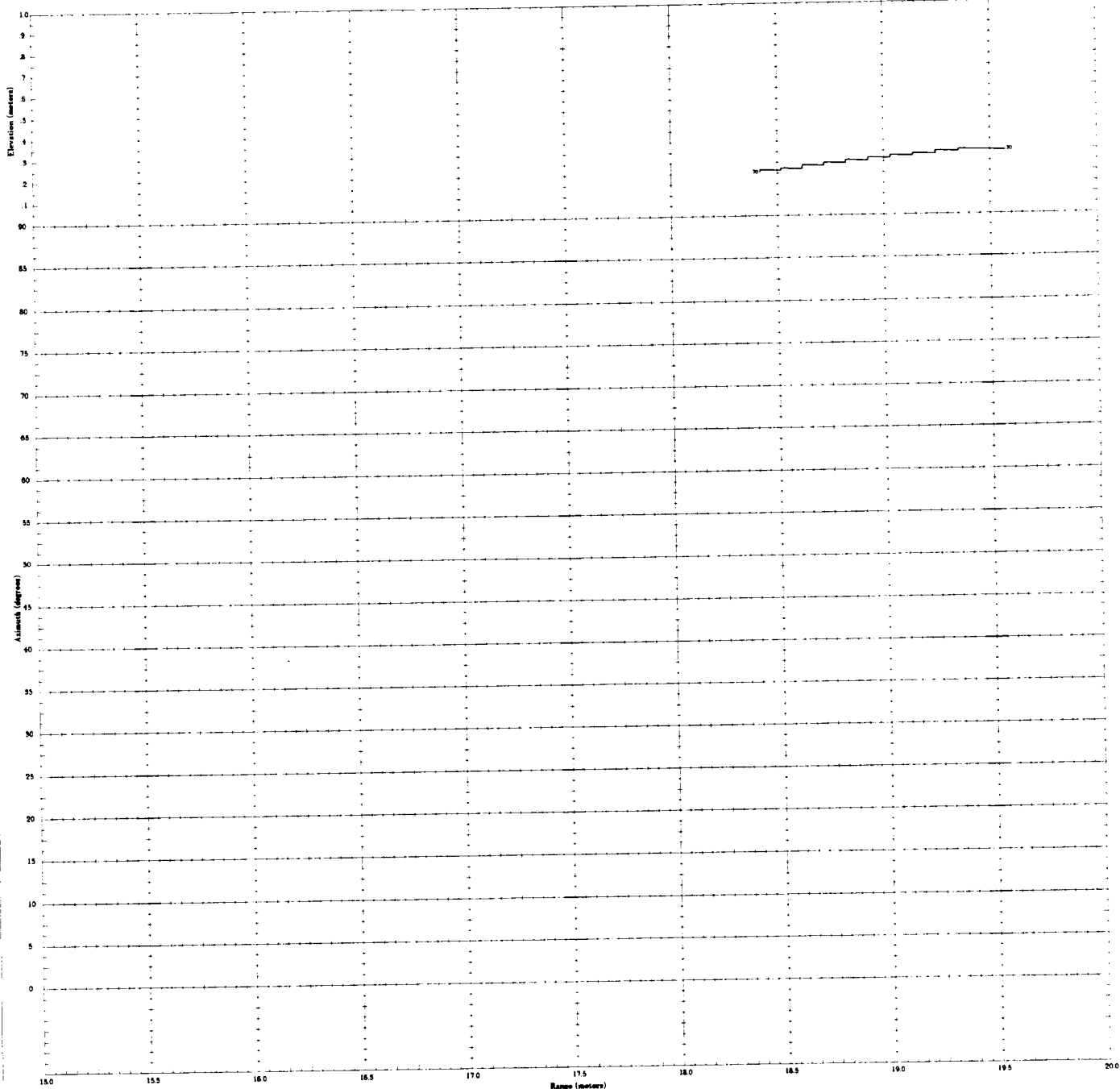


For parameters used in the Lander Gravity System (LGS) coordinate system. The LGS system is orthogonal and right handed. The origin coincides with that of the Lander Inertial Coordinate System (LICS). The 2 axis points toward the south. The 1 axis points in the direction of the projection of the LGS 2 axis onto the horizontal plane through the origin. The plot abscissa (horizontal range) is the distance from the LGS 1 axis. The ordinate (elevation) is the value of the LGS 2 coordinate. The range of elevation for each profile is shown by the associated profile azimuth angle. Azimuth is measured clockwise from above relative to a pair of reference lines, the LGS negative 1 axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:100 RANGE 50 TO 100 METERS 0 TO 90 DEGREES

Range Data Set File
JPL
L1-RSV-EDR-PROF1 1978
L1-RSV-PROF1 1978
L1-RSV-BACK 1978
L1-RSV-EDR-BACK 1978

These data were generated by a data setting system of the Earth Department and the Department of Geology, University of California, San Diego. The data were generated by the JPL Propulsion Laboratory under contract NAs1-101-001 and JPL 55170P and from NAS 150-1000000. APRIL 1980.

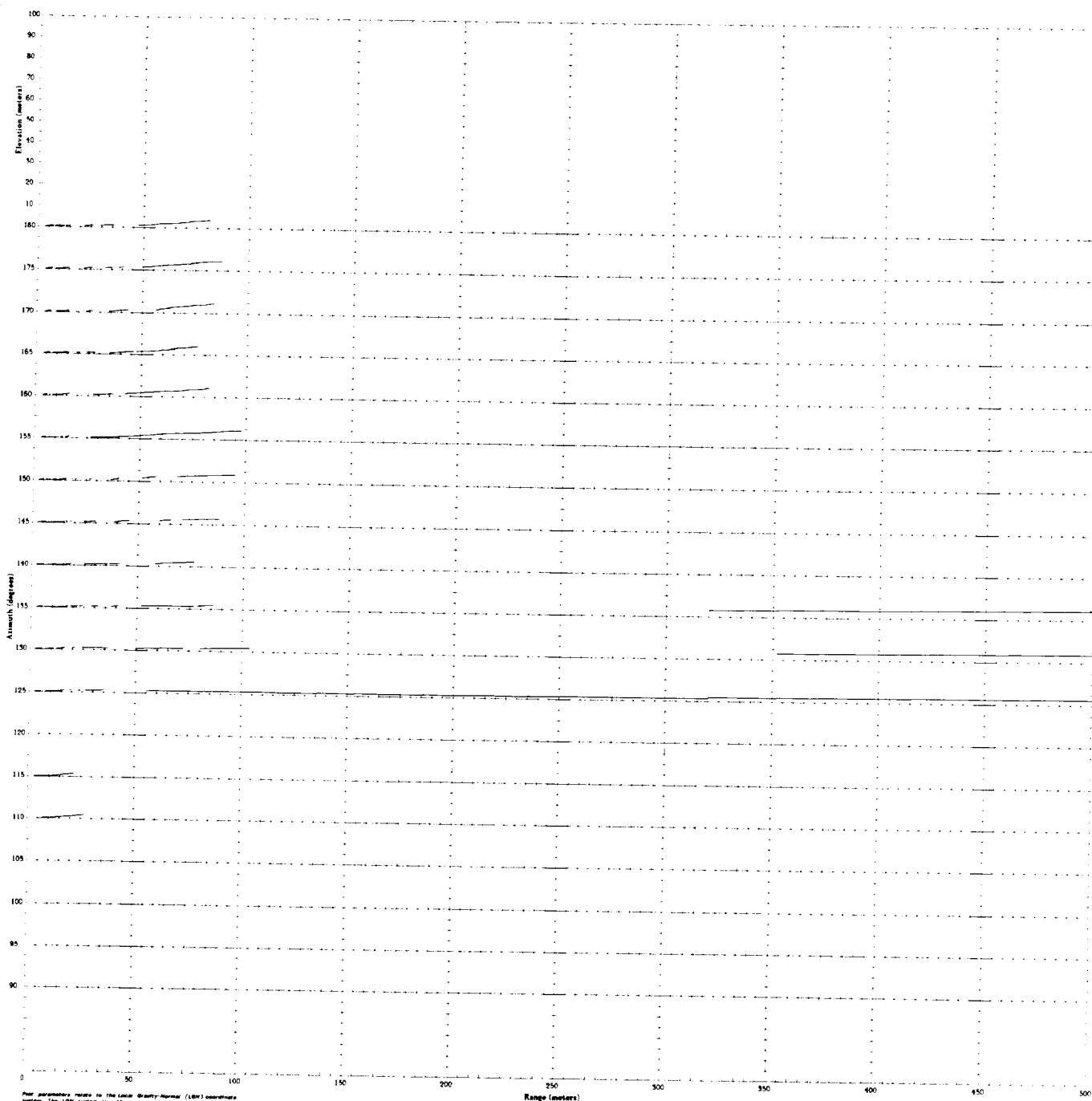


For parameters used in the Lunar Gravity-Sensor (LGS) coordinate system. The LGS system is orthogonal and right-handed. The origin coincides with that of the Lunar Aligned Coordinate System (LACS). The Z-axis points toward the earth. The Y-axis points in the direction of the projection of the LACS Z-axis normal to the horizontal plane through the origin. The X-axis (horizontal range) is the distance from the LGS Z-axis. The elevation (elevation) is the angle of the LGS Z-axis. The range of elevation for each profile is labeled by the associated profile number. Elevation is measured clockwise from above, relative to a zero of reference using the LGS negative Y-axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 15 TO 20 METERS 0 TO 90 DEGREES

Range Data Set Files
JPL
L1 RDDP 10M FRONT 1197A
L1 RDDP 10M REAR 1197B
L1 RDDP 10M BACK 1198A
L1 RDDP 10M BACK 1198B

This study was produced by 1. Lander, utilizing
images of the Camera System and the
Altimeter System. Images were
University (from 2 min data generated at the
University Laboratory) under contract N-65-6042
and JPL 951249 and from N-65-7142
N-65-7142

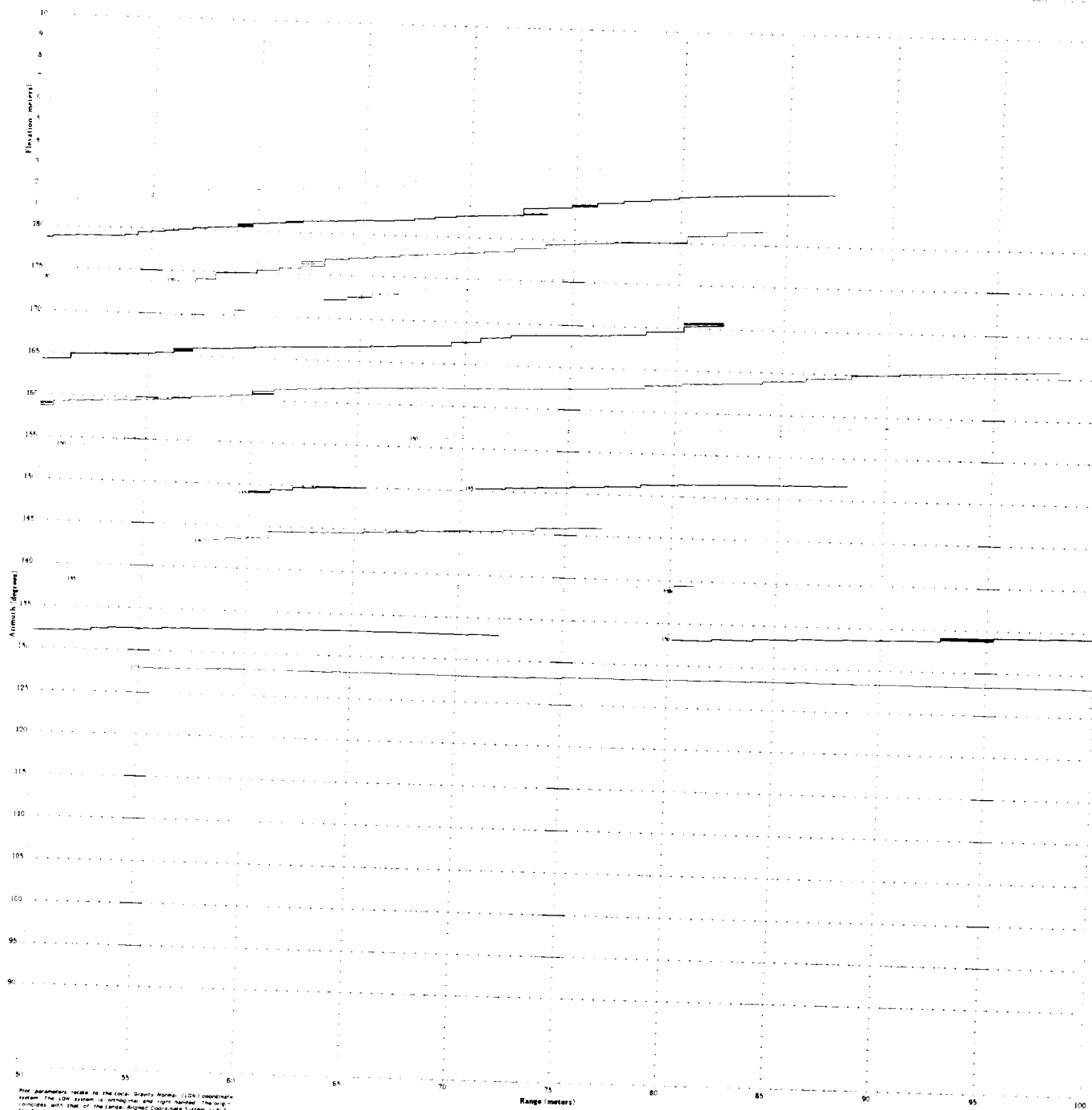


Plot parameters relative to the local Gravity-thermal (LGT) coordinate system. The LGT system is orthogonal and right-handed. The origin coincides with that of the Lander Instrument Coordinate System (LICS). The Z axis points toward the zenith. The Y axis points in the direction of the projection of the LICS Z axis normal to the horizontal plane through the origin. The plot abscissa (horizontal range) is the distance from the LGT Z axis. The ordinate (elevation) is the value of the LGT Z coordinate. The range of elevation for each profile is obtained by the associated profile azimuth angle. Azimuth is measured clockwise from above, relative to a zero of reference along the LGT negative Y axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1 : 1000 RANGE 0 TO 500 METERS 90 TO 180 DEGREES

Range Data Set File
JPL
17 ADDV LDR FRONT 179A
17 ADDV FRONT 179B
17 ADDV BACK 179C
17 ADDV LDR BACK 179D

These maps were produced by 1. Load, editing, reviewing, if the Lander 1, Department and the associated instruments, cameras, scanning equipment from a data set generated by the JPL Propulsion Laboratory's main computer, PDS-1, and JPL PDS-1 and then by JPL PDS-1.

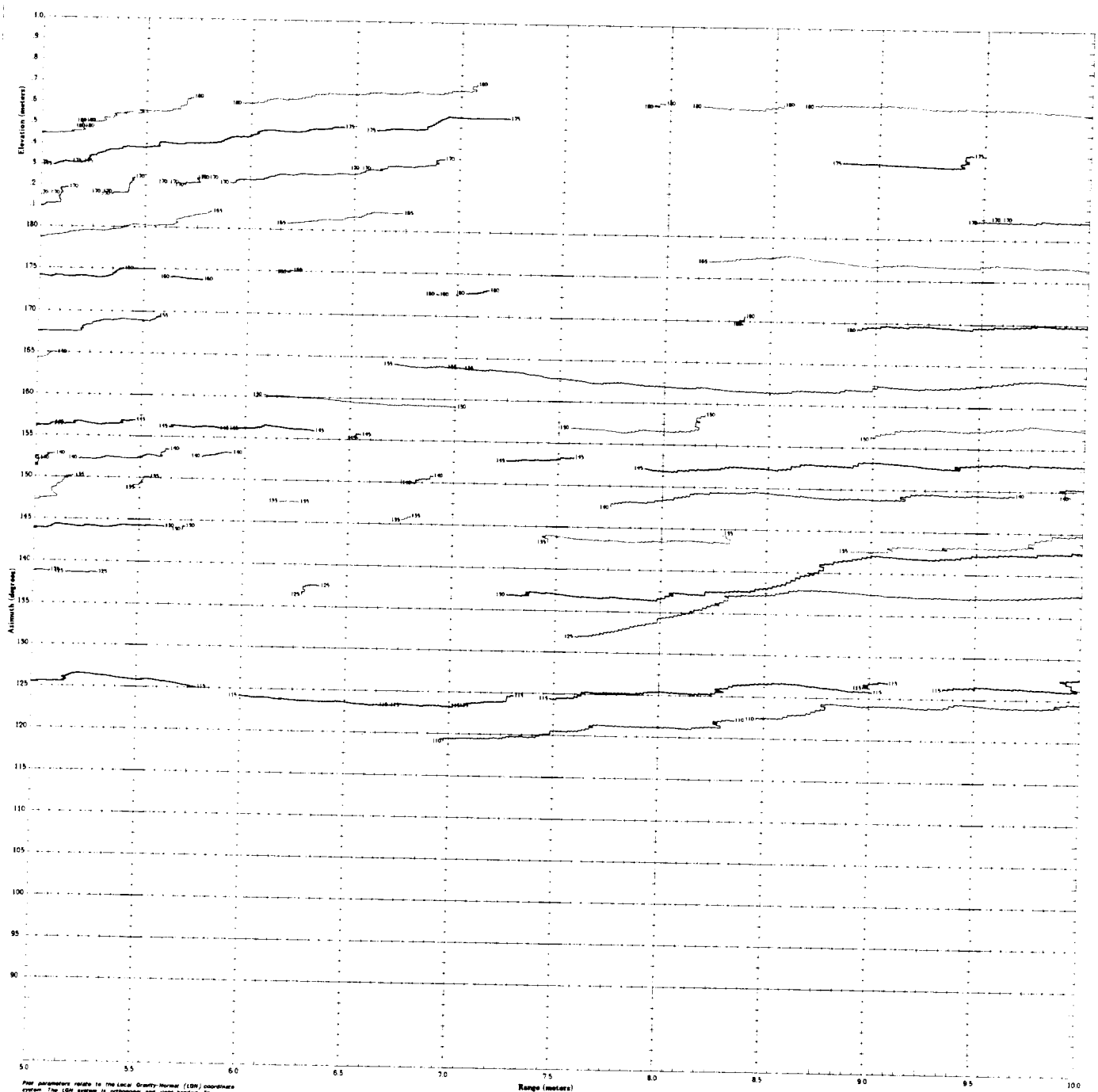


VIKING LANDER 1 SYSTEMATIC VERTICAL PROFILES

SCALE 1:100 RANGE 50 TO 100 METERS 90 TO 180 DEGREES

Range Data Set Files	
JPL	Stanford
LT.RDSV.EDR.FRONT	LTIFA
LT.RDSV.FRONT3	LTIFP
LT.RDSV.BACK	LTIBP
LT.RDSV.EDR.BACK	LTIBA

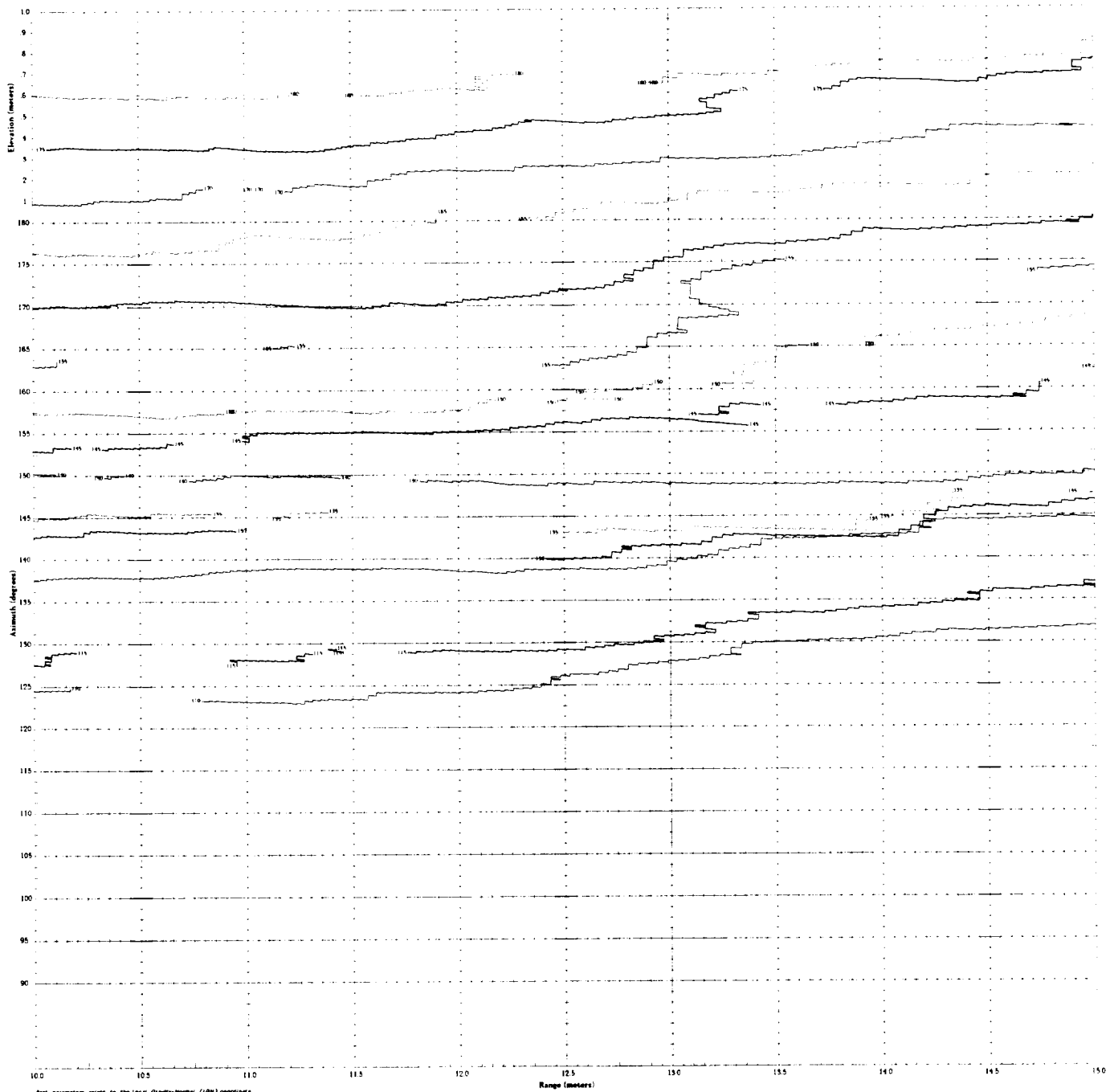
These maps were produced by S. Lohr utilizing resources of the Genetic Engineering and the Artificial Intelligence Laboratory, Stanford University from a data base generated at the Jet Propulsion Laboratory under contract A33-668-1 and JPL 67-146 and grant NSC 1722.



VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGES 5 TO 10 METERS 90 TO 180 DEGREES

Range Data Set Files
JPL
1) ADVI FOR FRONT 1978
1) ADVI FOR BACK 1978
1) ADVI FOR BACK 1984

These maps were produced by S. Lander, utilizing
software of the Lander, Department of
Geology, University of California, San Diego
University. They are not intended to be
used for navigation purposes. Under contract N00019-78-1-0017
and N00019-79-1-0017.

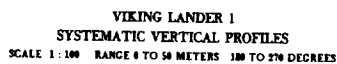


Profile parameters relate to the Local Gravity-Normal (LGN) coordinate system. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lander-Instrument Coordinate System (LICS). The Z-axis points toward the south. The X-axis points in the direction of the projection of the LICS Z-axis normal to the horizontal plane through the origin. The Y-axis (normal range) is the distance from the LGN Z-axis. The elevation (elevation) is the value of the LGN Z-coordinate. The range of elevation for each profile is labeled by the associated profile number range. Elevation is measured clockwise from above, relative to a zero of reference using the LGN negative Y-axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 10 TO 15 METERS 90 TO 130 DEGREES

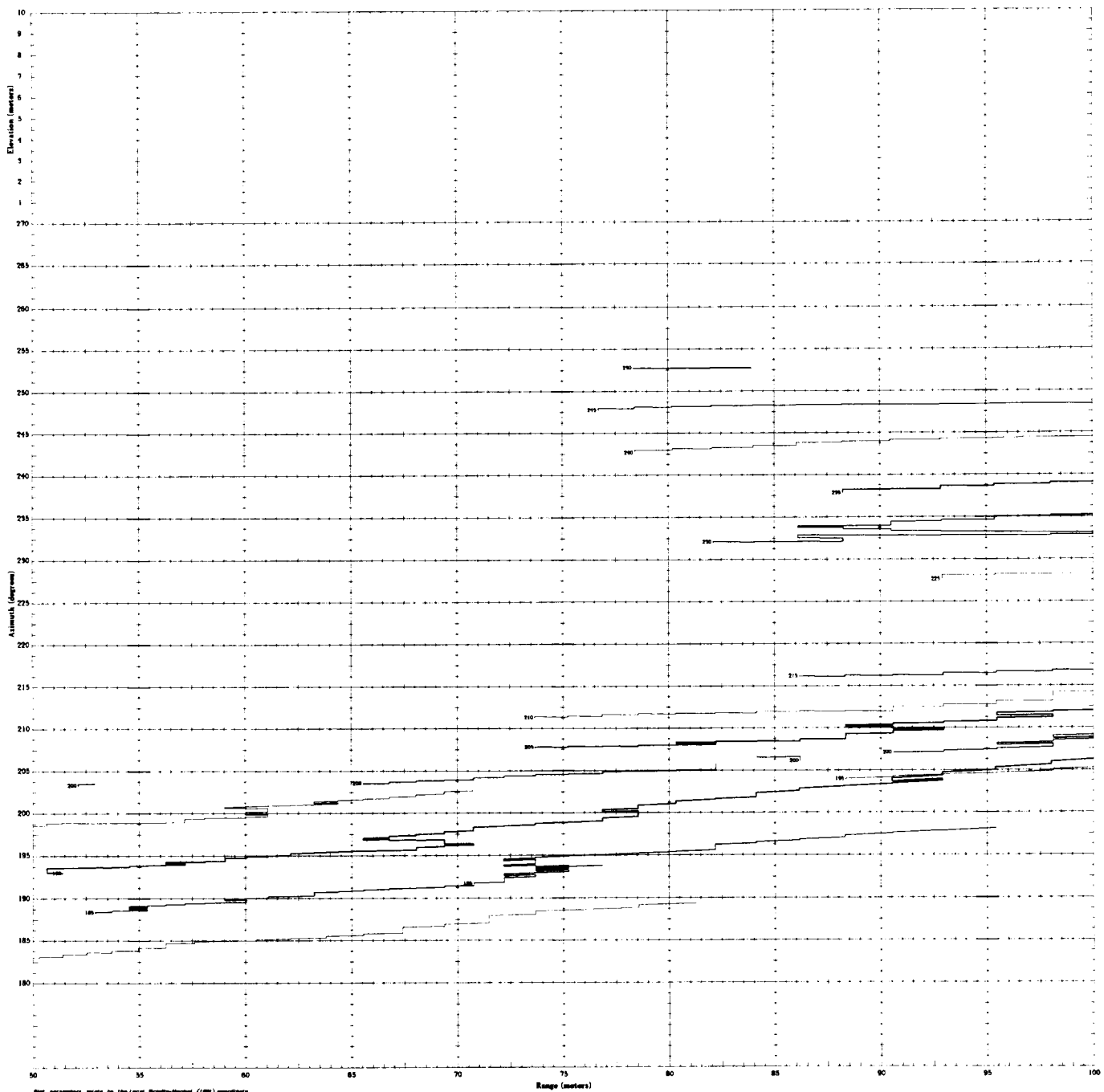
Range Data Set File
 1) ADV FOR FACT 1974
 1) ADV FOR FACT 1974
 1) ADV BACK 1974
 1) ADV FOR BACK 1974

These maps were produced by 3 Lander software systems: the Lander Software and the Lander Software Library. The Lander Software Library is a data base generated at the Jet Propulsion Laboratory under contract NAS-7-602-001. The Lander Software Library is a data base generated at the Jet Propulsion Laboratory under contract NAS-7-602-001. The Lander Software Library is a data base generated at the Jet Propulsion Laboratory under contract NAS-7-602-001.



Range Date Set Files	
JPL	Stanford
L1.RDSV.ZDR.FRONT	L1VFA
L1.RDSV.FRONT3	L1VFP
L1.RDSV.BACK	L1VBP
L1.RDSV.ZDR.BACK	L1VBA

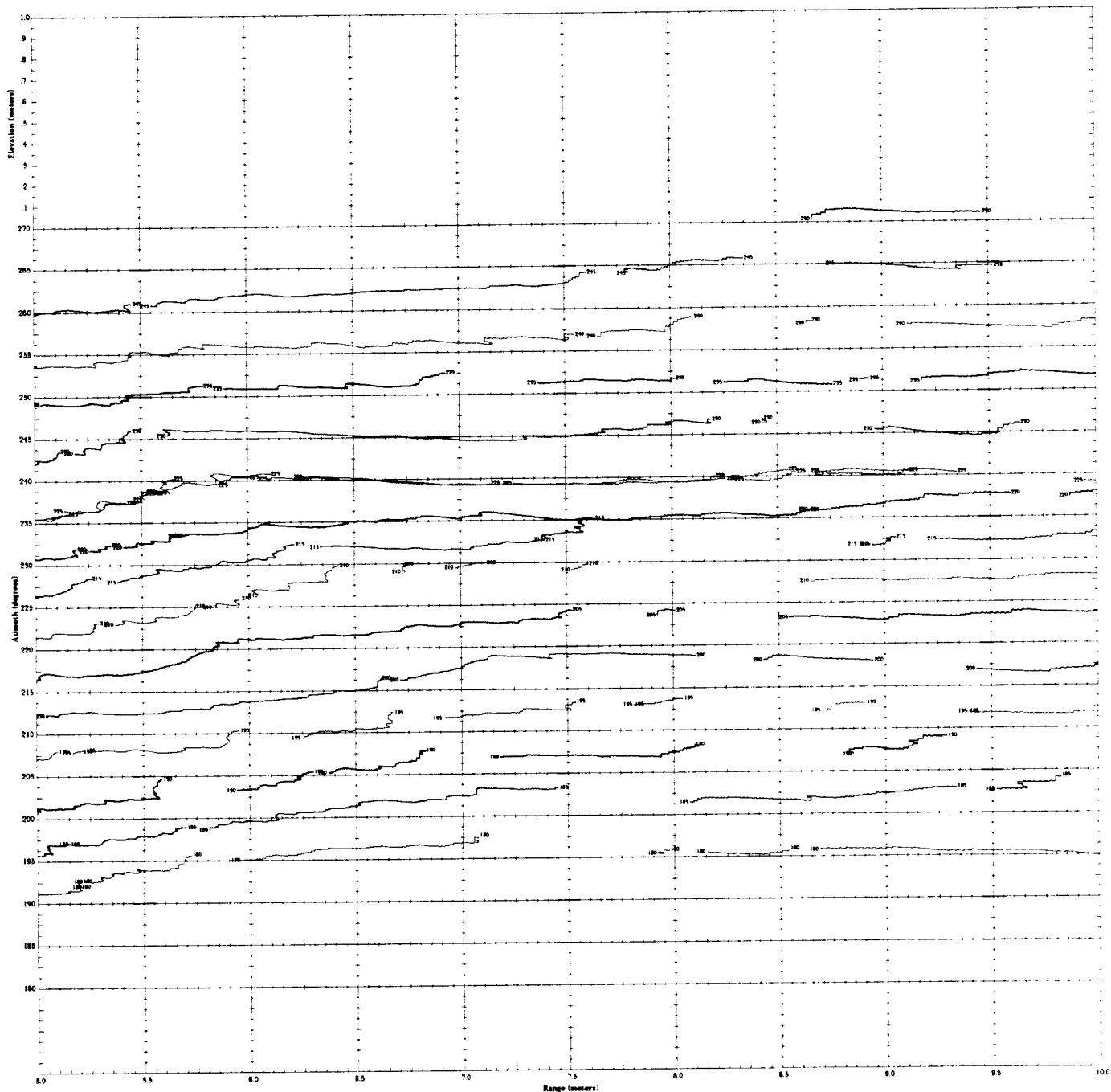
These maps were produced by S. Lieber, utilizing resources of the Central Department and the Analytical Intelligence Laboratory, Stanford University (from a data base generated at the JPL Propulsion Laboratory); under contracts #AS7-9603 and JPL 933249, and grant N5C 7938.



VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:100 RANGE 50 TO 100 METERS 180 TO 310 DEGREES

Range Data Set Files
L1:RDSV 180 FRONT L1:RDSV 180 BACK
L1:RDSV 190 FRONT L1:RDSV 190 BACK
L1:RDSV 200 FRONT L1:RDSV 200 BACK
L1:RDSV 210 FRONT L1:RDSV 210 BACK
L1:RDSV 220 FRONT L1:RDSV 220 BACK
L1:RDSV 230 FRONT L1:RDSV 230 BACK
L1:RDSV 240 FRONT L1:RDSV 240 BACK
L1:RDSV 250 FRONT L1:RDSV 250 BACK
L1:RDSV 260 FRONT L1:RDSV 260 BACK
L1:RDSV 270 FRONT L1:RDSV 270 BACK

These maps were produced by S. Liden, utilizing
software of the Canadian Department of the
Geological Survey, and the
University of Toronto. The data were generated at the
University of Toronto, using programs P-271-001
and P-271-002, and were processed by S. Liden, 1980.

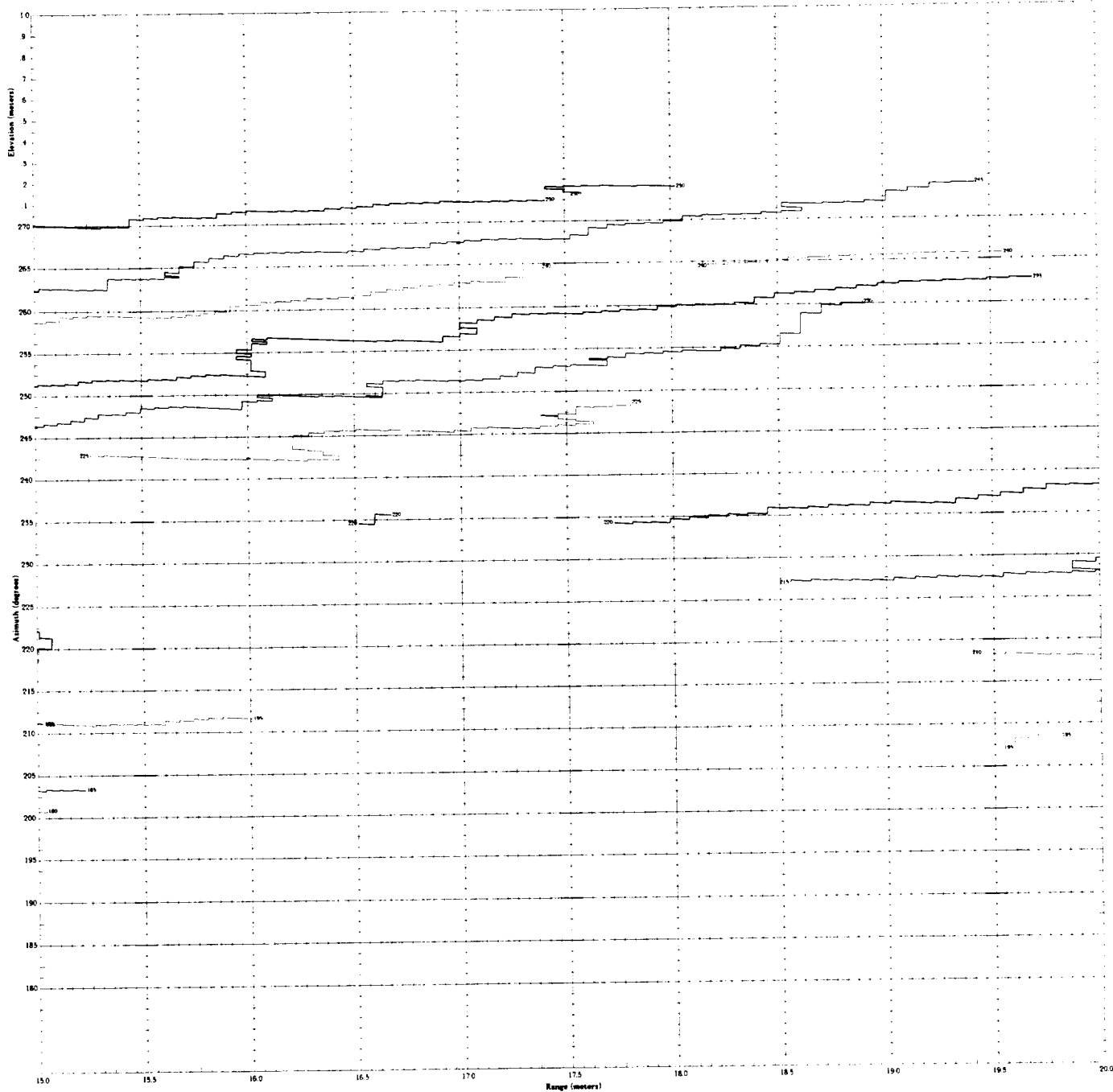


Plot parameters relate to the Lander Global Reference (LGR) coordinate system. The LGR system is orthogonal and right-handed. The origin coincides with that of the Lander Reference Coordinate System (LACS). The Z-axis points toward the planet. The Y-axis points in the direction of the projection of the LACS Z-axis normal to the horizontal plane through the origin. The plot displays (horizontal range) in the distance from the LGR Z-axis. The ordinate (elevation) is the value of the LGR Z-coordinate. The date of acquisition for each profile is indicated by the sequential profile azimuth angle. Azimuth is measured clockwise from ahead, relative to a zero of reference using the LGR negative Y-axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 5 TO 10 METERS 180 TO 270 DEGREES

Range Data Set Files
JPL Superfund:
L1/ASV/LAND FRONT 1197A
L1/ASV/FRONT 1197B
L1/ASV/LAND BACK 1198A
L1/ASV/LAND BACK 1198B

These maps were prepared by J. Lander, Mission Manager, of the Science Department and the Engineering Department, Mars Exploration Rover Mission, NASA/JPL. The data were generated at the Jet Propulsion Laboratory under contract N-001-60-1 and JPL 551279 and given NAC 7350 180-00000000.

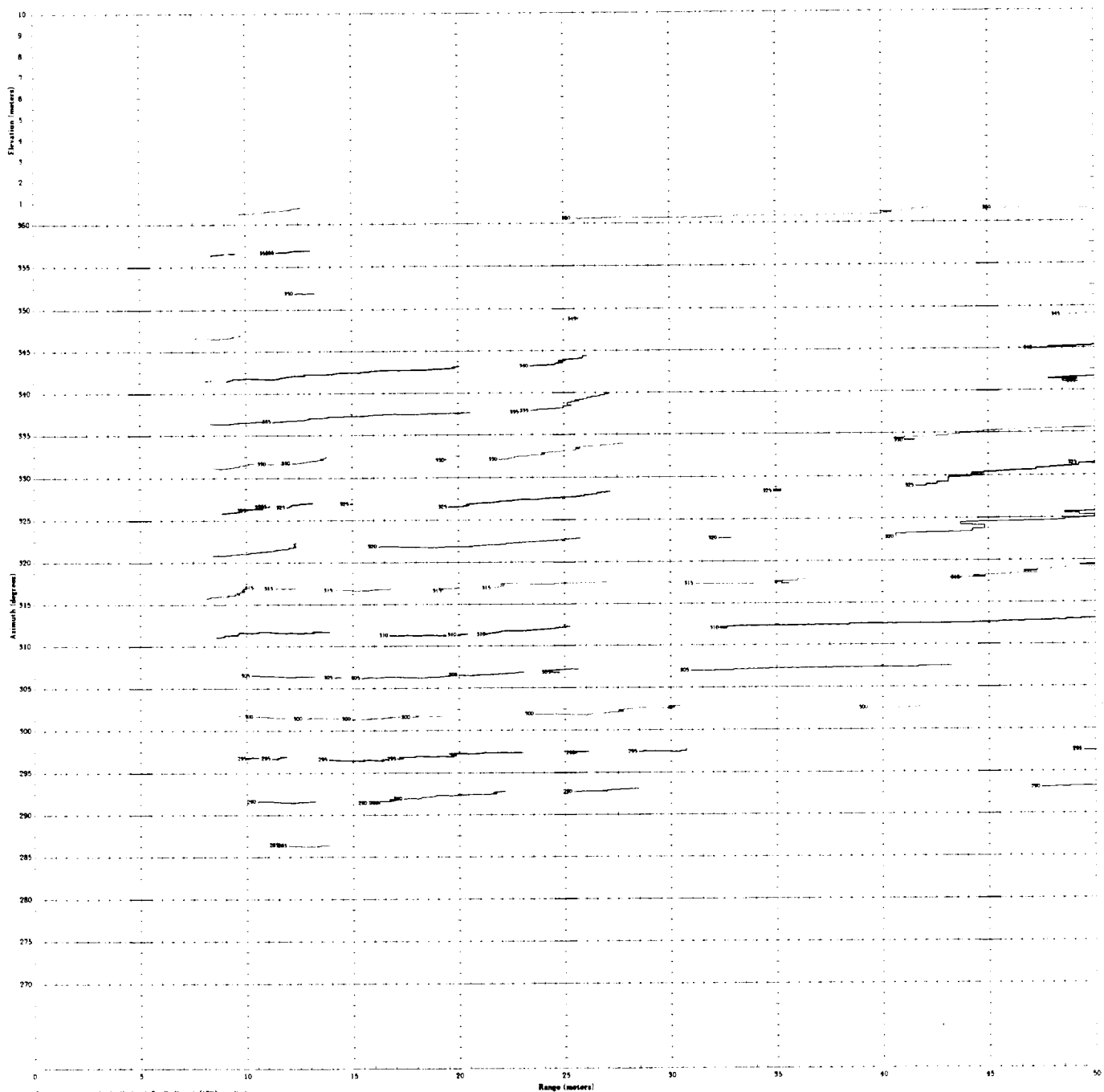


Plot parameters relate to the Local Gravity-Warner (LGR) coordinate system. The LGR system is orthogonal and right-handed. The origin coincides with that of the Lander Relative Coordinate System (LRCS). The Z-axis points toward the center. The X-axis points in the direction of the projection of the LRCS Z-axis normal to the horizontal plane through the origin. The plot abscissa (horizontal range) is the distance from the LGR Z-axis. The ordinate (elevation) is the value of the LGR Z-coordinate. The zero of elevation for each profile is labeled by the abscissa profile azimuth angle. Azimuth is measured counter-clockwise from above, relative to a zero of reference along the LGR negative Y axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 15 TO 20 METERS 180 TO 210 DEGREES

Range Data Set File
JPL
L1 RDSV FOR FRONT L178
L1 RDSV FRONT L178
L1 RDSV BACK L178
L1 RDSV FOR BACK L178

These maps were produced by J. Lander, using
images of the Lander's cameras and the
Angstrom, Infrared, and Ultraviolet
cameras. The maps were produced at the
University of Arizona, Tucson, Arizona.
The maps were produced on 11/22/78.
The maps were produced on 11/22/78.
The maps were produced on 11/22/78.

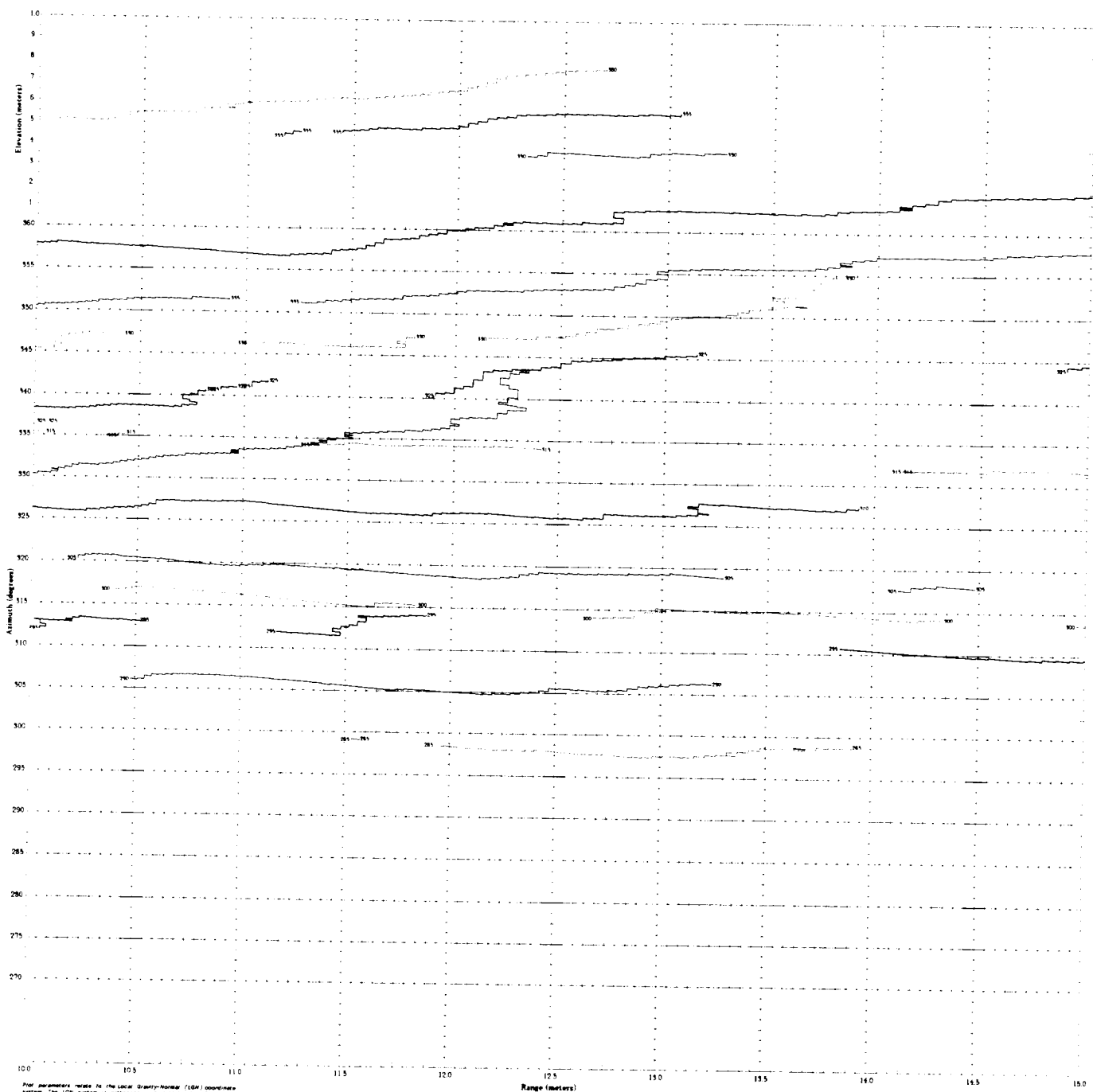


For parameters, refer to the Local Gravity-Vertical (LGV) coordinate system. The LGV system is orthogonal and right-handed. The origin coincides with that of the Lunar Reference Coordinate System (LRCS). The Z-axis points toward the center of the Moon. The X-axis points in the direction of the projection of the LRCS Z-axis onto the horizontal plane through the origin. The Y-axis points in the direction of the projection of the LRCS X-axis onto the horizontal plane through the origin. The elevation (elevation) is the value of the LGV Z-coordinate. The range of elevation for each profile is indicated by the second and third vertical angles. Elevation is measured clockwise from above, relative to a zero of reference along the LGV negative Y-axis.

VIKING LANDER 1 SYSTEMATIC VERTICAL PROFILES SCALE 1:100 RANGE 0 TO 50 METERS 270 TO 360 DEGREES

Range Data Set File
JPL
L1: RGSZ FOR FRONT 11/78
L1: RGSZ FOR BACK 11/78
L1: ADVZ FOR BACK 11/78
L1: ADVZ FOR BACK 11/78

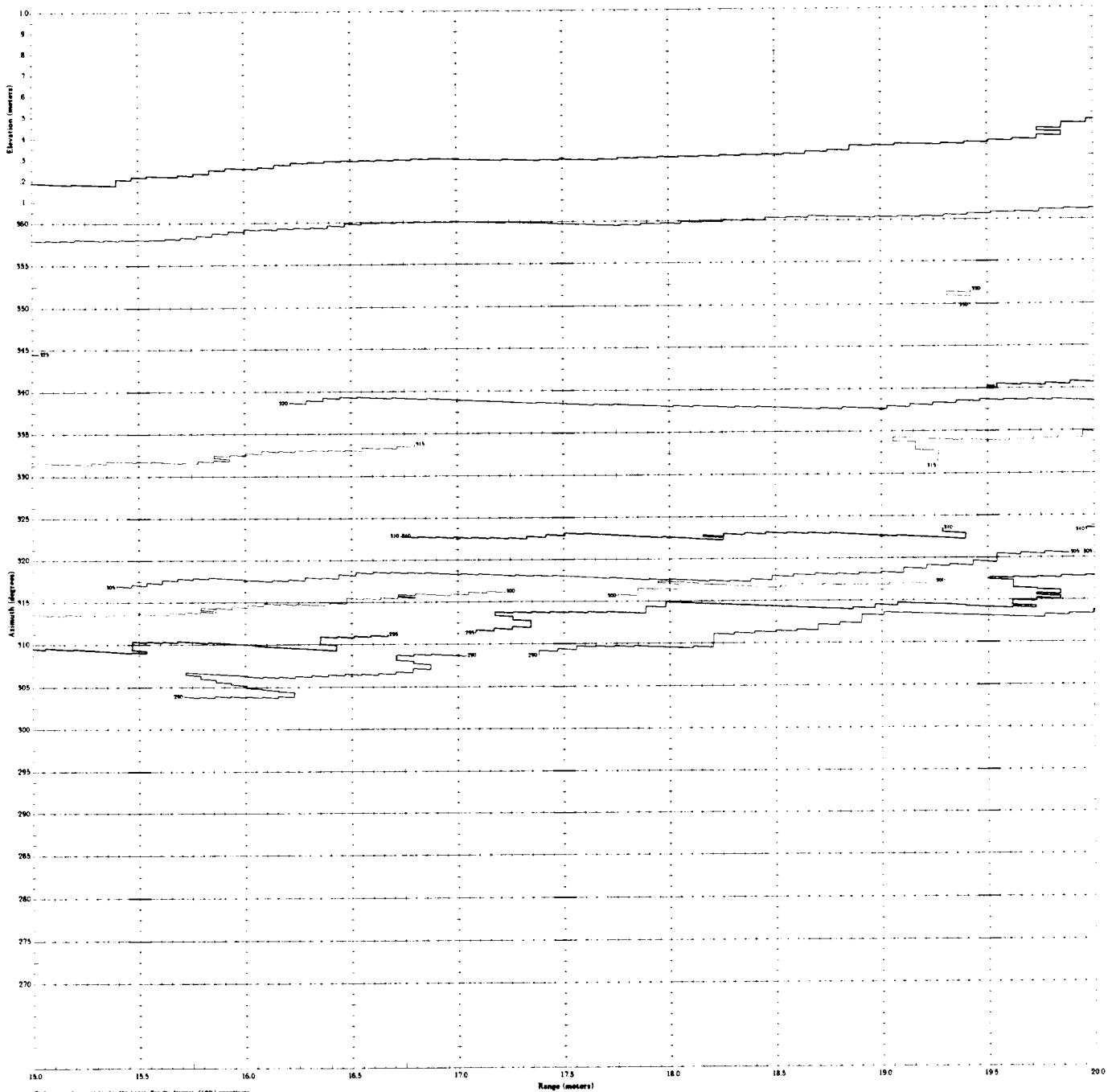
These maps were produced by J. Smith, utilizing resources of the Geomatics Department and the Jet Propulsion Laboratory. The maps were produced at the Jet Propulsion Laboratory, under contract NAs-13-100-001, and JPL D-151149, and given NAs-13-100-001.



VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 10 TO 15 METERS 270 TO 360 DEGREES

Range Data Set Title
JPL
ADSV-EDM FRONT
ADSV-EDM BACK
ADSV-EDM BACK

These maps were produced by J. Lander, adding
numbers of the Lander's drawings and the
official photographs. Lander's original
drawings (then a map) were generated at the JPL
Propulsion Laboratory under contract N0001-67-1-001
and JPL 77-74, and given to JPL 77-74.



Plot parameters relate to the Local Gravity-Header (LGH) instrument system. The LGH system is orthogonal and right-handed. The origin coincides with that of the Lander-Inertial Coordinate System (LICS). The 2 axis points toward the south. The 1 axis points in the direction of the direction of the LICS 2-axis normal to the horizontal plane through the origin. The plot abscissa (horizontal range) is the distance from the LGH 1-axis. The ordinate (elevation) is the value of the LGH 2-coordinate. The scale of elevation for each profile is defined by the associated profile abscissa angle. Abcissa is measured counter-clockwise from above, relative to a zero of reference along the LGH negative 1-axis.

VIKING LANDER 1
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 15 TO 20 METERS 270 TO 360 DEGREES

Range Data Set Title
JPL Standard
L1 ADDV LDR FRONT L177A
L1 ADDV FRONT L177A
L1 ADDV BACK L178P
L1 ADDV LDR BACK L178A

These data were produced by 3 Lander working
members of the Lander Instrument and
Software Engineering Laboratory, Space
Science Applications, Inc., under the direction of the JPL
Project Manager, under the direction of the JPL
and JPL 1977 and JPL 1978
contractors.

6. ELEVATION CONTOURS

6.1 Elevation Contour Mosaic Overlay Stereo Pairs

This section contains mosaic stereo pairs of images into which have been overlayed the systematic elevation contour lines. The general remarks of section 4.1 apply.

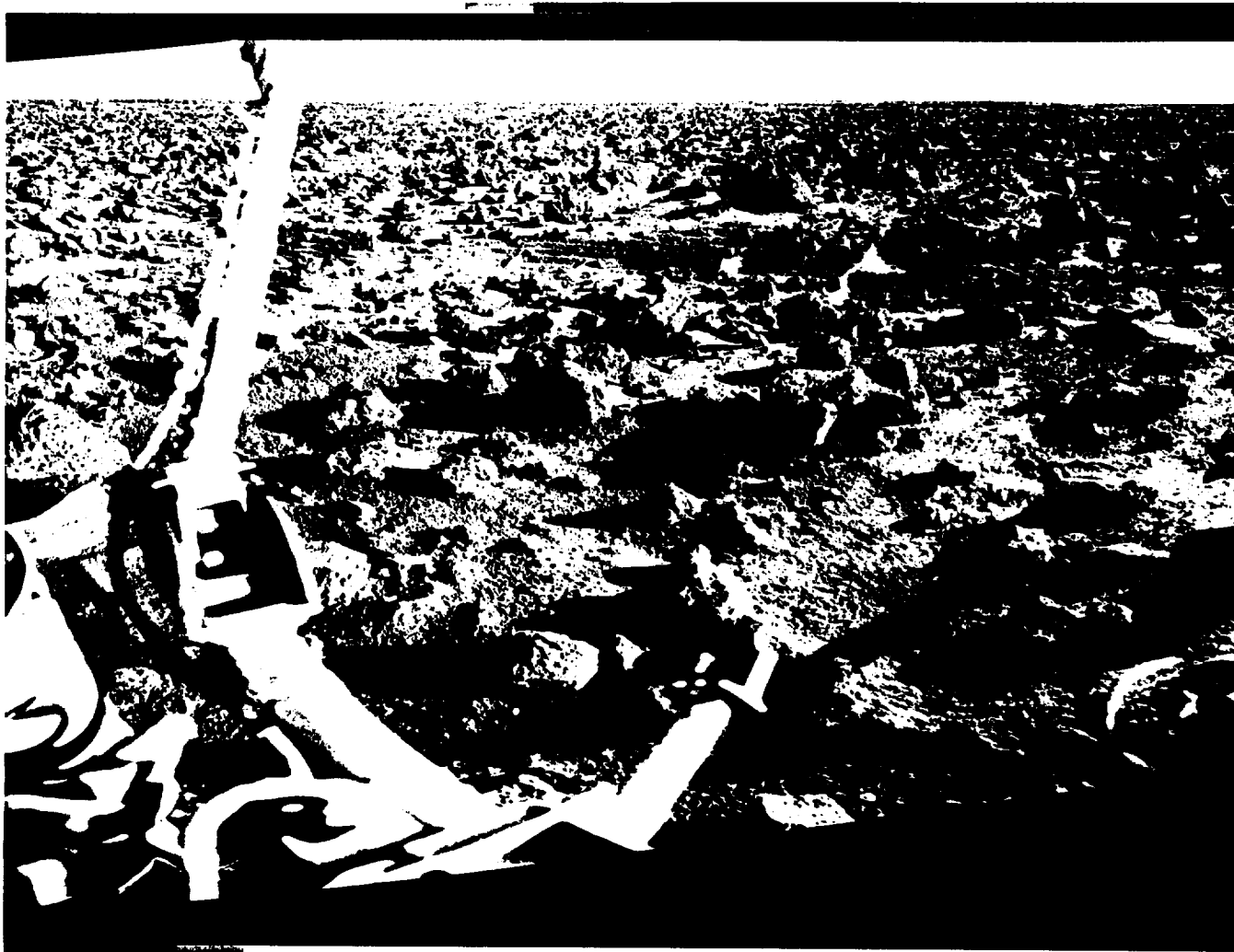


Fig. 35. Elevation Contour Overlay Mosaic: Camera 1 (left), front left quadrant —
from IPL PIC ID 79/10/11/055137.



Fig. 36. Elevation Contour Overlay Mosaic: Camera 2 (right), front left quadrant — from IPL PIC ID 79/03/15/022628.

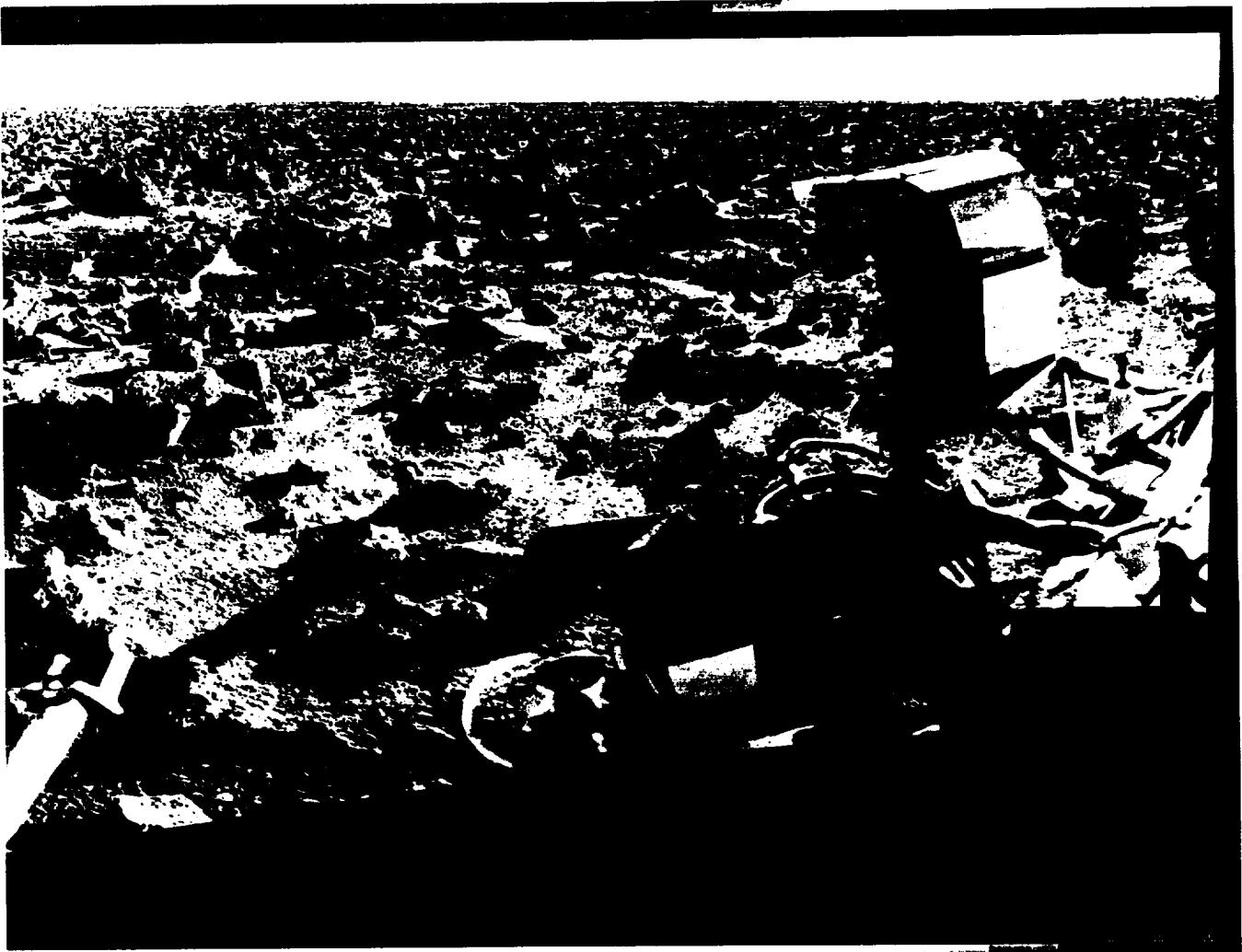


Fig. 37. Elevation Contour Overlay Mosaic: Camera 1 (left), front right quadrant — from IPL PIC ID 79/10/11/055137.



Fig. 38. Elevation Contour Overlay Mosaic: Camera 2 (right), front right quadrant —
from IP1, PIC ID 79/03/15/022628.



Fig. 39. Elevation Contour Overlay Mosaic: Camera 2 (left), back left quadrant —
from IPL PIC ID 79/10/08/233417.

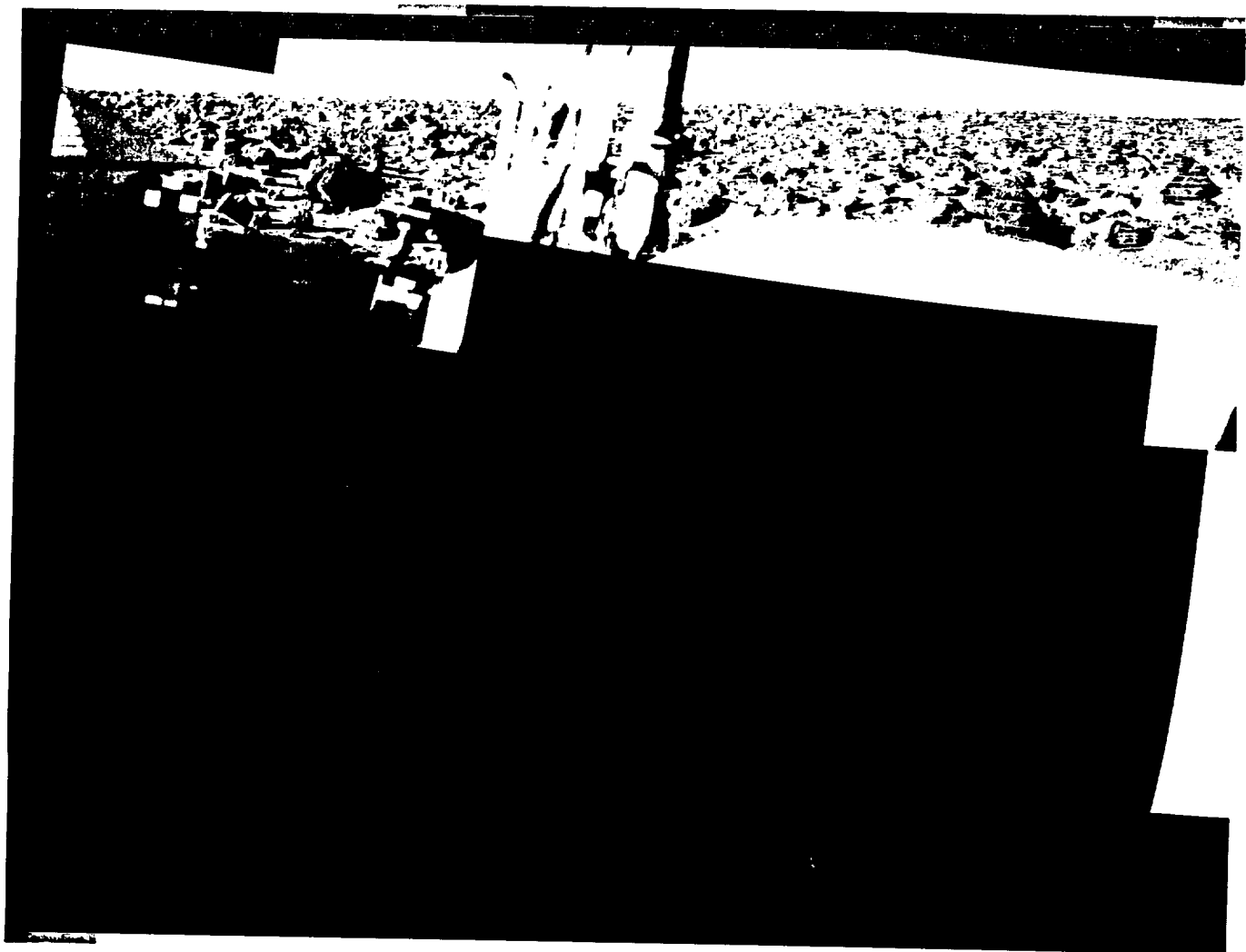


Fig. 40. Elevation Contour Overlay Mosaic: Camera 1 (right), back left quadrant —
from IPL PIC ID 79/04/04/030319.

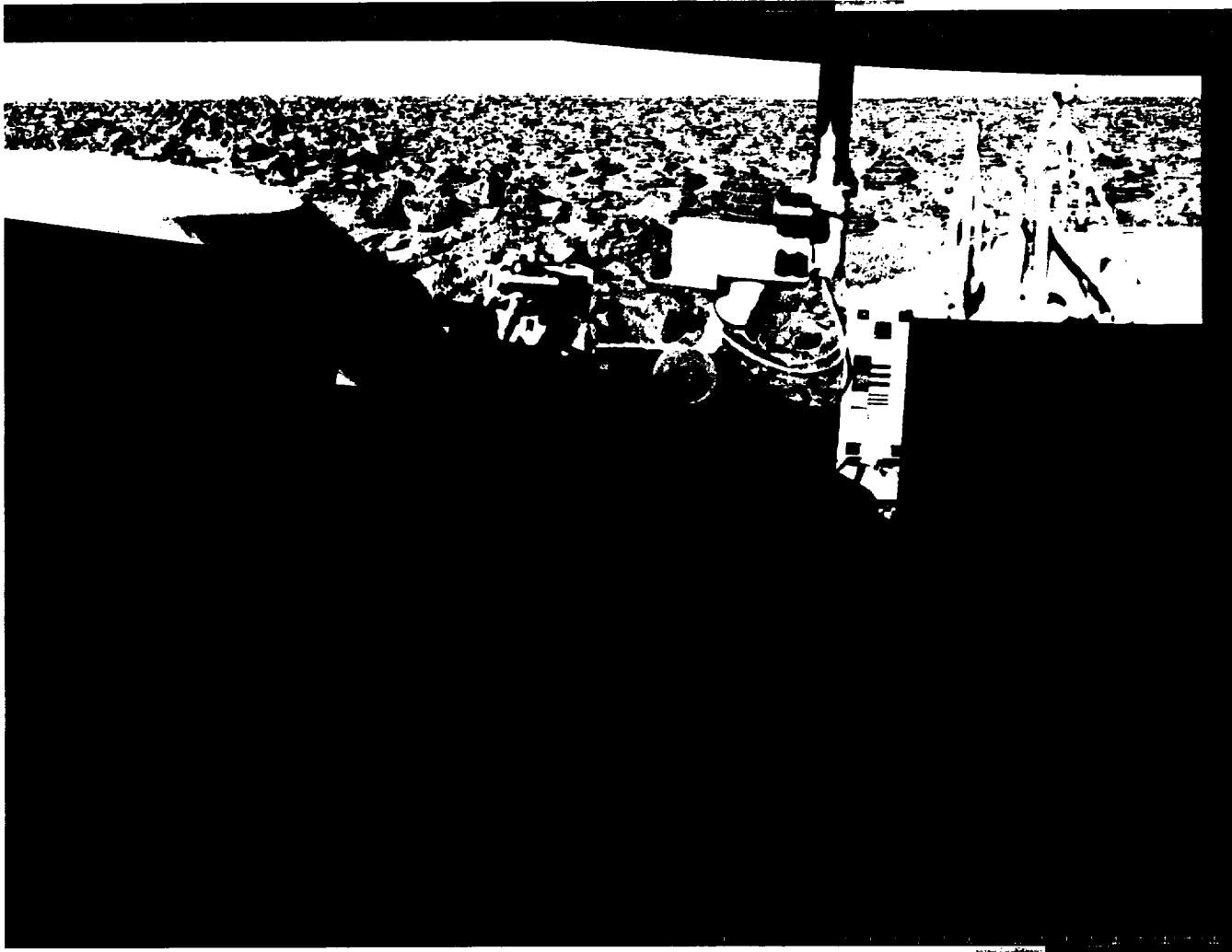


Fig. 41. Elevation Contour Overlay Mosaic: Camera 2 (left), back right quadrant —
from IPL PIC ID 79/10/08/233417.

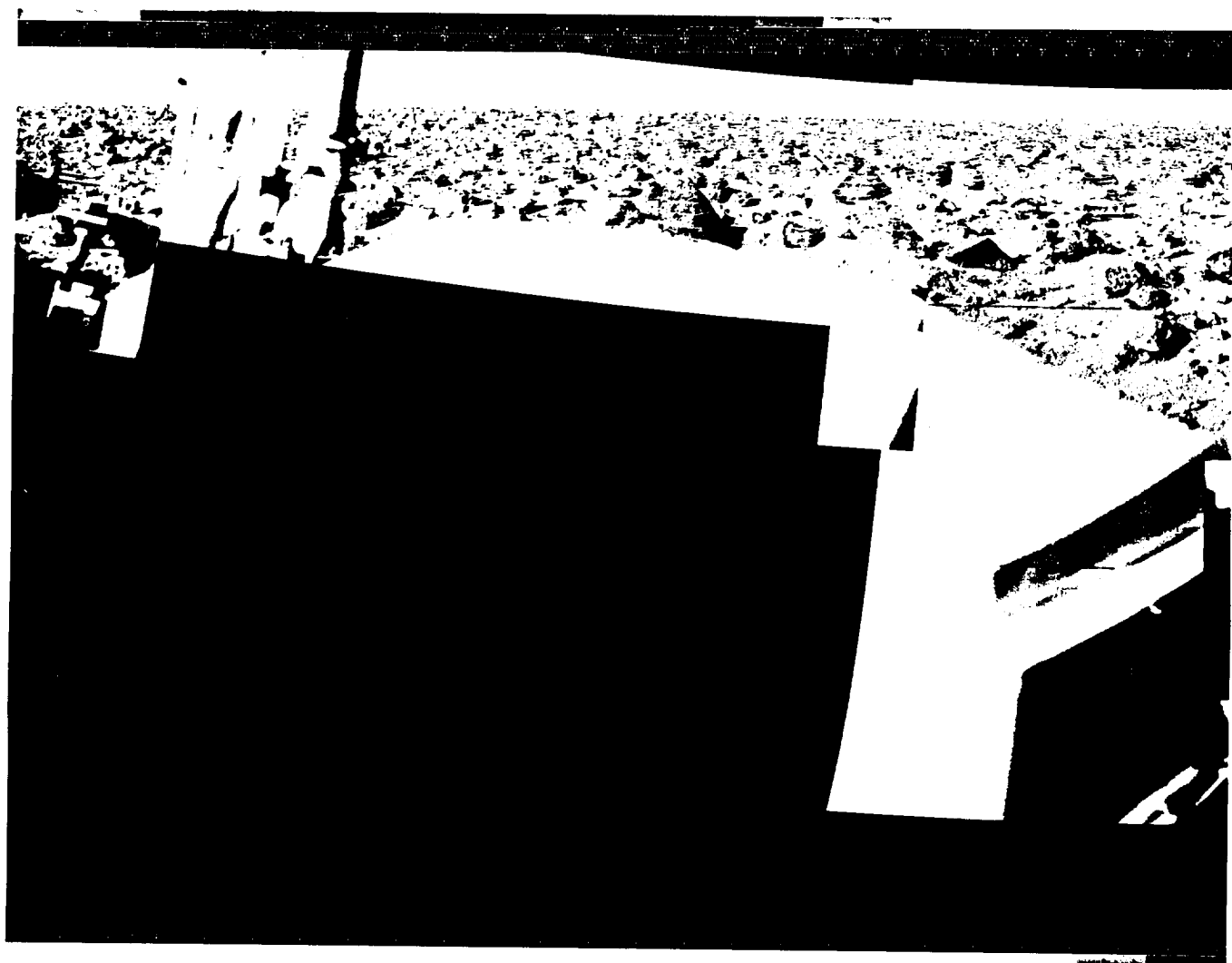


Fig. 42. Elevation Contour Overlay Mosaic: Camera 1 (right), back right quadrant --
from IPL PIC ID 79/04/04/030319.

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6.2 Camera Perspective Annotated Elevation Contours — Camera 1

This section contains camera-1 perspective representations of the elevation contour map data. The nature and purpose of these representations has been explained in section 3.2.1. On opposing pages are presented mosaic overlays of the elevation con-

tour lines, and corresponding camera perspective representations of the map data, with the individual lines of the map data annotated with elevation values. The general discussion of section 4.1 applies to the mosaics of this section.



Fig. 43. Elevation Contour Overlay Mosaic: Camera 1 (left), front left quadrant – from IPL PIC ID 79/10/11/055137.

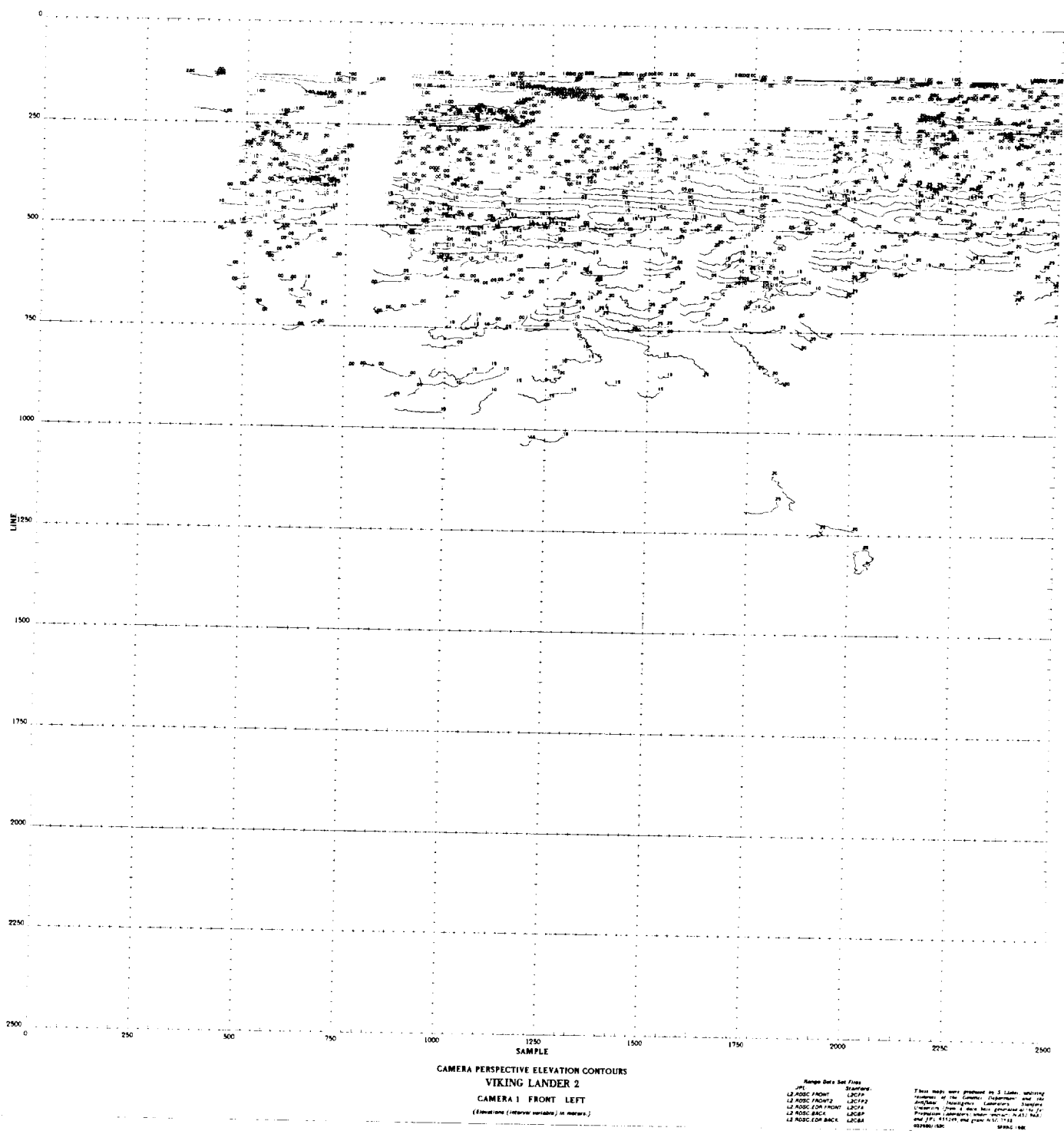




Fig. 45. Elevation Contour Overlay Mosaic: Camera 1 (left), front right quadrant — from IPL PIC ID 79/10/11/055137.

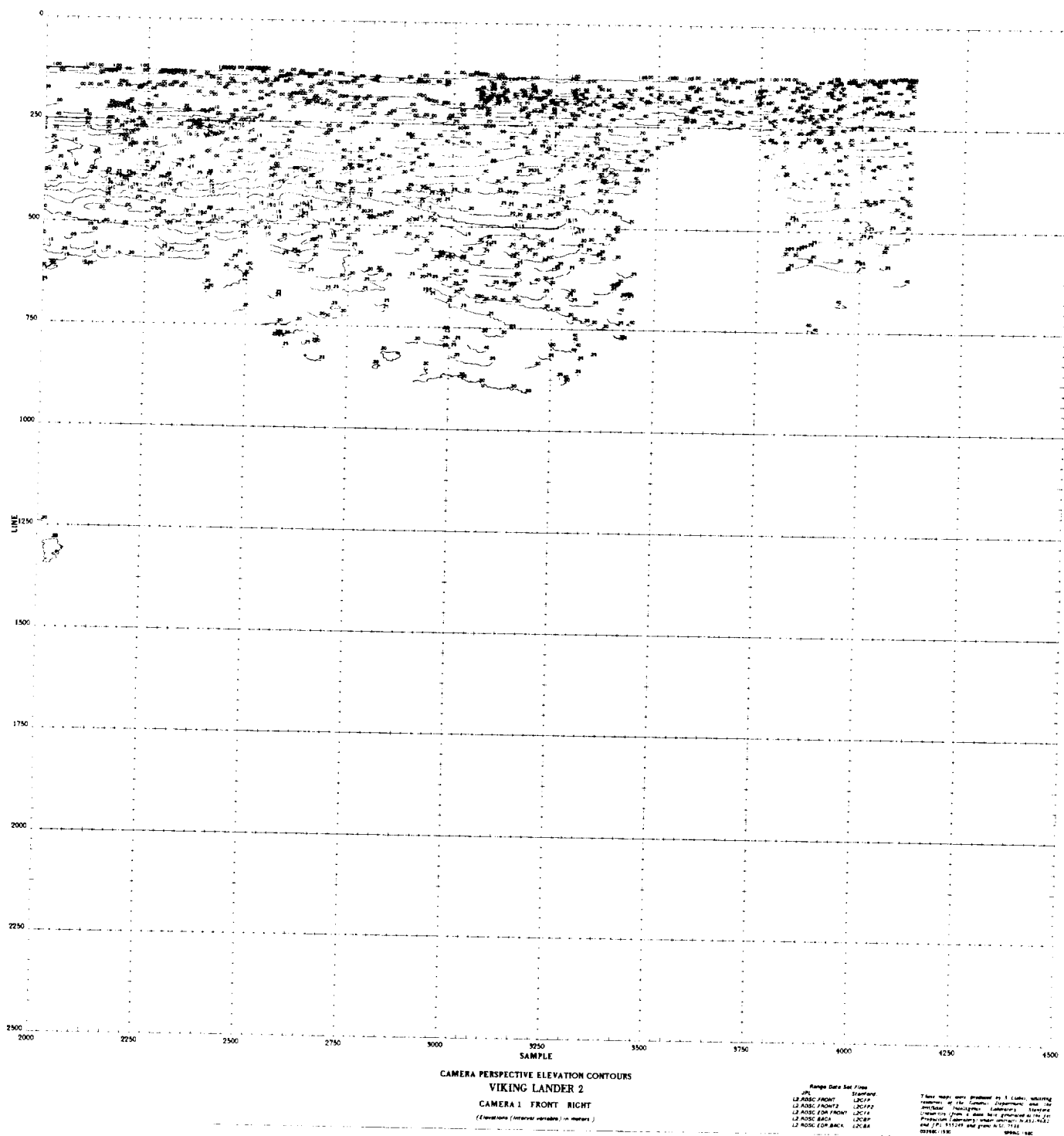


Fig. 46. Camera Perspective Annotated Elevation Contour Lines:
Camera 1 (left), front right quadrant.

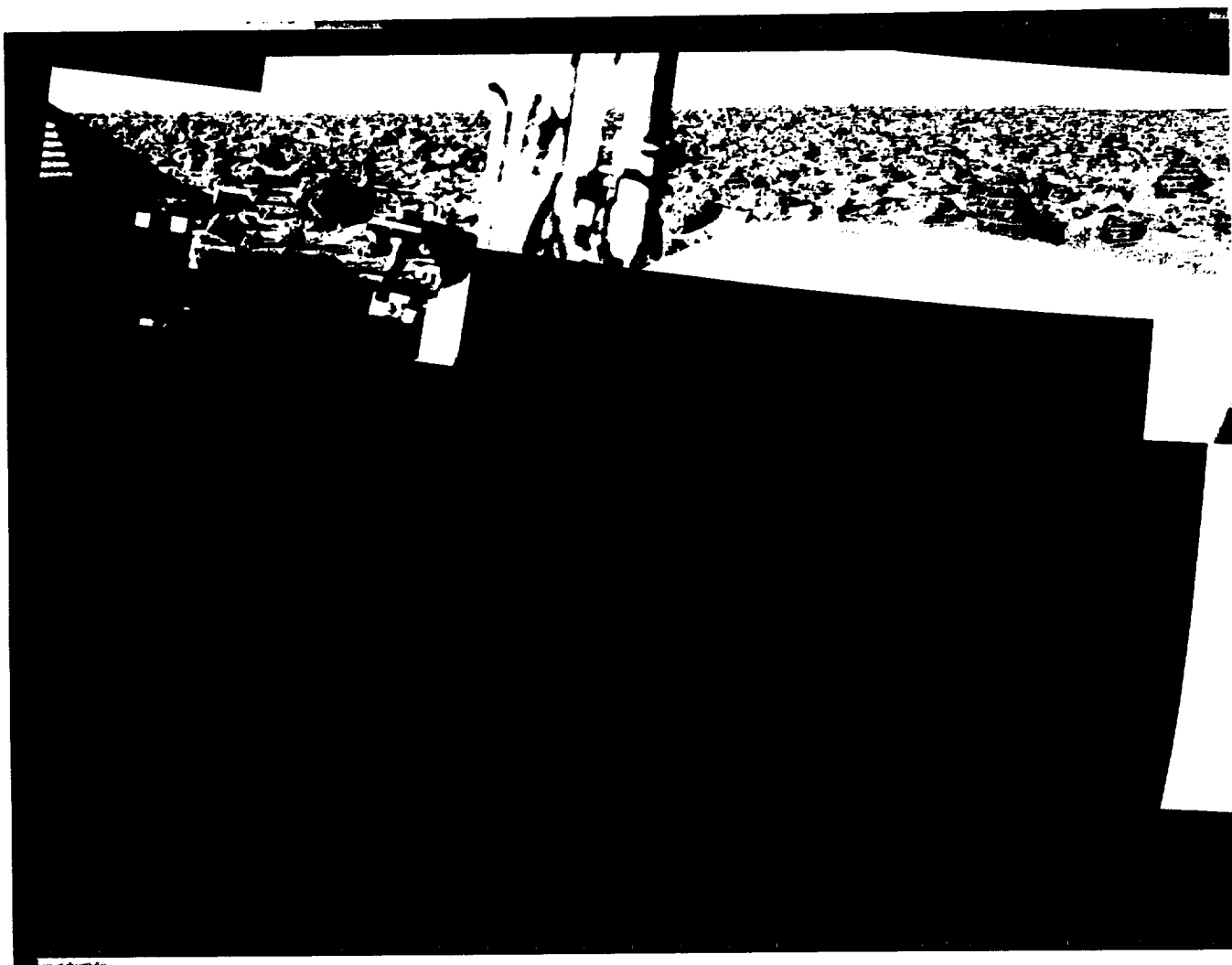


Fig. 47. Elevation Contour Overlay Mosaic: Camera 1 (right), back left quadrant —
from IPL PIC ID 79/04/04/030319.

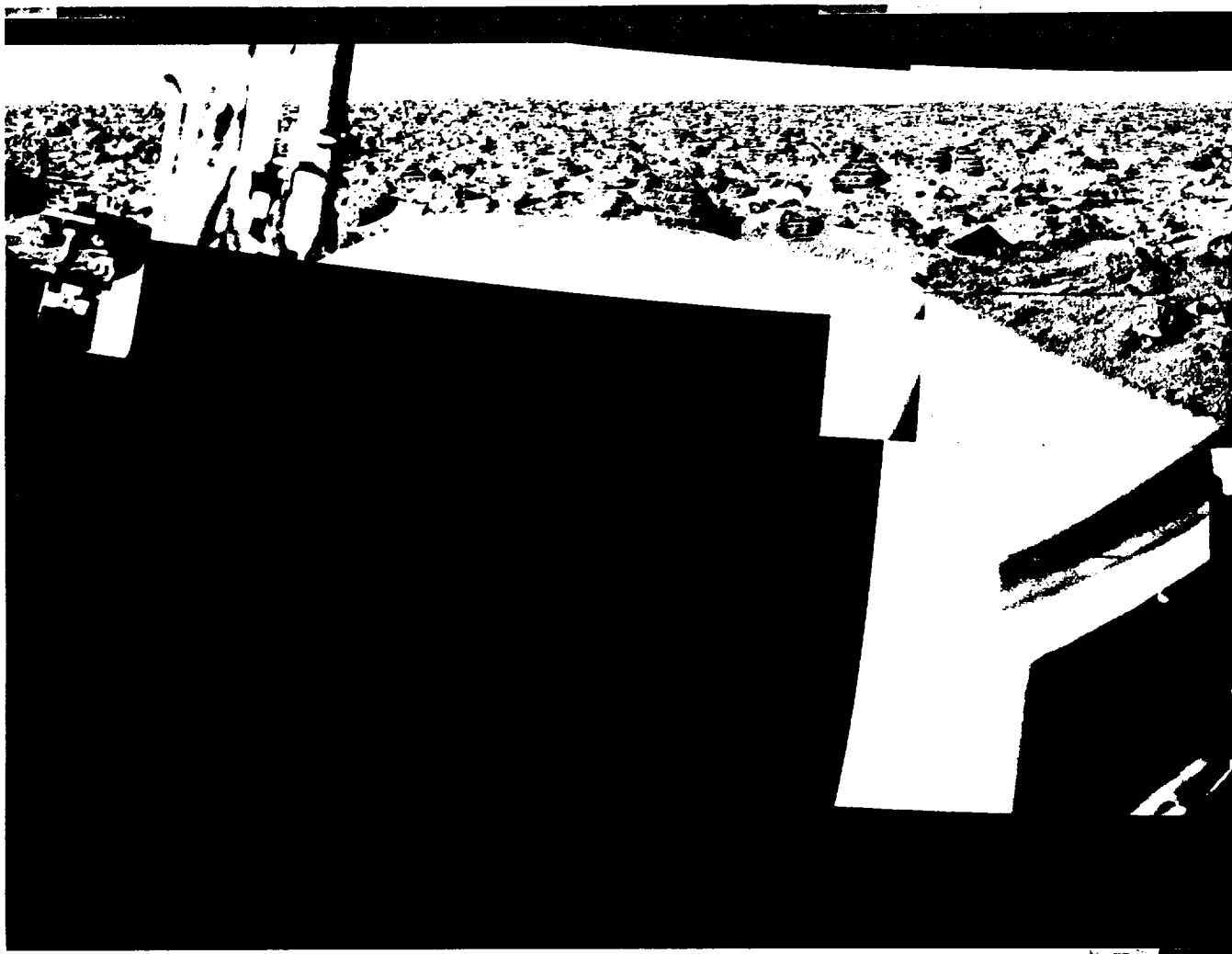


Fig. 49. Elevation Contour Overlay Mosaic: Camera 1 (right), back right quadrant — from IPL PIC ID 79/04/04/030319.

6.3 Elevation Contour Map Collection

6.3.1 Tabulation of Elevation Contour Map Sheets

The purpose of the following tabulation is to facilitate paging to desired map sheets, as well as to indicate which sheets have been produced. A separate tabular array is presented for

each scale. The rows and columns of the matrices are labeled with the parameters of the sheet array designators (see section 3.3) of the individual map sheets. The column labeling indicates the value of the first parameter of the designator, and the row labeling indicates the second parameter of the designator. The array entries associated with the designators indicate the page number upon which the sheet may be found.

VI.2 Contour Sheets								Scale 1:2000							
	-6	-4	-2	+0	+2	+4	+6								
+6															
+4															
+2															
+0				161											
-2															
-4															
-6															

VI.2 Contour Sheets								Scale 1:100							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5															
+3															
+1				168	169										
-1				170	171										
-3															
-5															
-7															

VI.2 Contour Sheets								Scale 1:5							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5															
+3			203	204	205	206									
+1			207	208	209	210									
-1															
-3			211												
-5															
-7															

VI.2 Contour Sheets								Scale 1:1000							
	-6	-4	-2	+0	+2	+4	+6								
+6															
+4															
+2															
+0				162											
-2															
-4															
-6															

VI.2 Contour Sheets								Scale 1:50							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5															
+3															
+1				172	173										
-1				174	175										
-3															
-5															
-7															

VI.2 Contour Sheets								Scale 1:2							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5			212	213	214	215									
+3		216	217	218	219	220	221								
+1															
-1															
-3															
-5															
-7															

VI.2 Contour Sheets								Scale 1:500							
	-6	-4	-2	+0	+2	+4	+6								
+6															
+4															
+2															
+0				163											
-2															
-4															
-6															

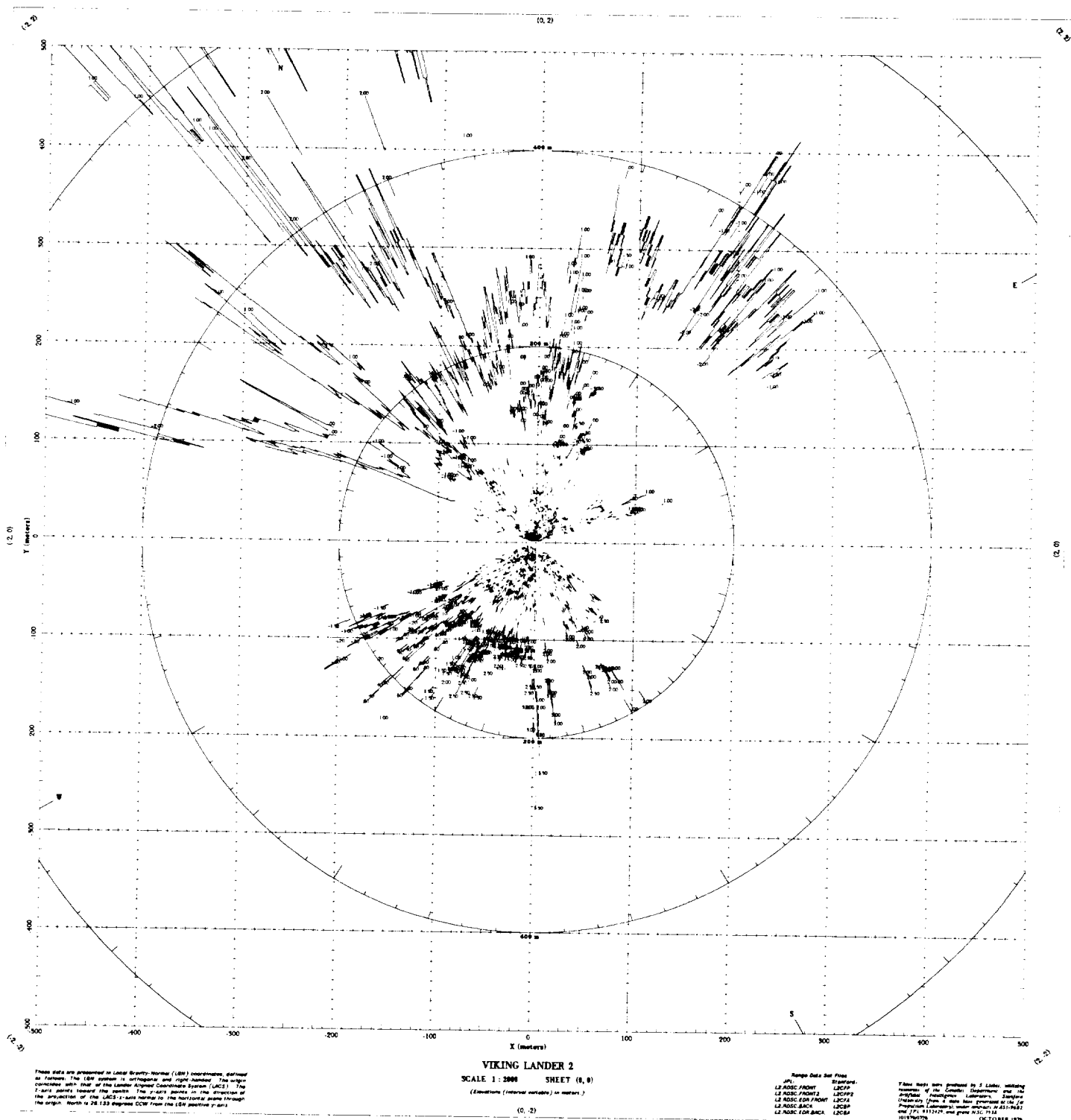
VI.2 Contour Sheets								Scale 1:20							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5															
+3			176	177	178	179									
+1			180	181	182	183									
-1			184	185											
-3			186	187	188	189									
-5															
-7															

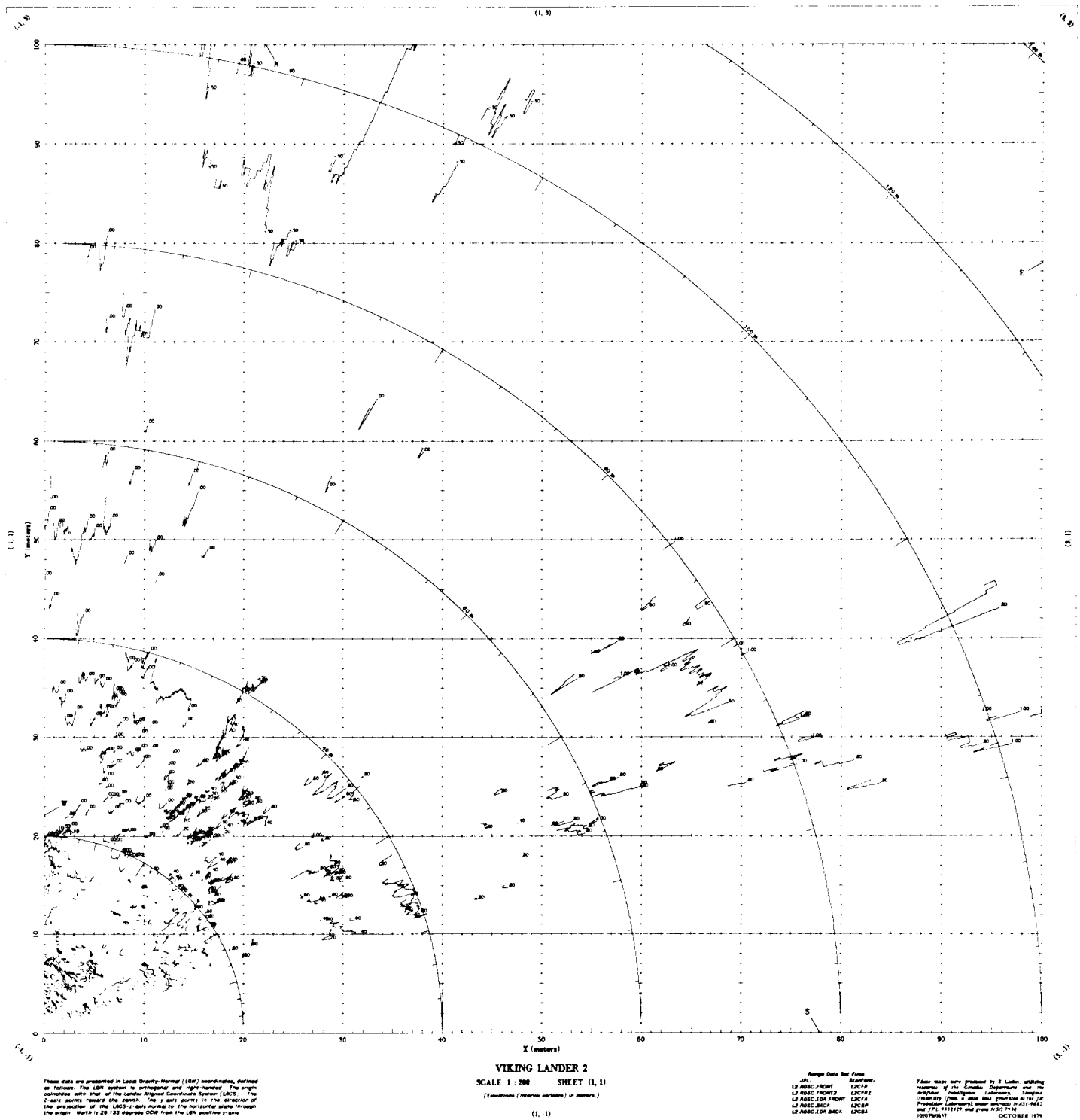
VI.2 Contour Sheets								Scale 1:1							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7	222	223	224			225									
+5														226	
+3															
+1															
-1															
-3															
-5															
-7															

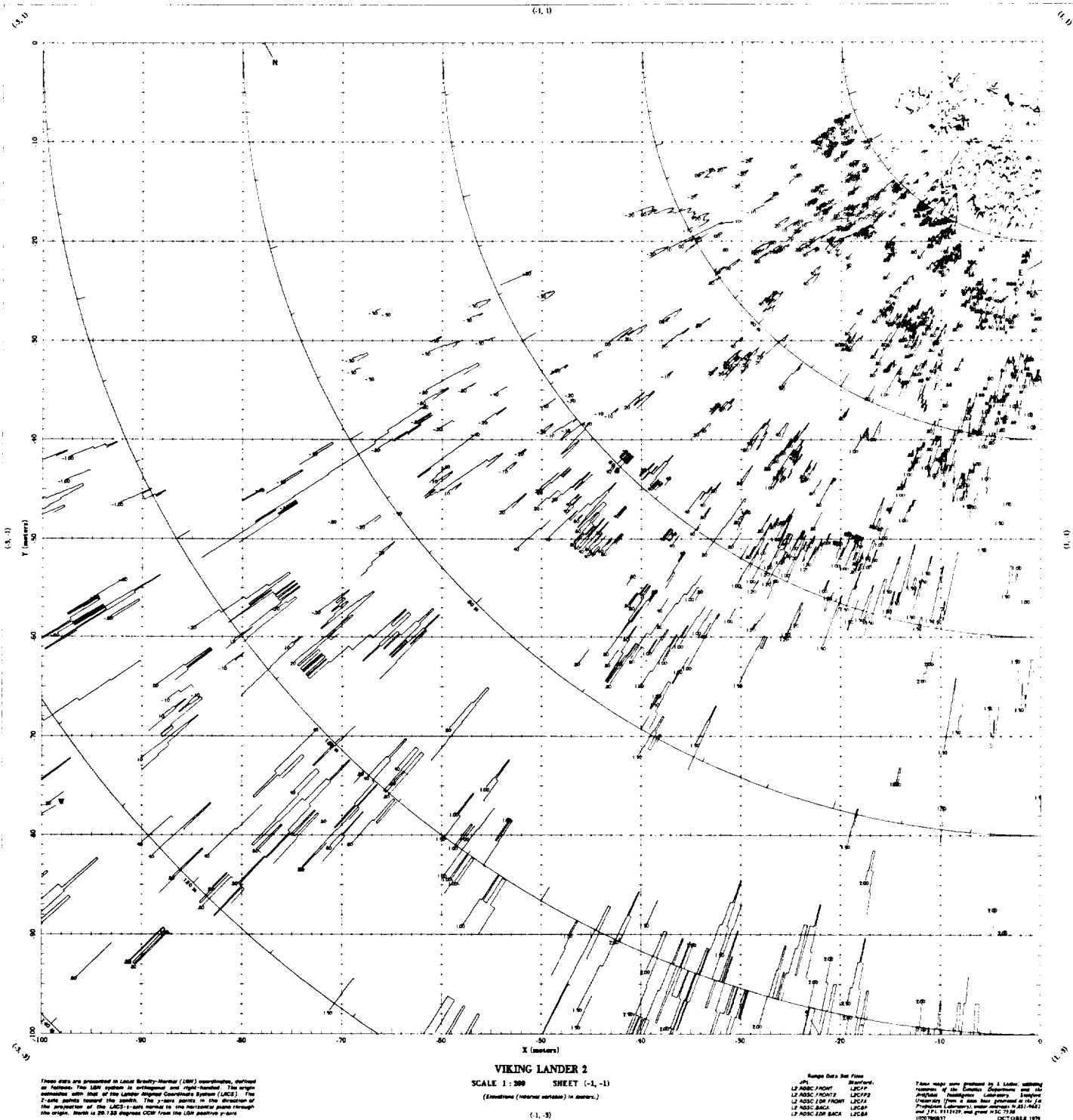
VI.2 Contour Sheets								Scale 1:200							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5															
+3															
+1				164	165										
-1				166	167										
-3															
-5															
-7															

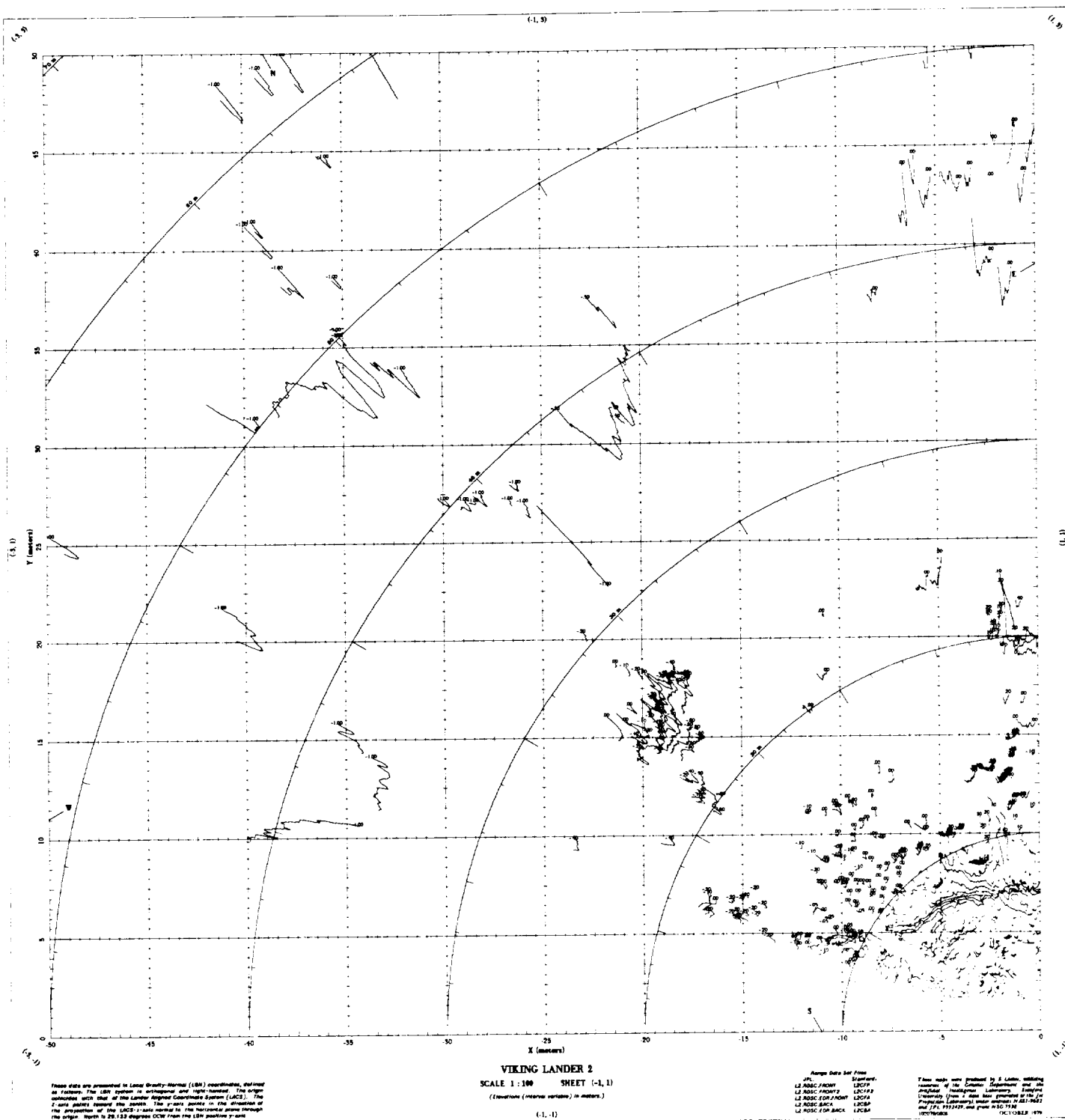
VI.2 Contour Sheets								Scale 1:10							
	-7	-5	-3	-1	+1	+3	+5	+7							
+7															
+5															
+3			190	191	192	193									
+1			194	195	196	197									
-1			198	199											
-3			200	201		202									
-5															
-7															

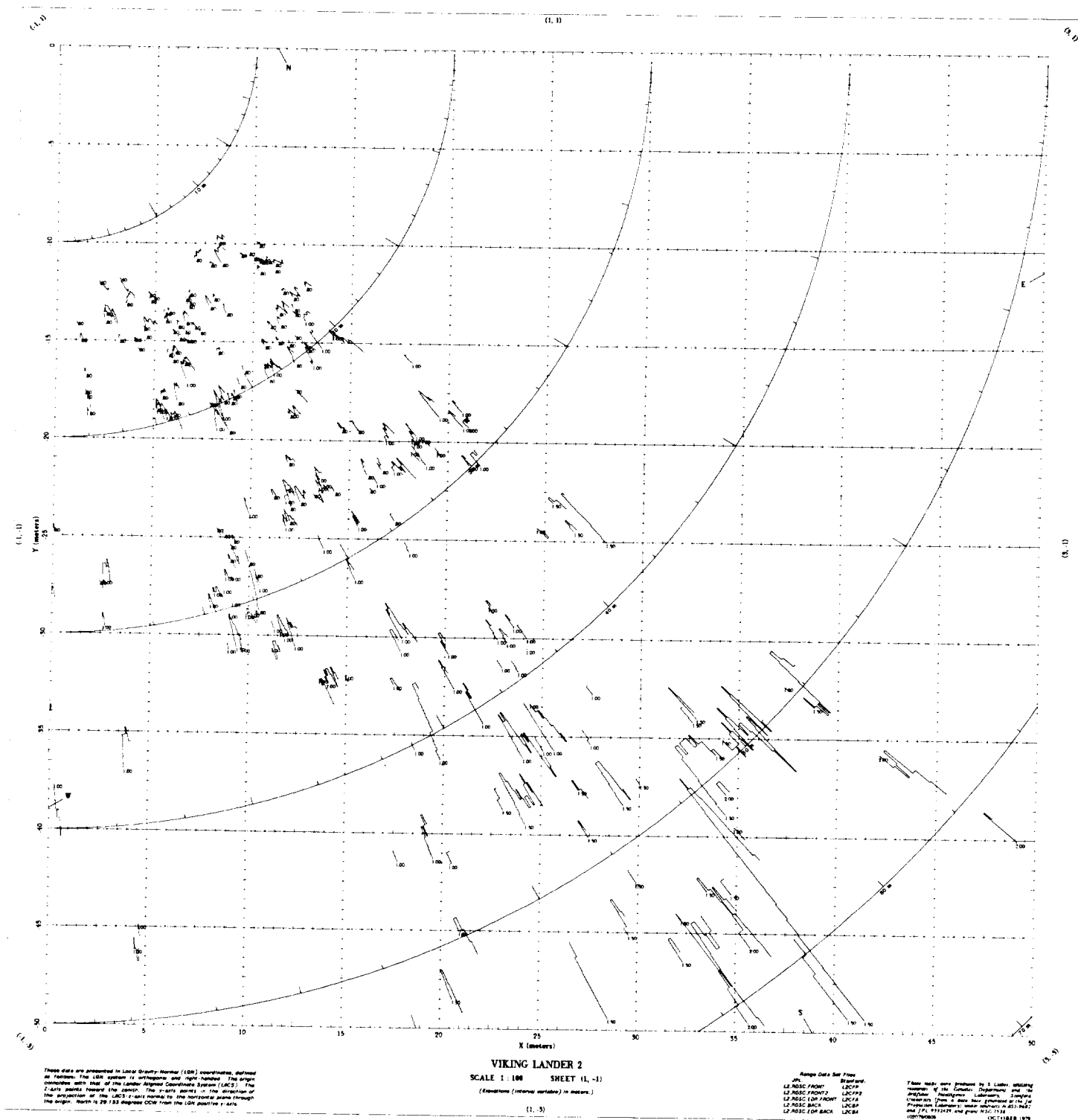
Table 2. Tabulation of Elevation Contour Map Sheets.

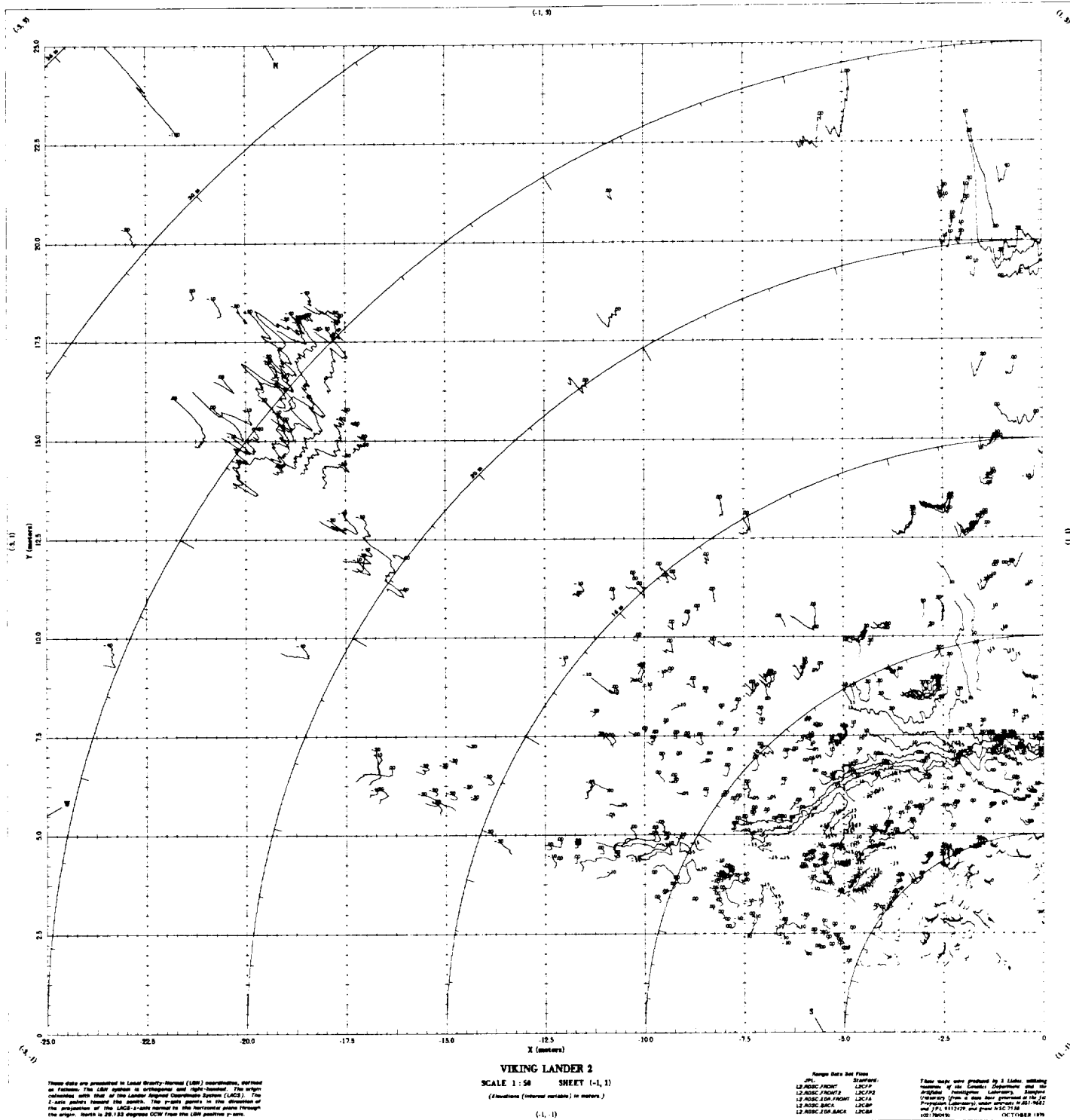


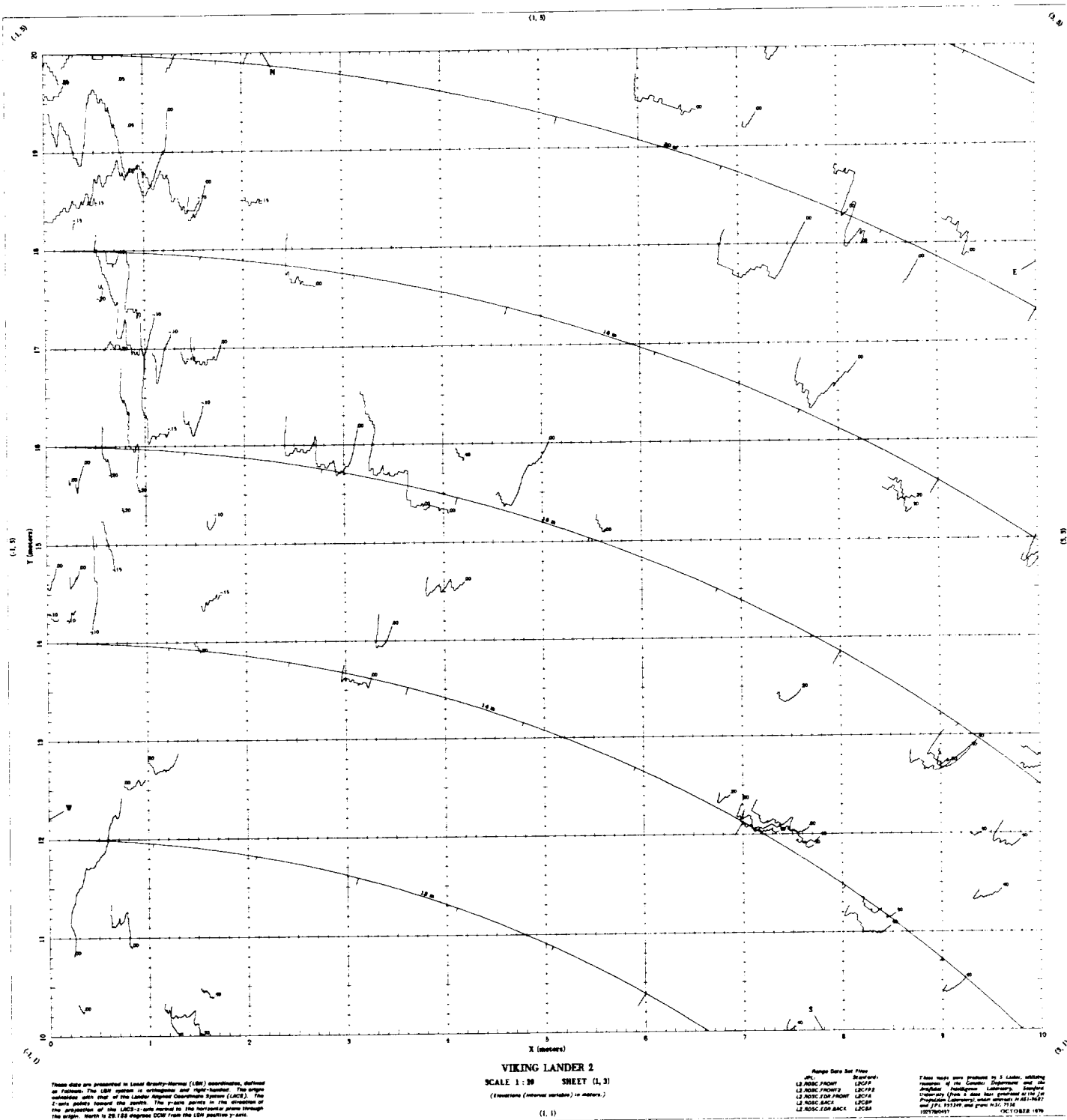


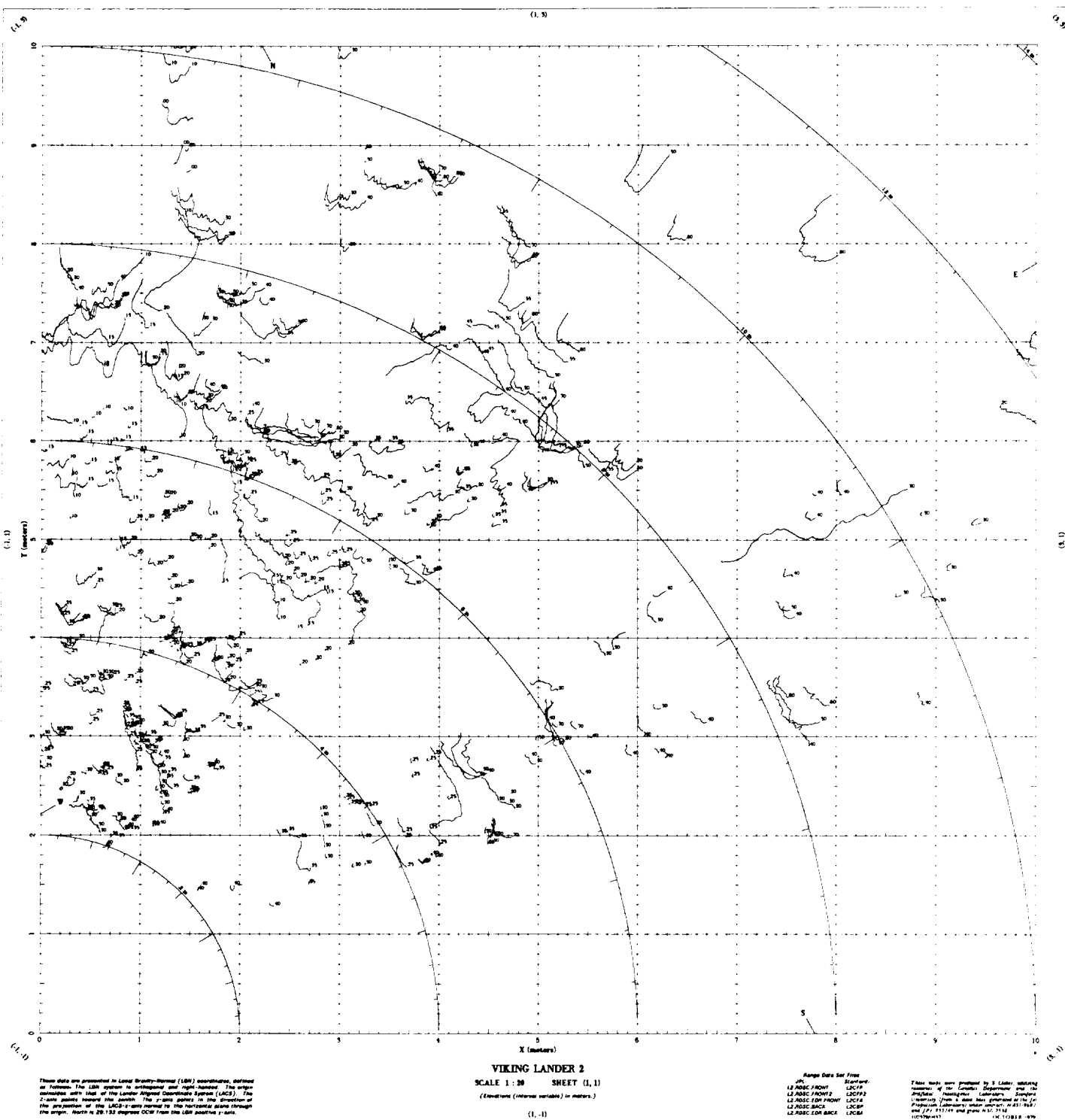










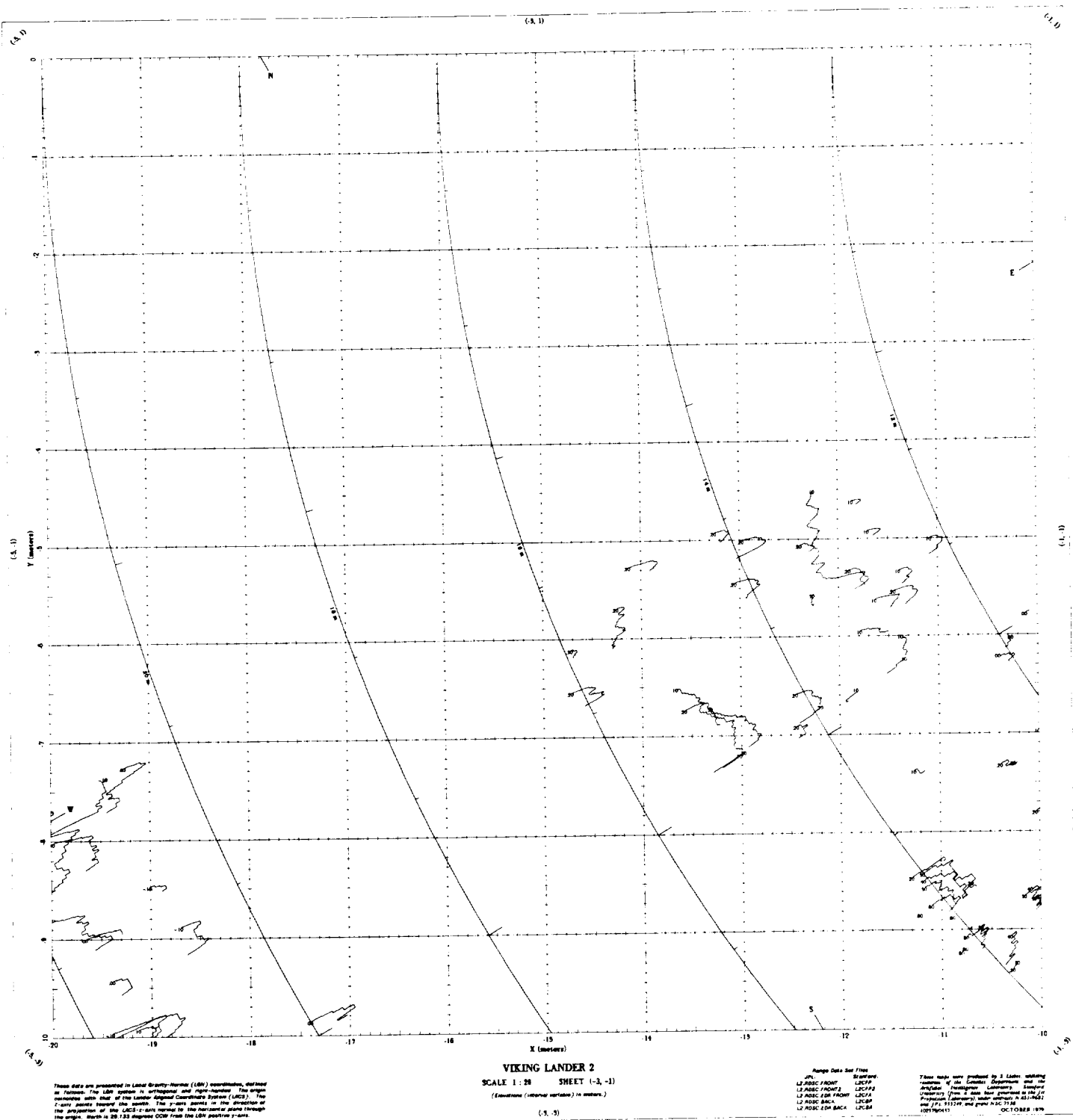


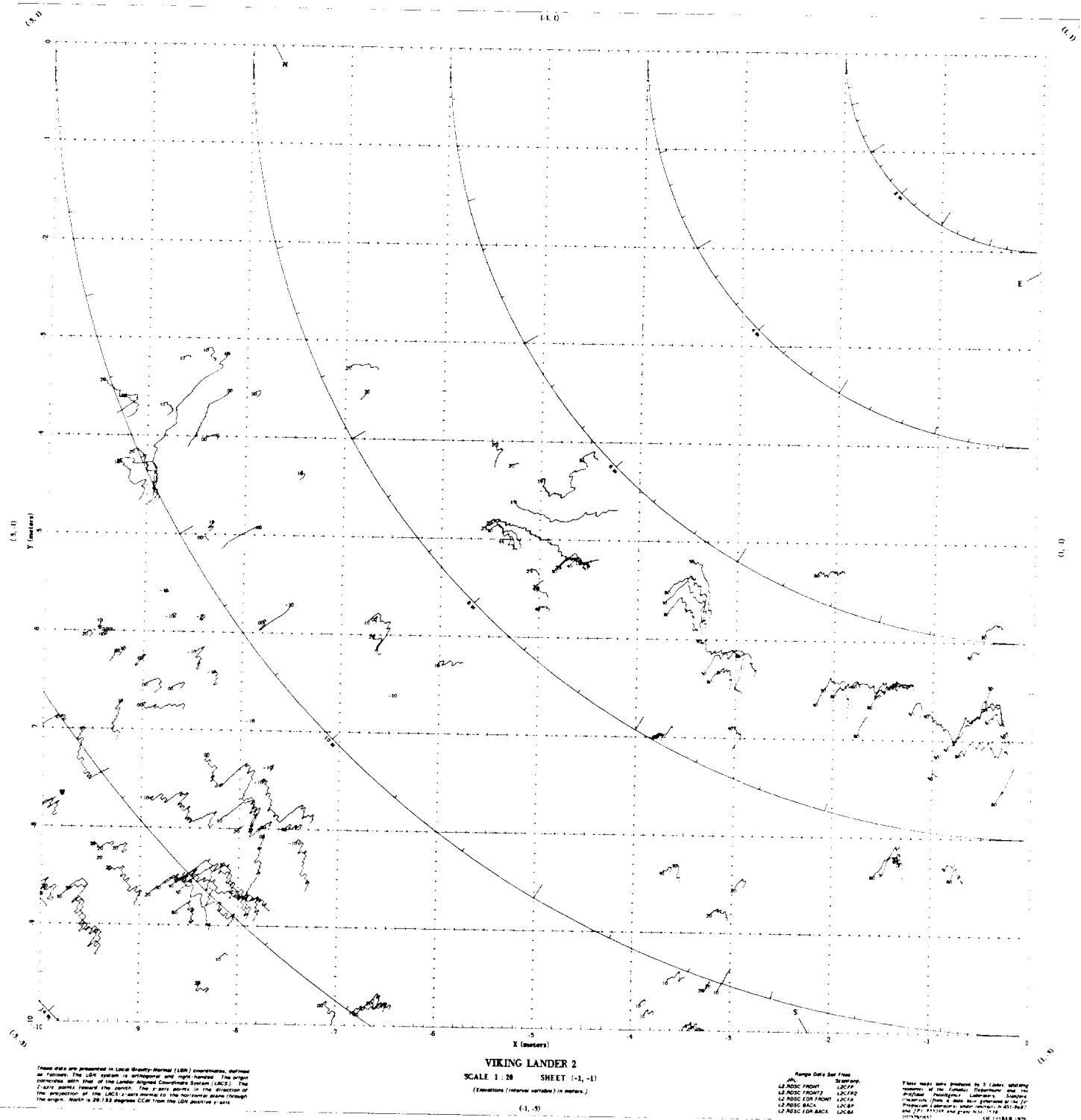
VIKING LANDER 2
SCALE 1:20 SHEET (1,1)

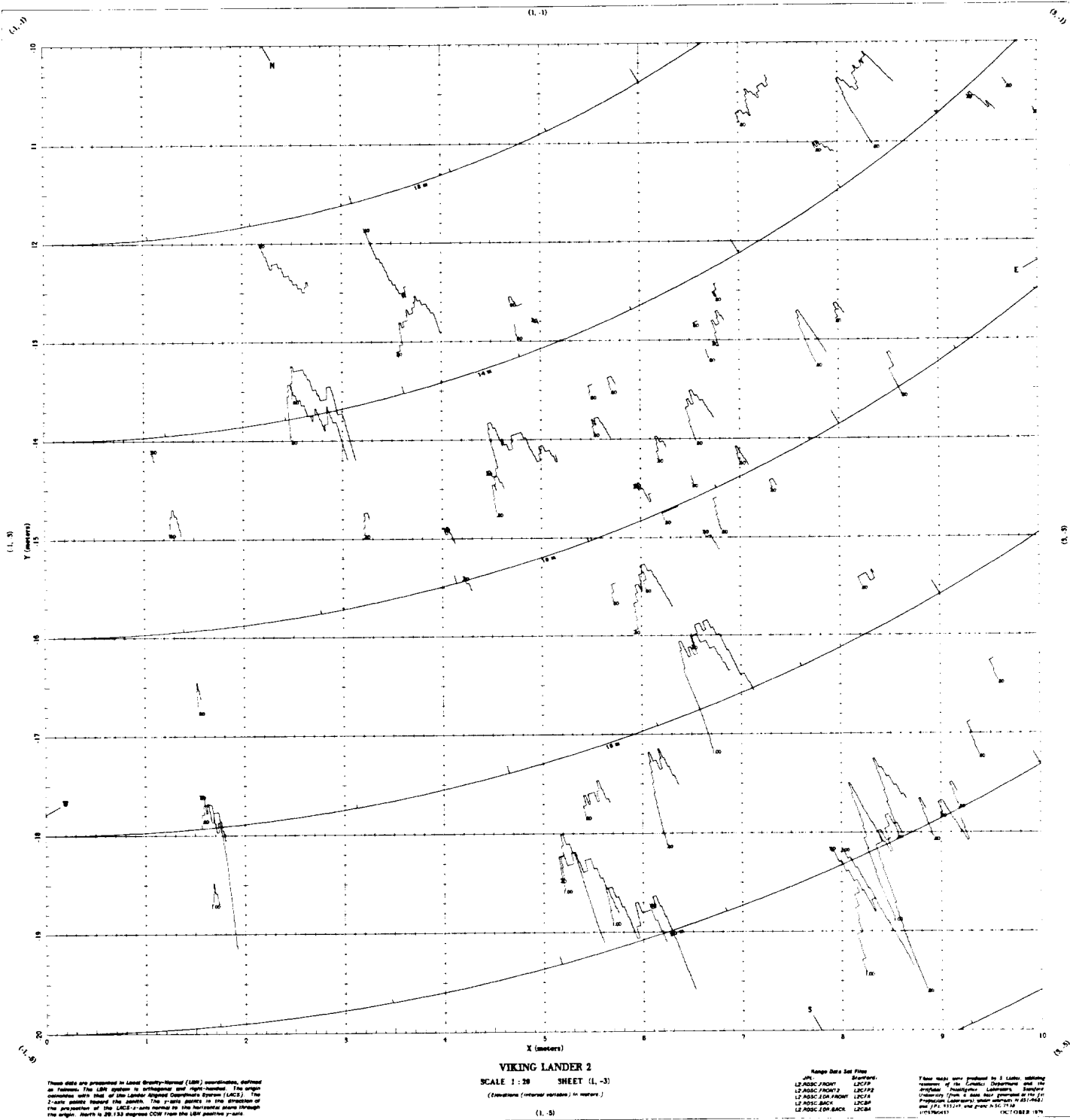
Range Data Set Files	
JPL	Starts
L2 ADSC.FRONT	L2CFA
L2 ADSC.FRONT2	L2CFF
L2 ADSC.EDR.FRONT	L2CFA
L2 ADSC.BACK	L2CBF
L2 ADSC.EDR.BACK	L2CBA

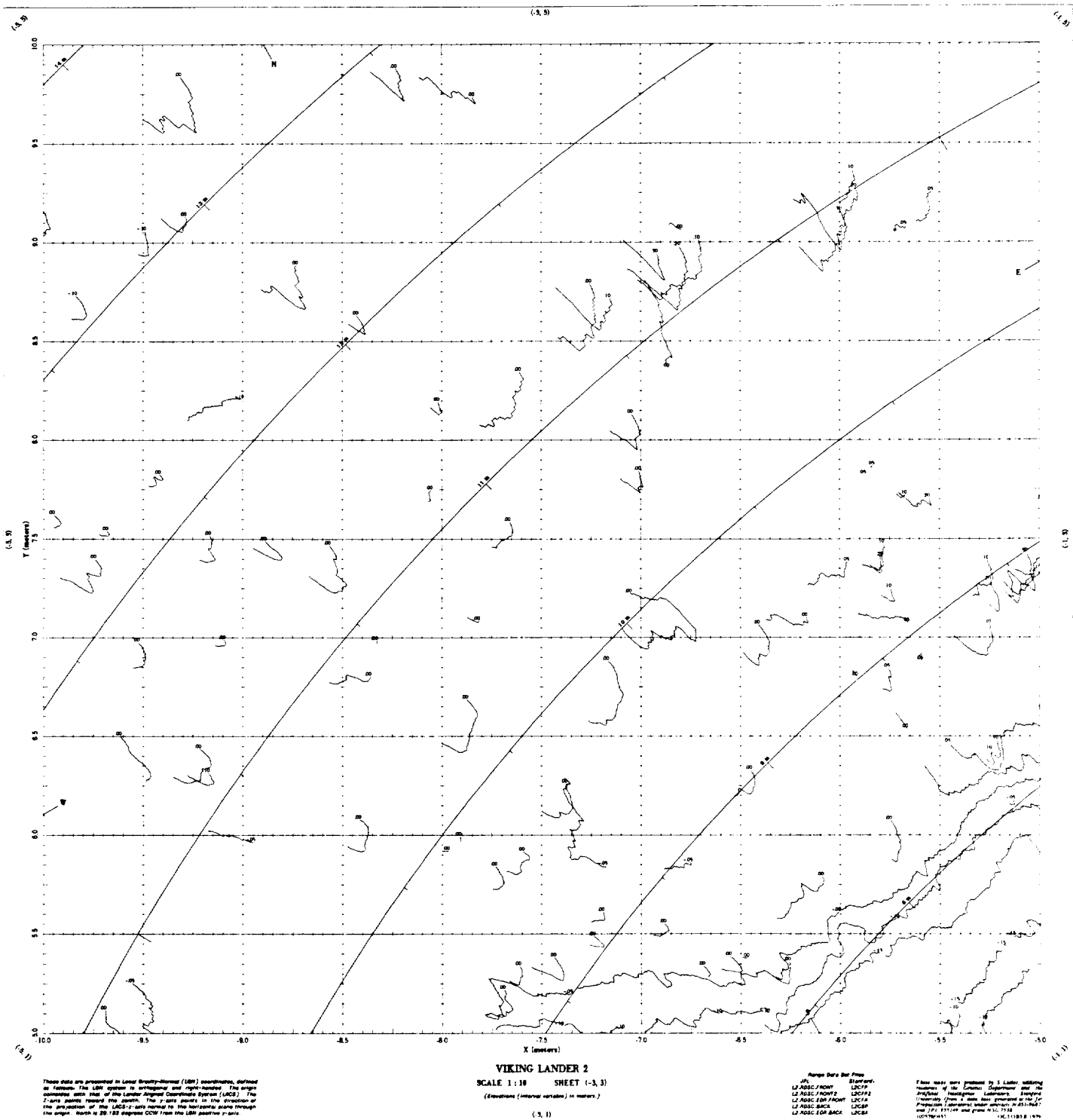
These maps were produced by S. Lister, student researcher of the Genetics Department and the Artificial Intelligence Laboratory, Stanford University from a data base generated at the Population Laboratory under contract N-851-96-2, and JPL 355-69 and grant NCS-75-52.

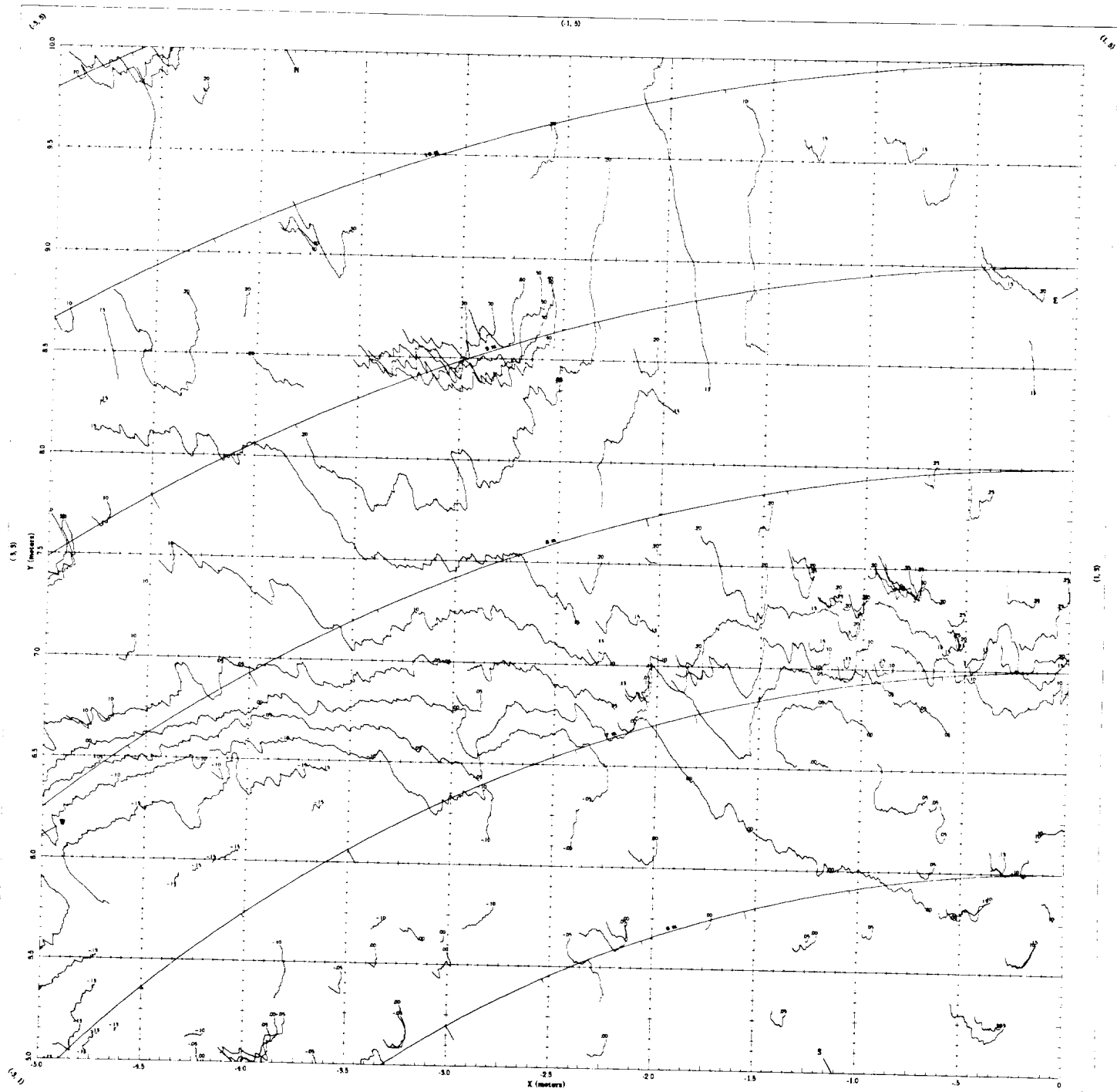
182









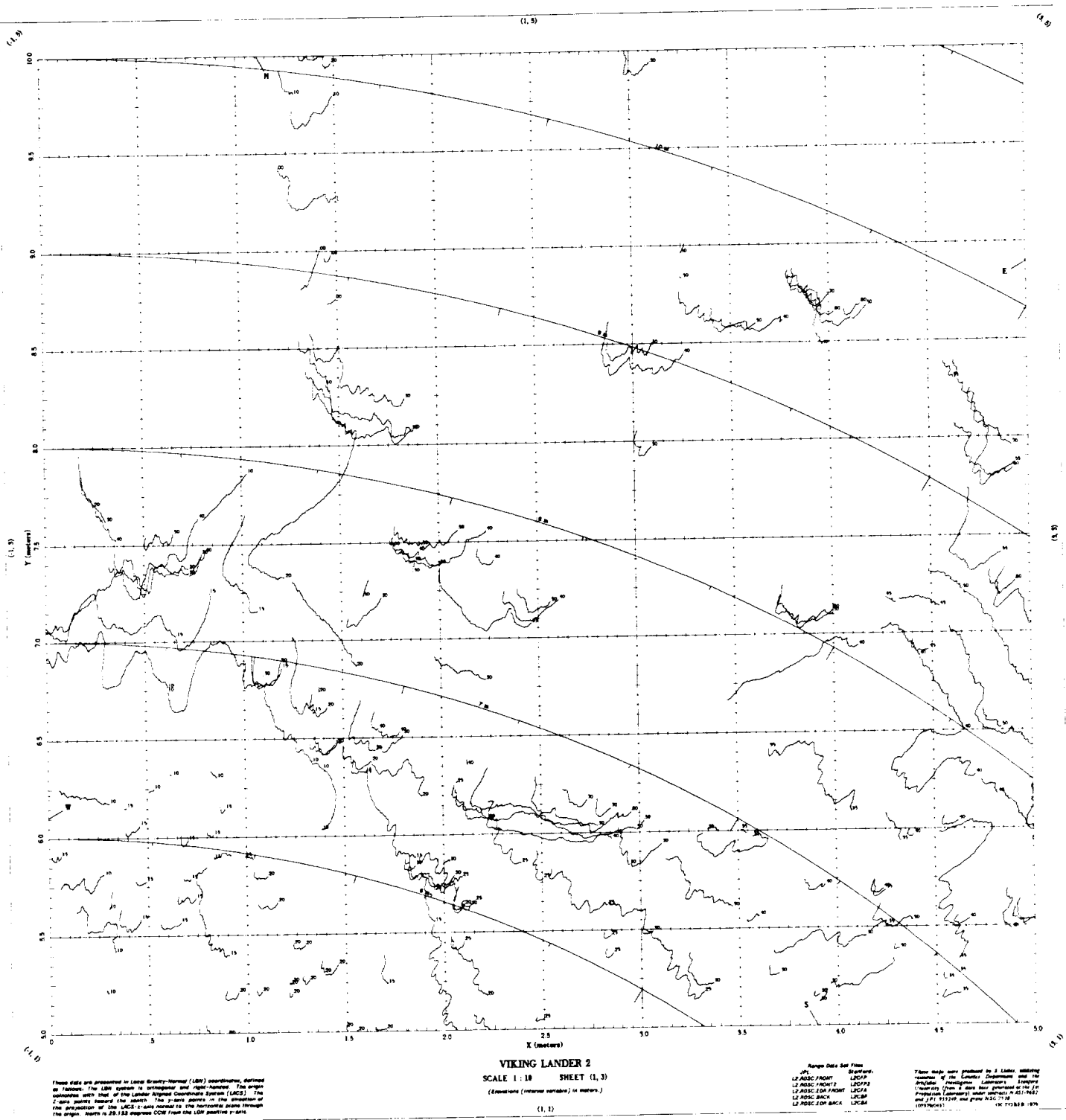


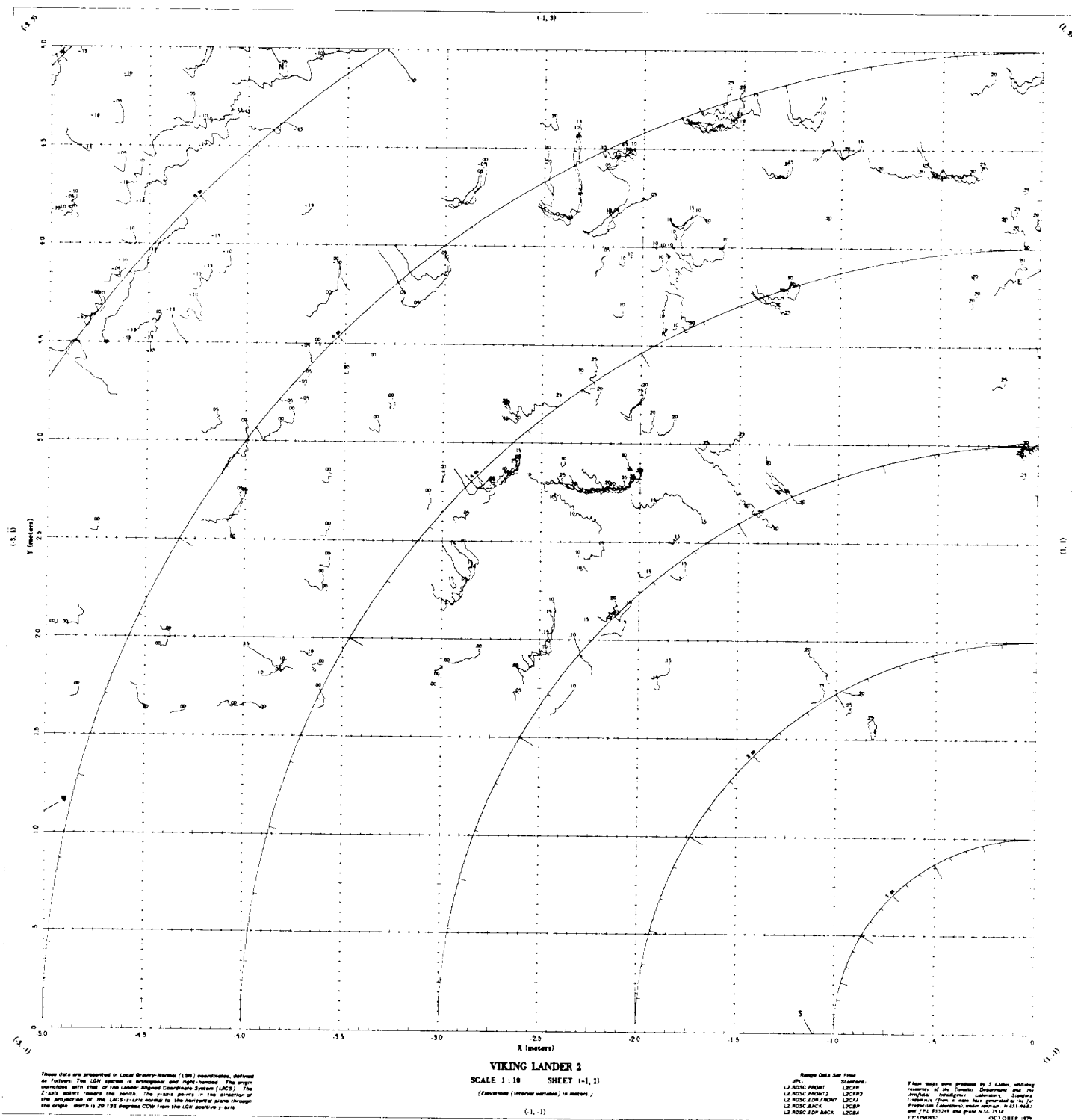
These data are presented in Local Gravity-Reference (LGR) coordinates, defined as follows: The LGR system is orthogonal and right-handed. The origin coincides with that of the Lander Reference Coordinate System (LRCS). The Y-axis points toward the south. The X-axis points in the direction of the projection of the LRCS Z-axis normal to the horizontal plane through the origin. North is 29.153 degrees CCW from the LGR positive Y-axis.

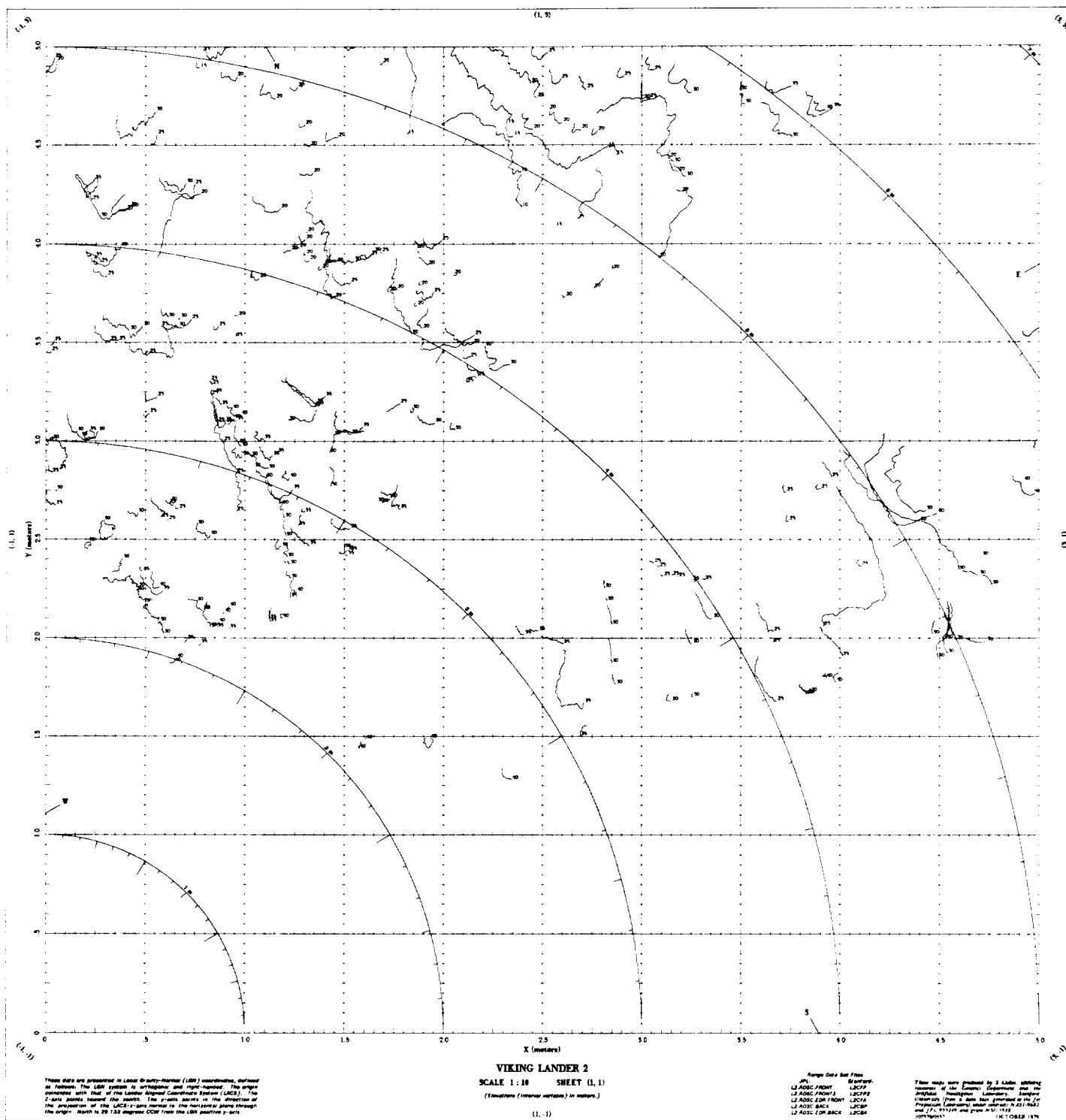
VIKING LANDER 2
SCALE 1:30 SHEET (-1,3)
(Elevations (vertical variable) in meters)
(-1, 1)

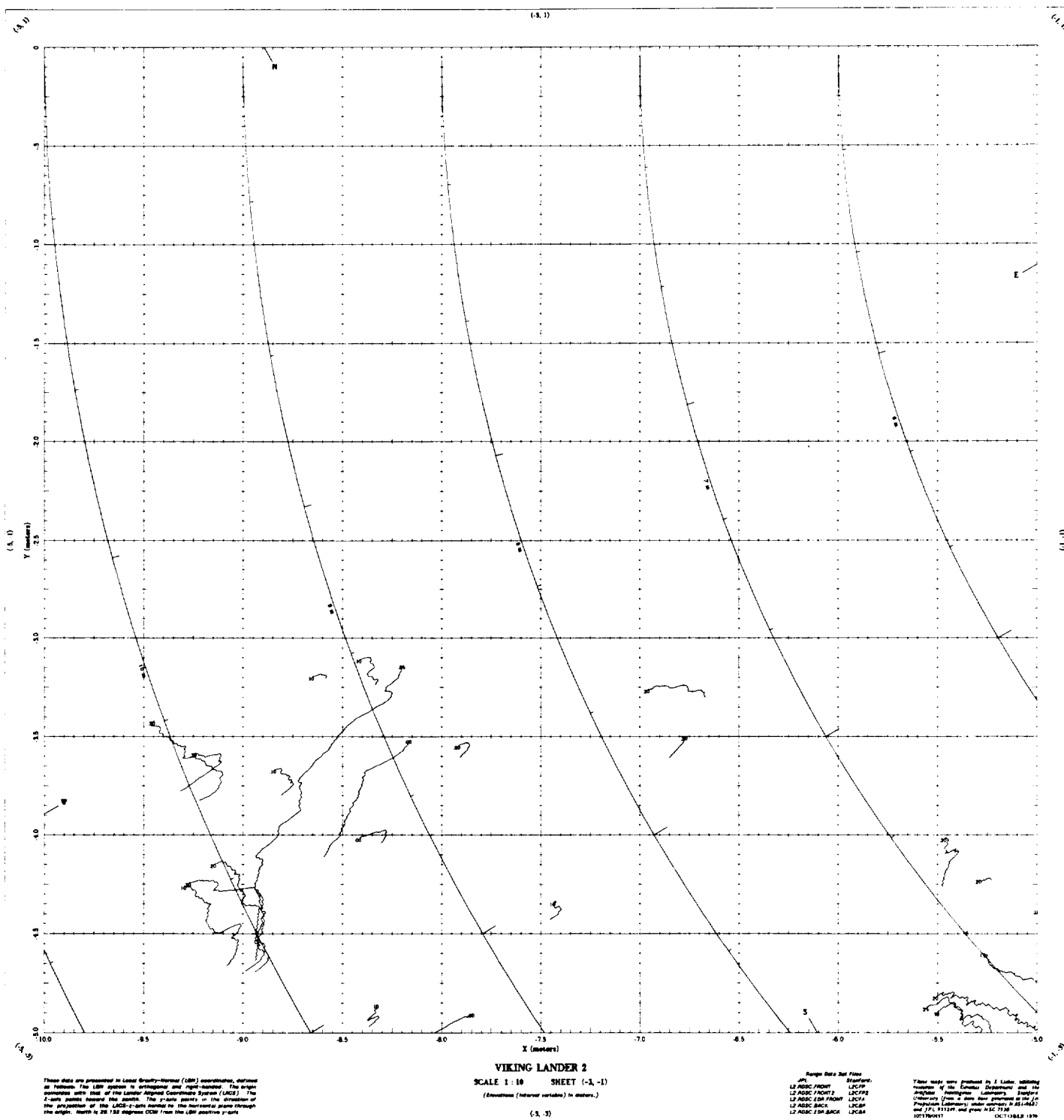
Range Data Set Files
JPL Stanford
L2.A00C.FRONT L2C00F
L2.A00C.FRONT2 L2C00F2
L2.A00C.FOR.FRONT L2C00F.F
L2.A00C.BACK L2C00B
L2.A00C.FOR.BACK L2C00B.F

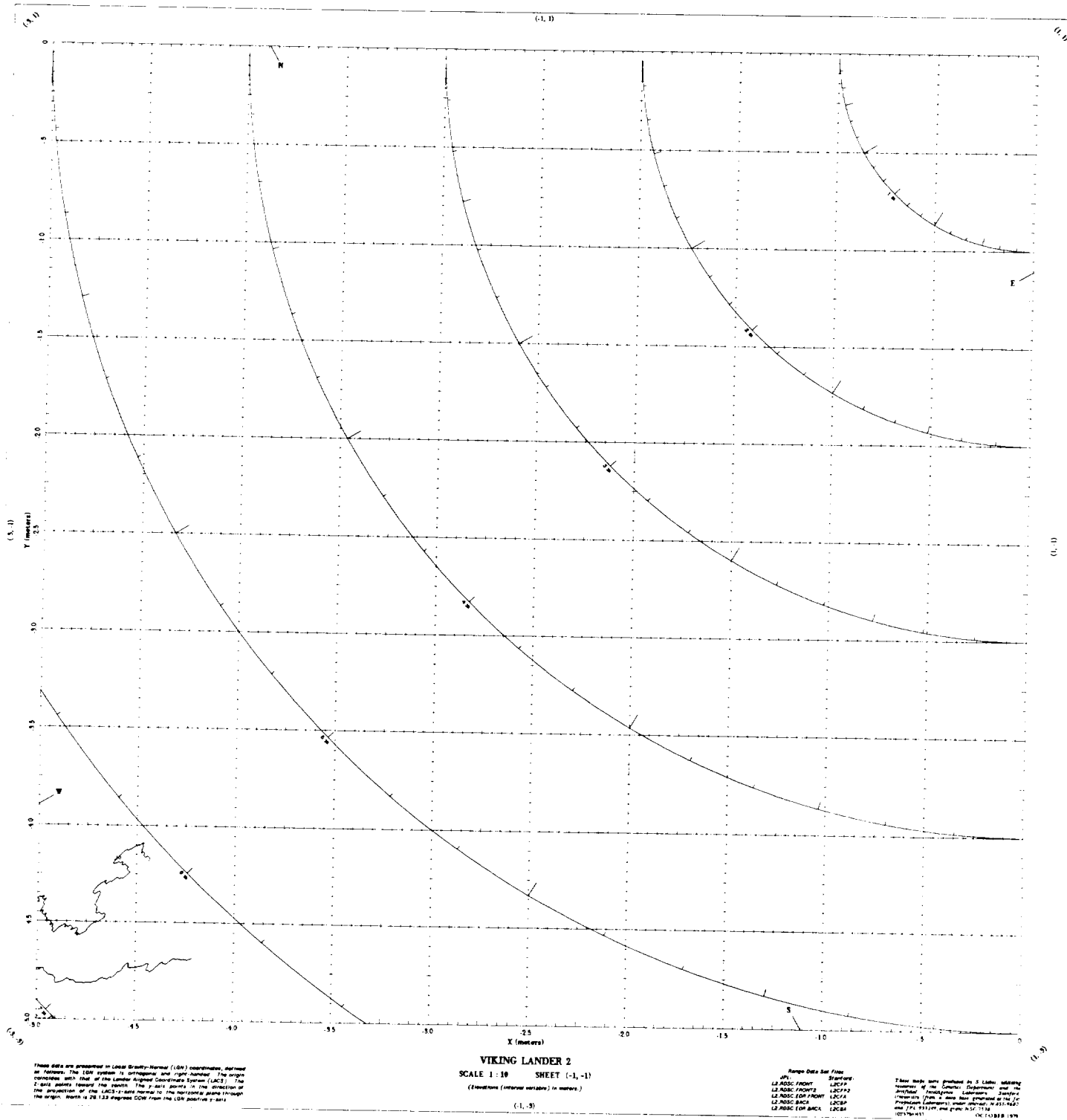
These data were generated by J. Linder, including
elevation, of the Lander, Camera, and LGR
coordinates. The data were generated at the JPL
Planetary Computer Center, under contract to JPL
and JPL, 915249 and given to SC 7114
10/1/79
OCTOBER 1979

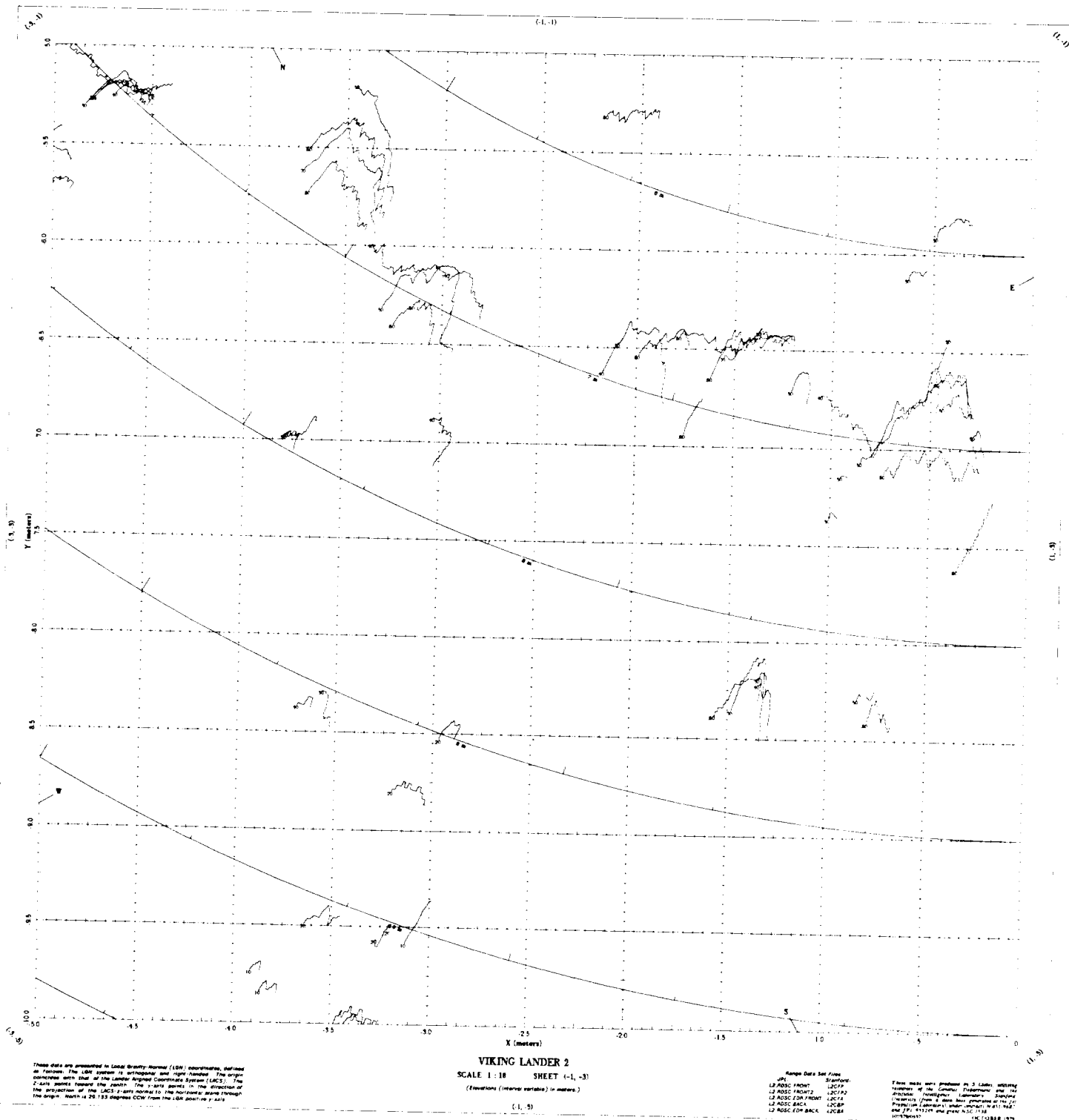


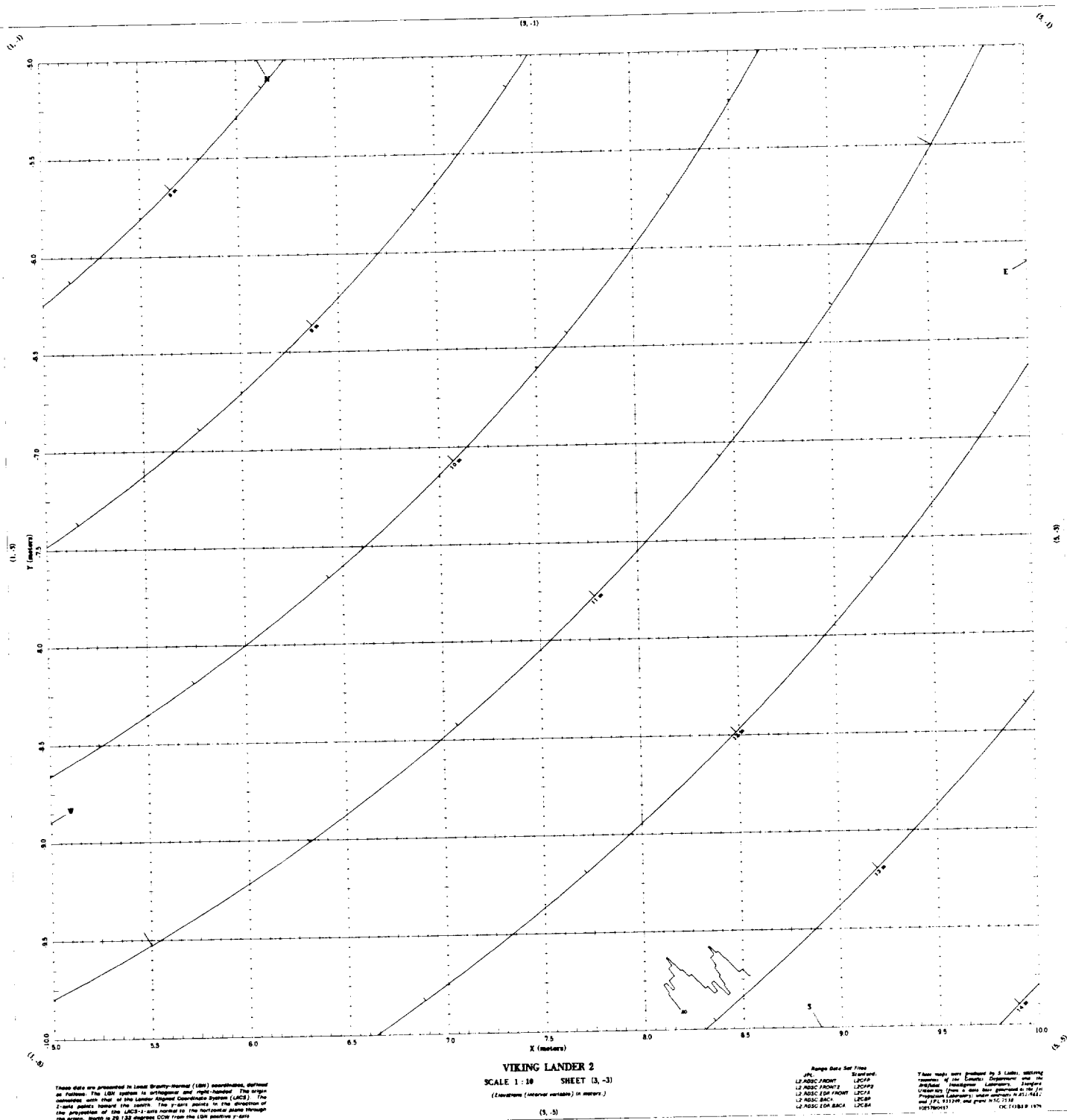


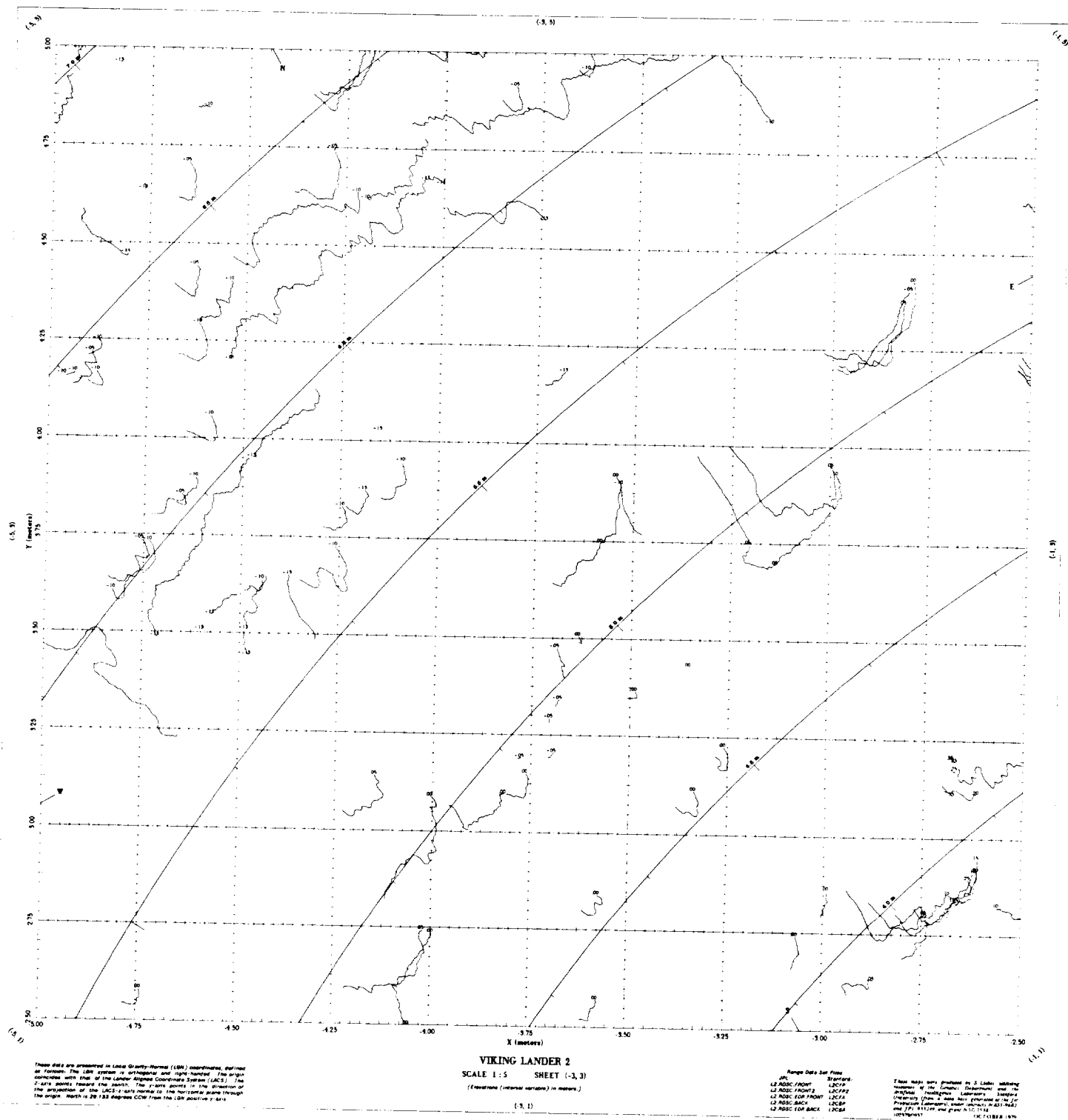


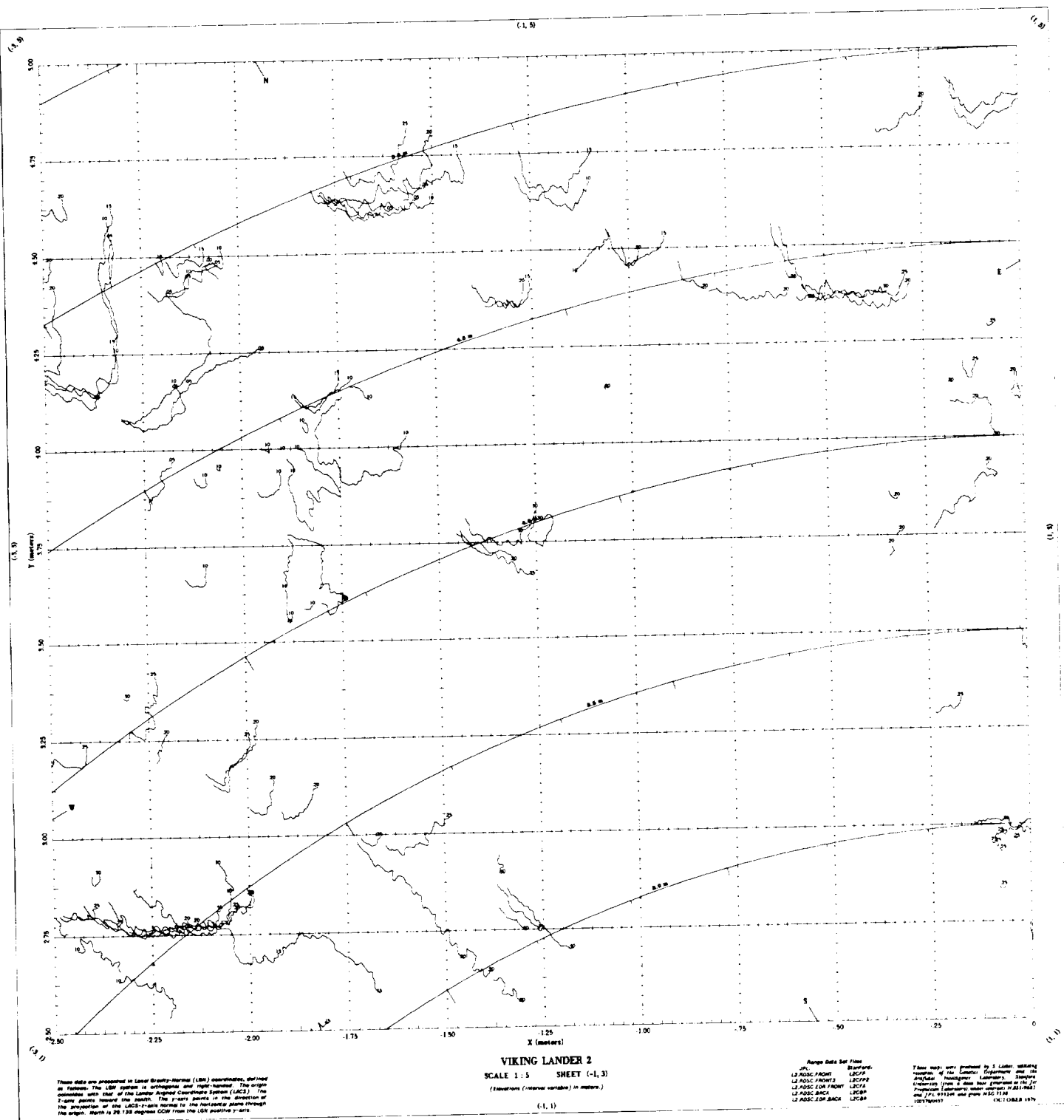


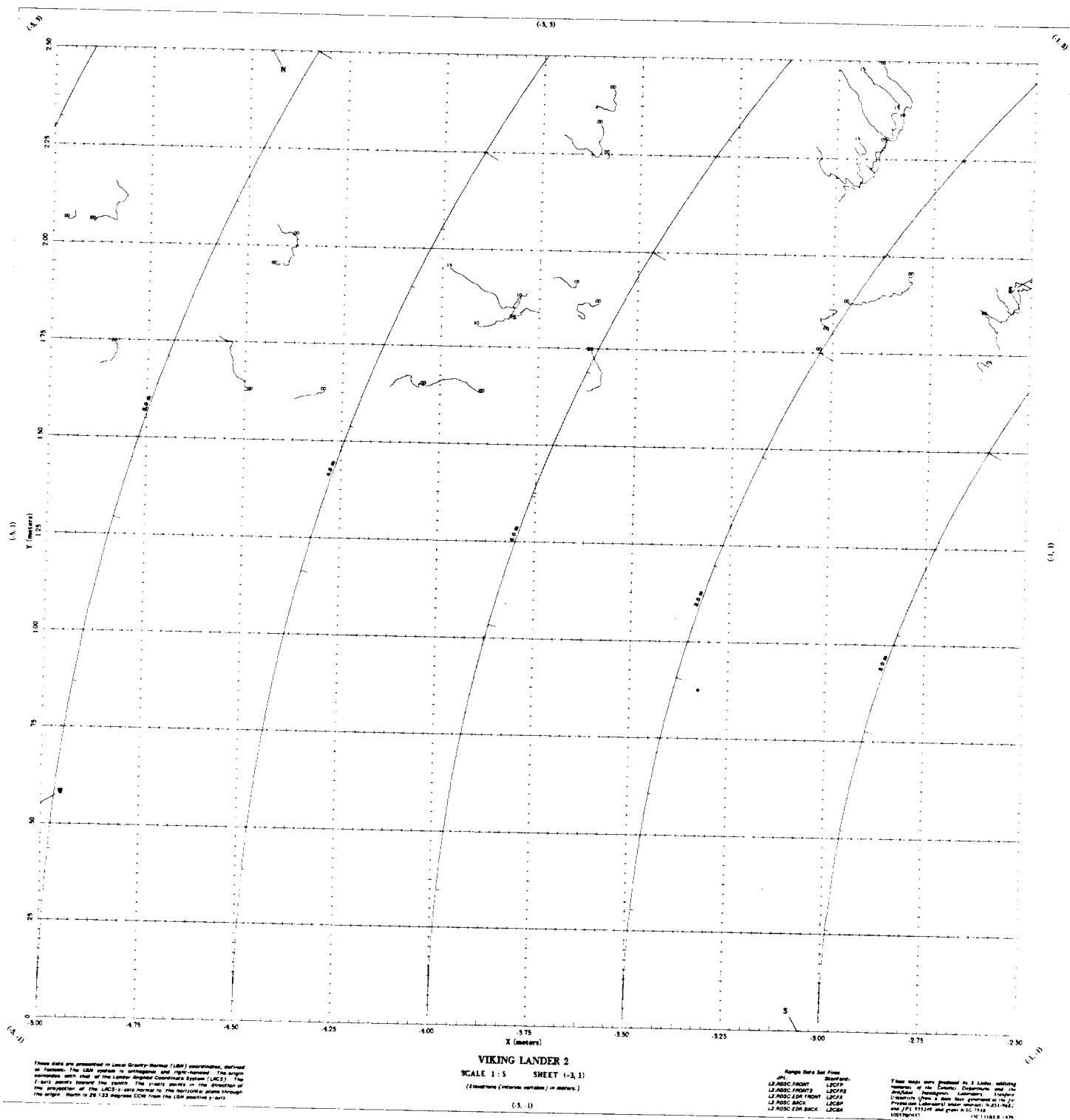


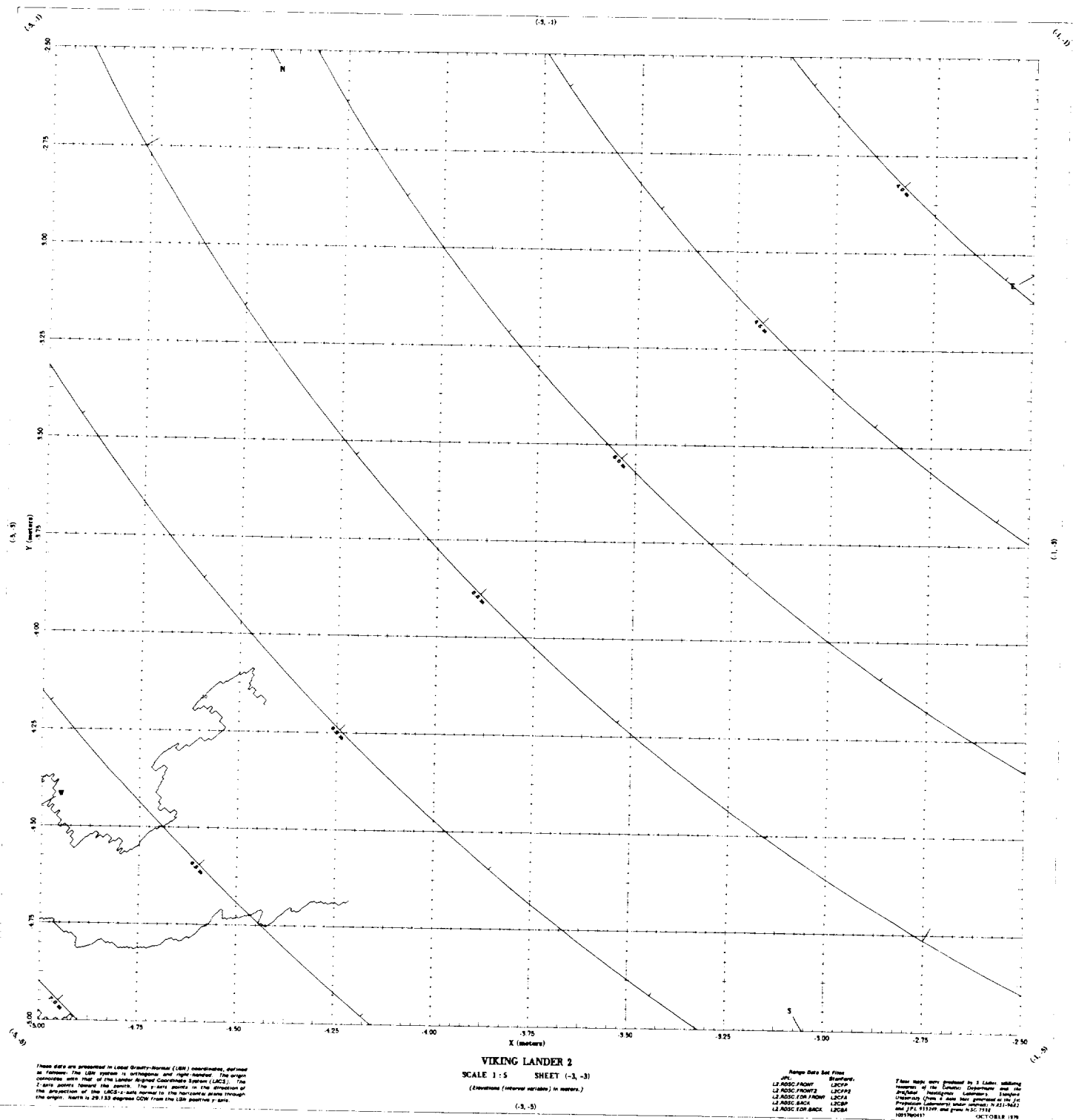


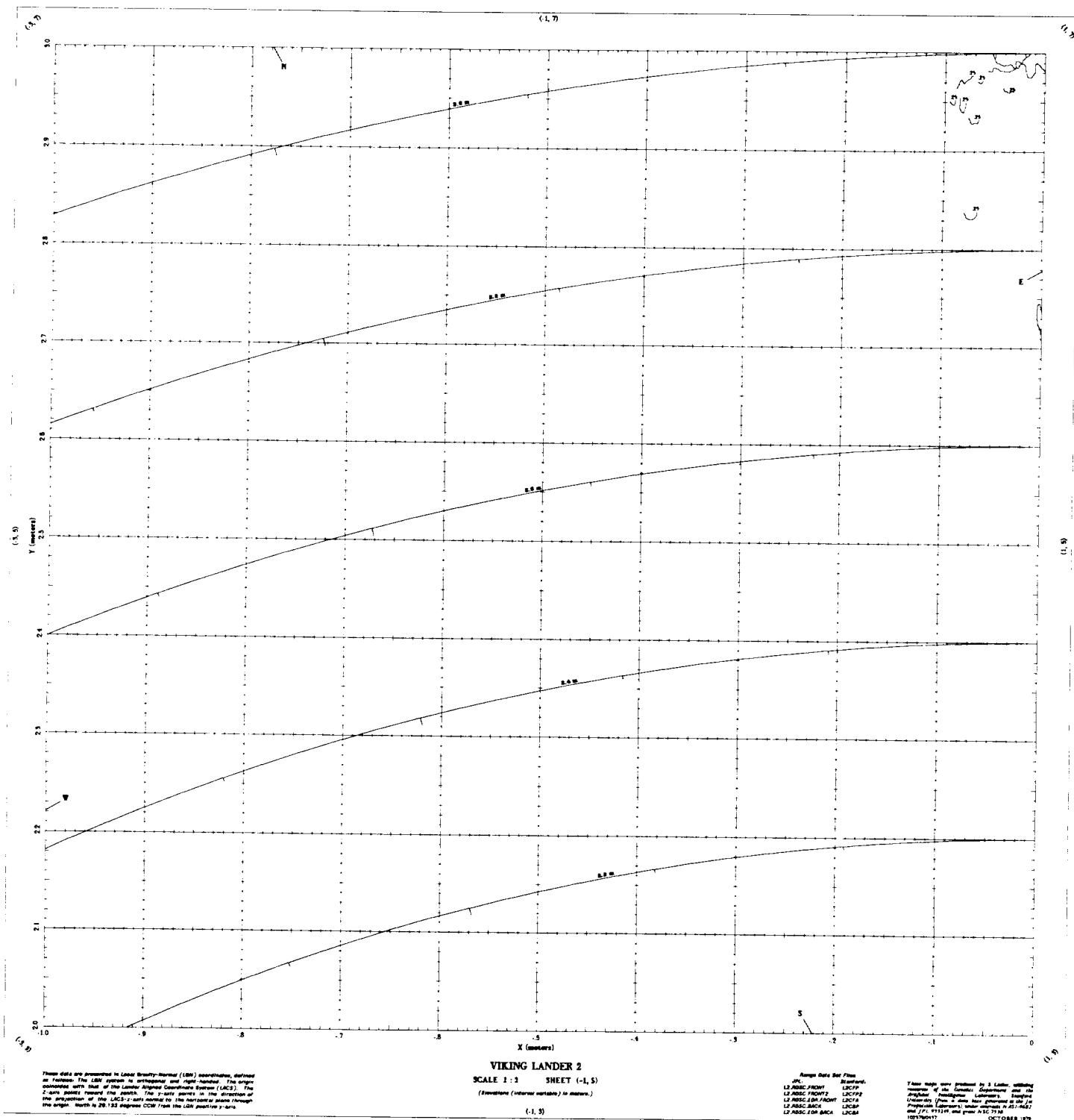


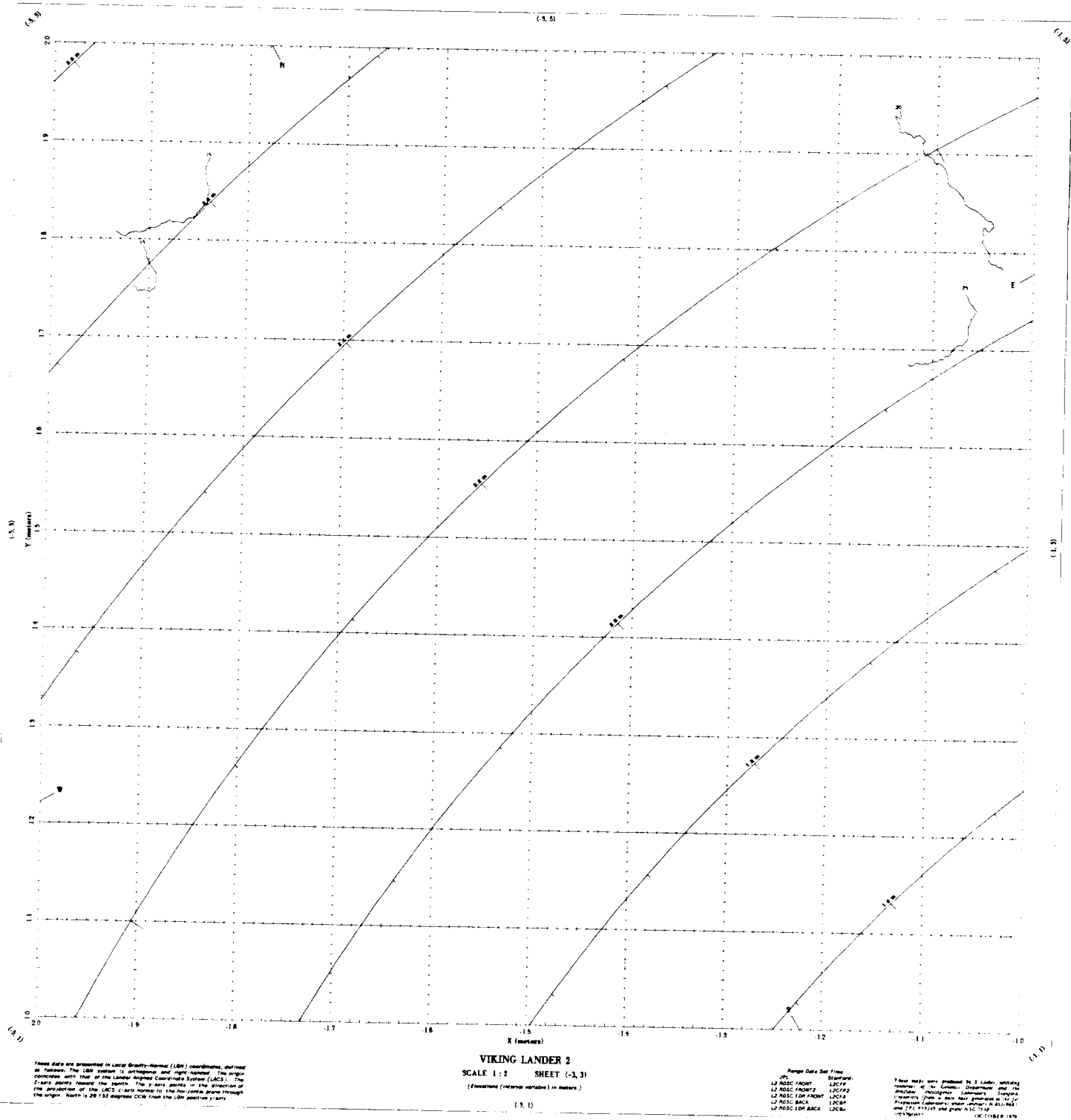


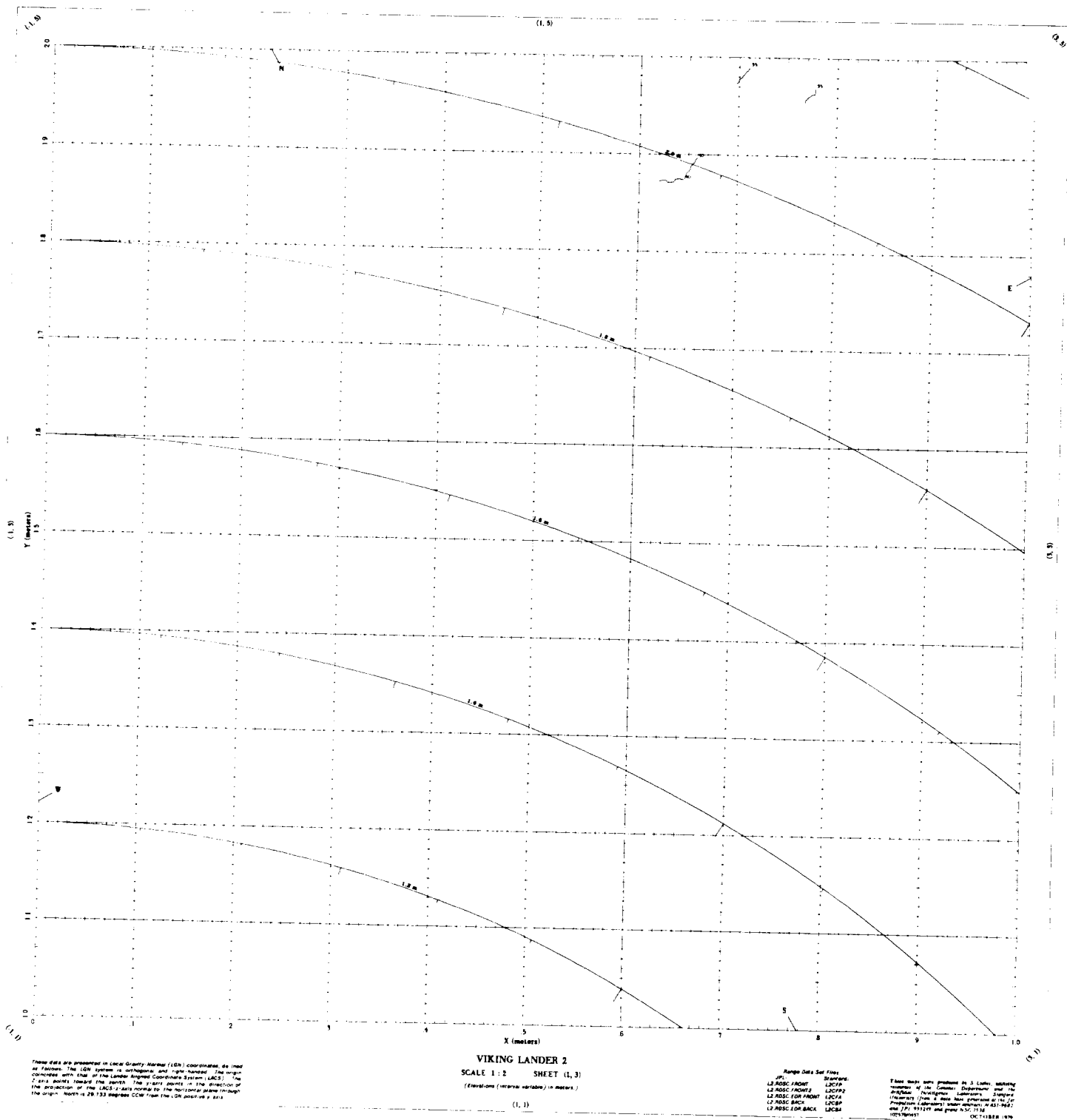


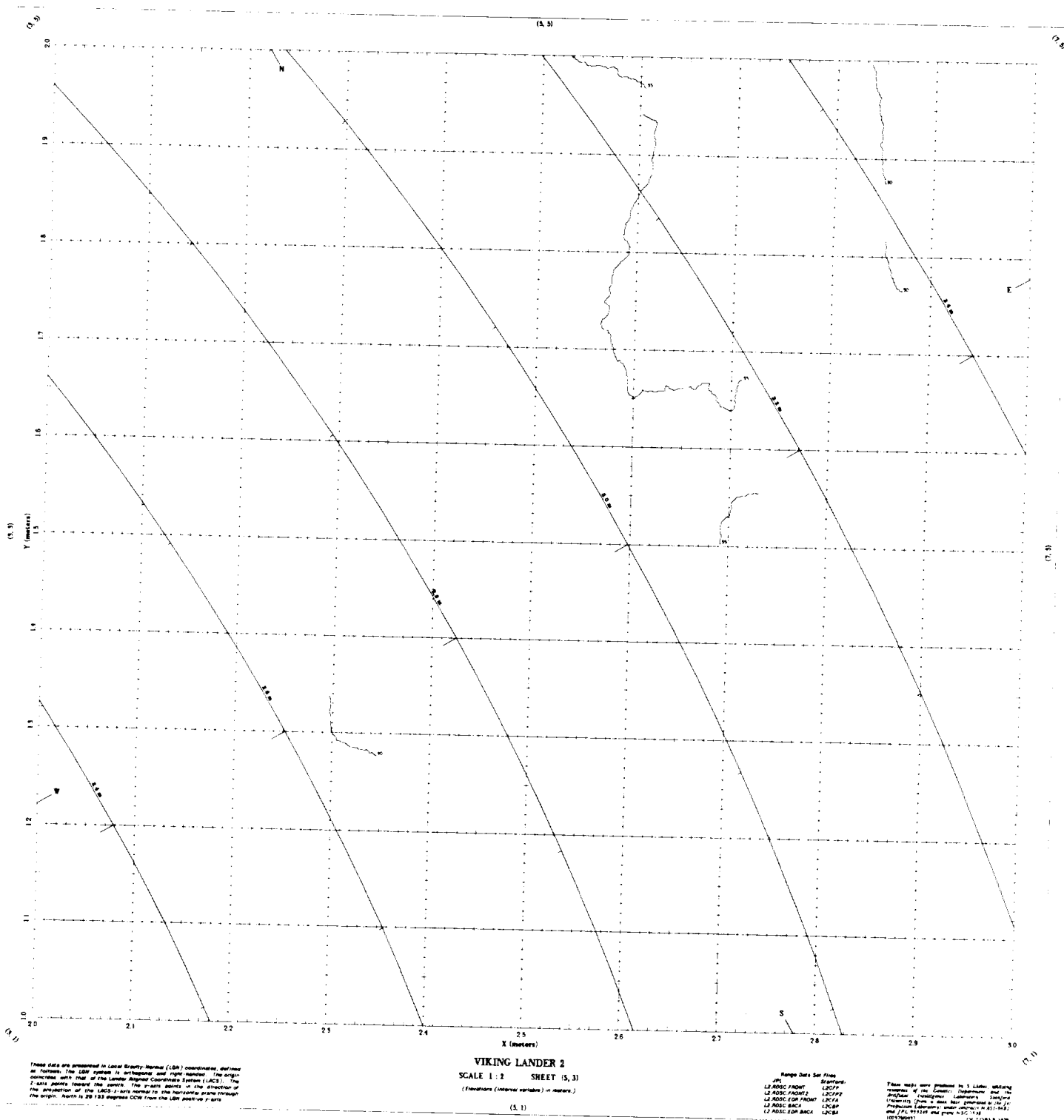


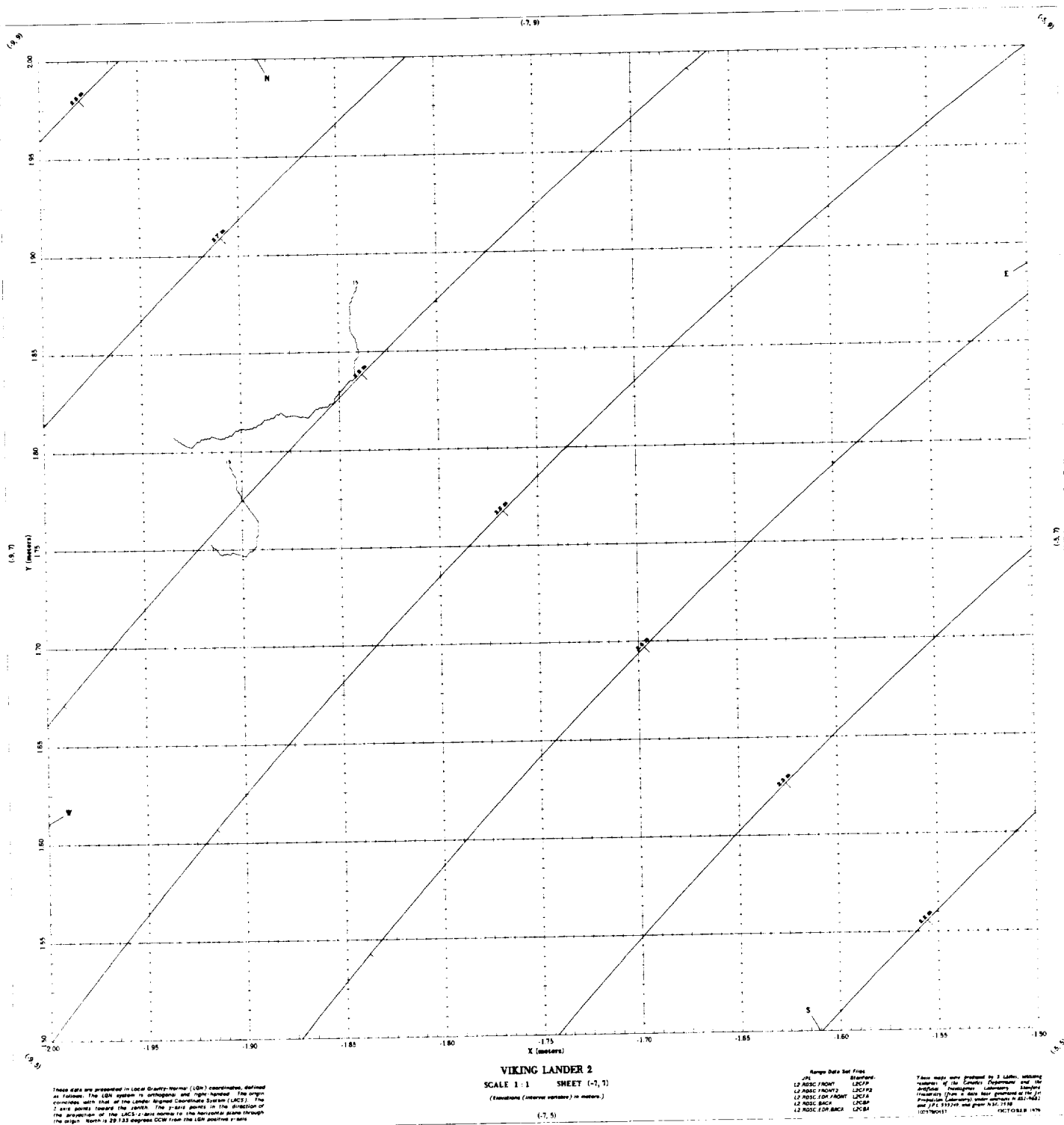


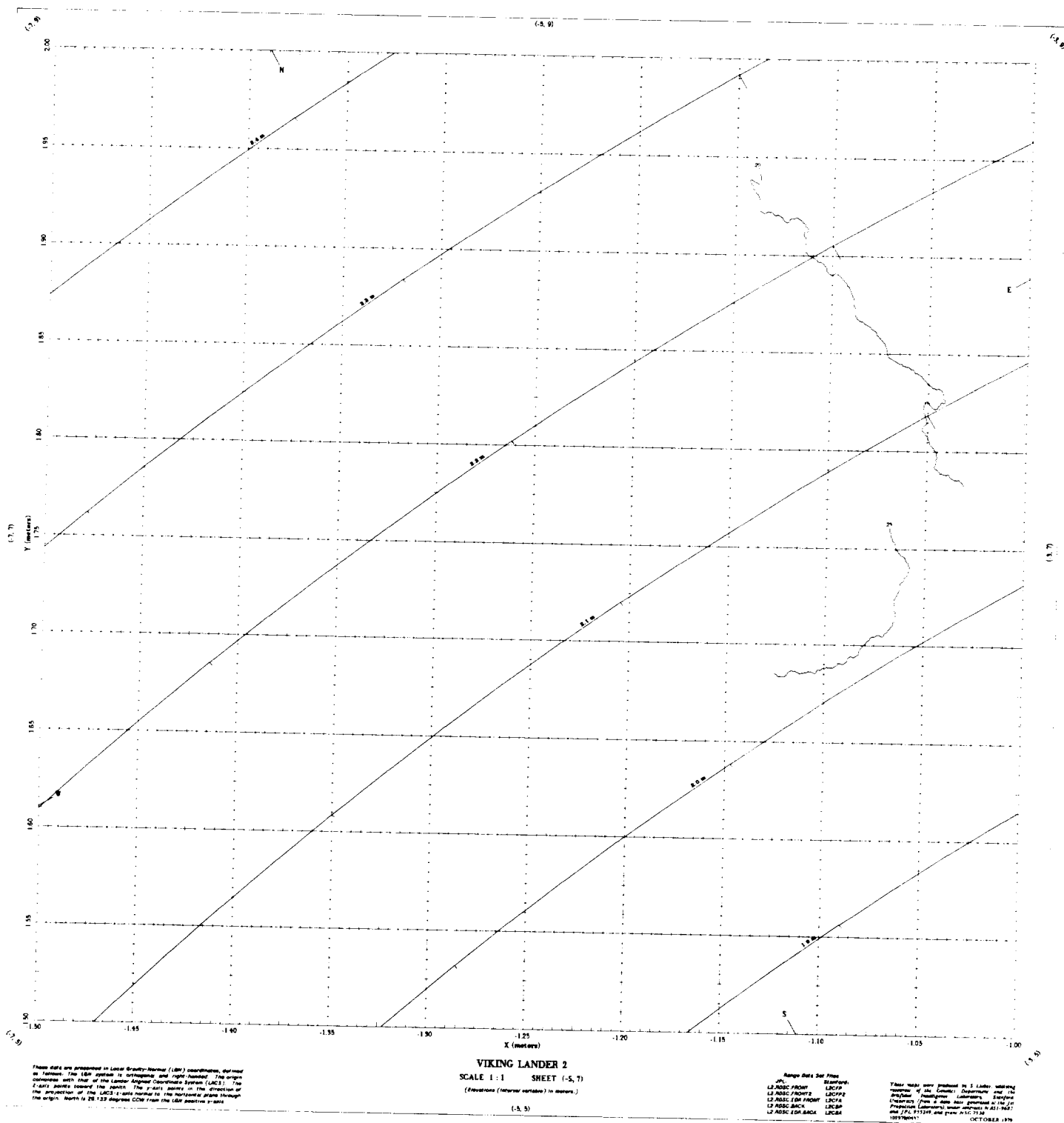












7. VERTICAL PROFILES

7.1 Vertical Profile Mosaic Overlay Stereo Pairs

This section contains mosaic stereo pairs of images into which have been overlayed the systematic vertical profiles. The general remarks of section 4.1 apply.



Fig. 51. Vertical Profile Overlay Mosaic: Camera 1 (left), front left quadrant — from IPL PIC ID 79/10/11/060732.



Fig. 52. Vertical Profile Overlay Mosaic: Camera 2 (right), front left quadrant — from IPL PIC ID 79/03/15/032905.

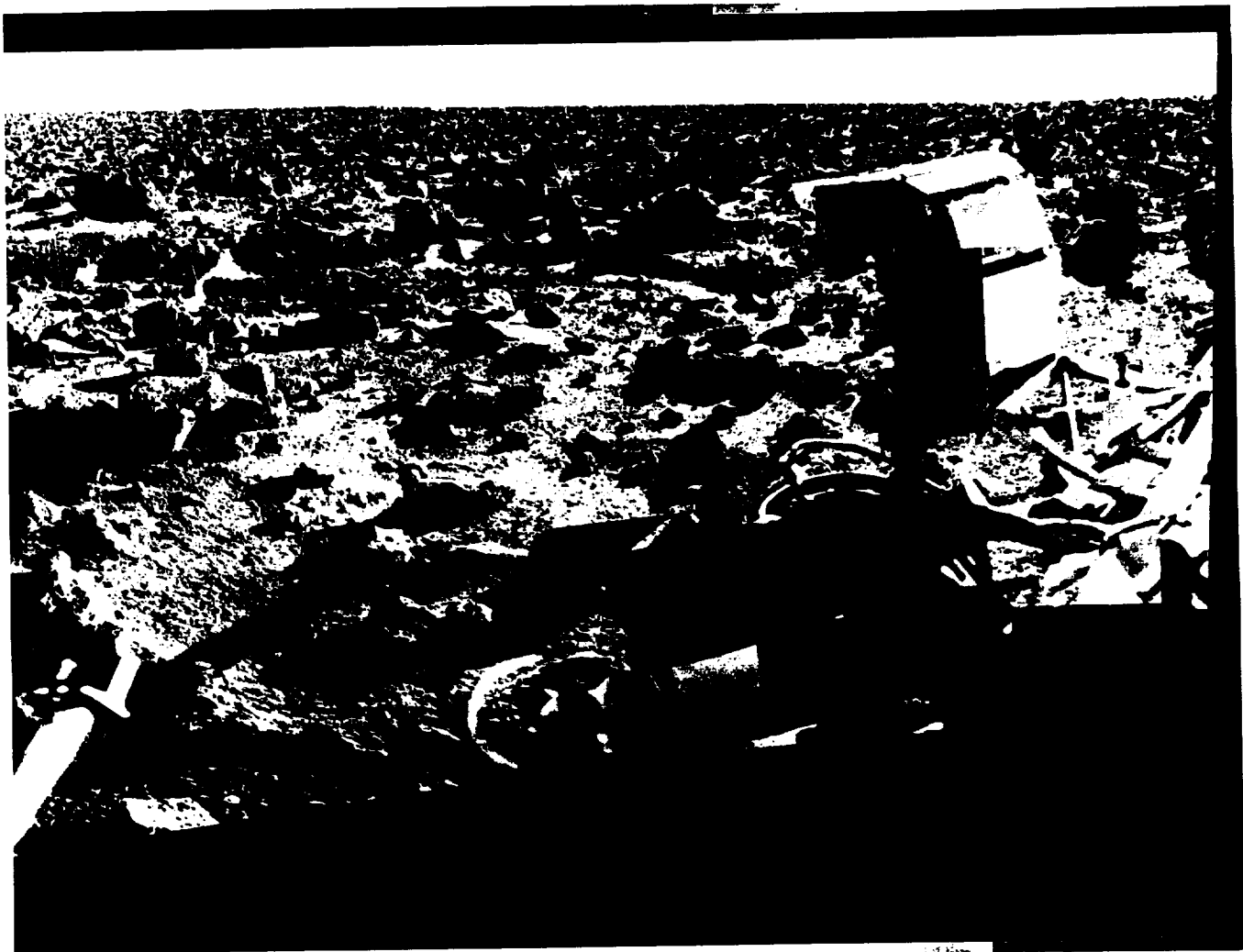


Fig. 53. Vertical Profile Overlay Mosaic: Camera 1 (left), front right quadrant —
from IPL PIC ID 79/10/11/060732.



Fig. 54. Vertical Profile Overlay Mosaic: Camera 2 (right), front right quadrant —
from IPI, PIC ID 79/03/15/032905.



Fig. 55. Vertical Profile Overlay Mosaic: Camera 2 (left), back left quadrant —
from IPL PIC ID 79/10/08/234537.

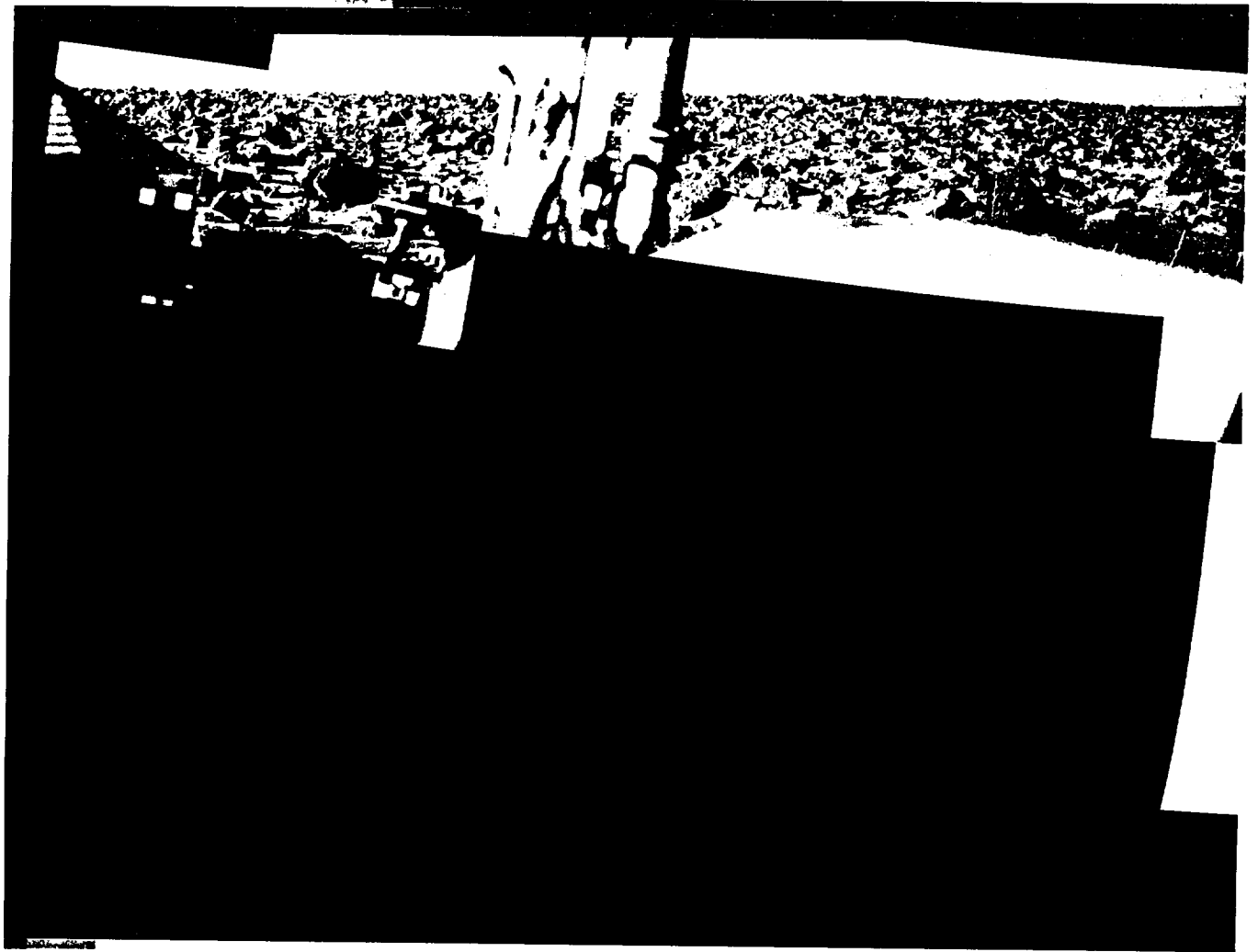


Fig. 56. Vertical Profile Overlay Mosaic: Camera 1 (right), back left quadrant — from H¹L PIC ID 79/04/04/041347.

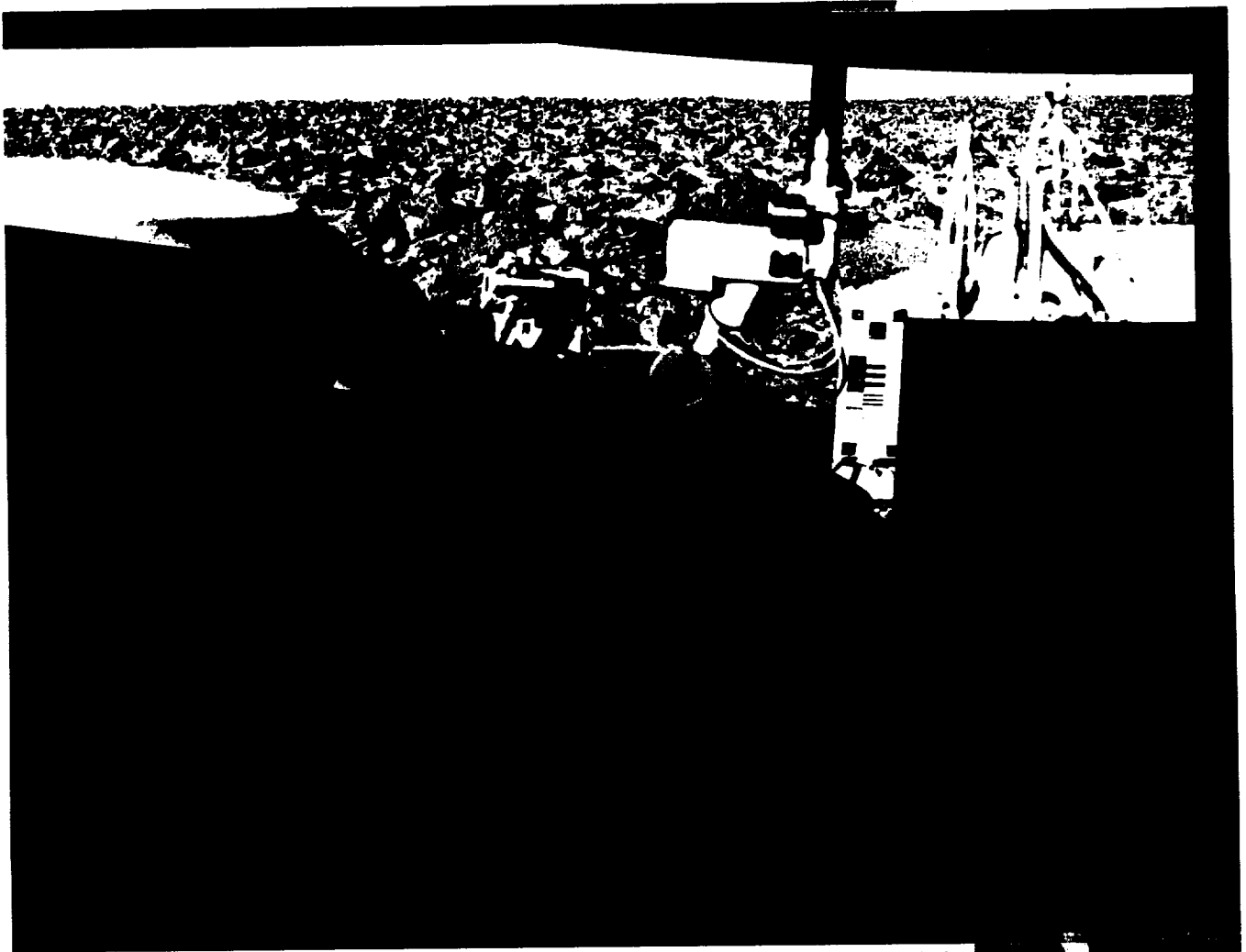


Fig. 57. Vertical Profile Overlay Mosaic: Camera 2 (left), back right quadrant —
from IPL PIC ID 79/10/08/234537.

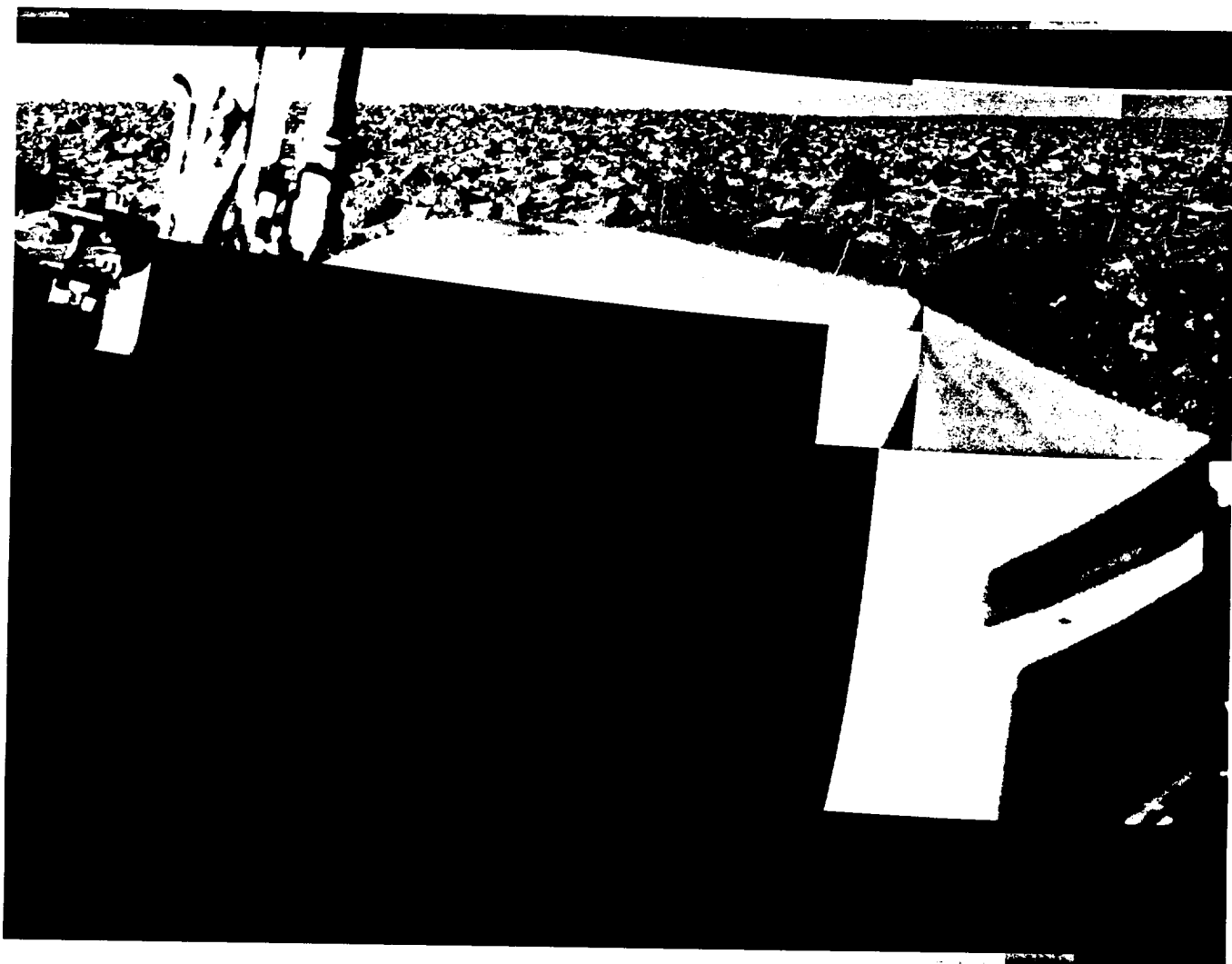


Fig. 58. Vertical Profile Overlay Mosaic: Camera 1 (right), back right quadrant — from IPL PIC ID 79/04/04/041347.

~~FILE~~ INTENTIONALLY BLANK

7.2 Camera Perspective Annotated Vertical Profiles — Camera 1

This section contains camera-1 perspective representations of the elevation contour map data. The nature and purpose of these representations has been explained in section 3.2.1. On opposing pages are presented mosaic overlays of the elevation con-

tour lines, and corresponding camera perspective representation of the map data, with the individual lines of the map data annotated with elevation values. The general discussion of section 4.1 applies to the mosaics of this section.



Fig. 59. Vertical Profile Overlay Mosaic: Camera 1 (left), front left quadrant —
from IPL PIC ID 79/10/11/060732.

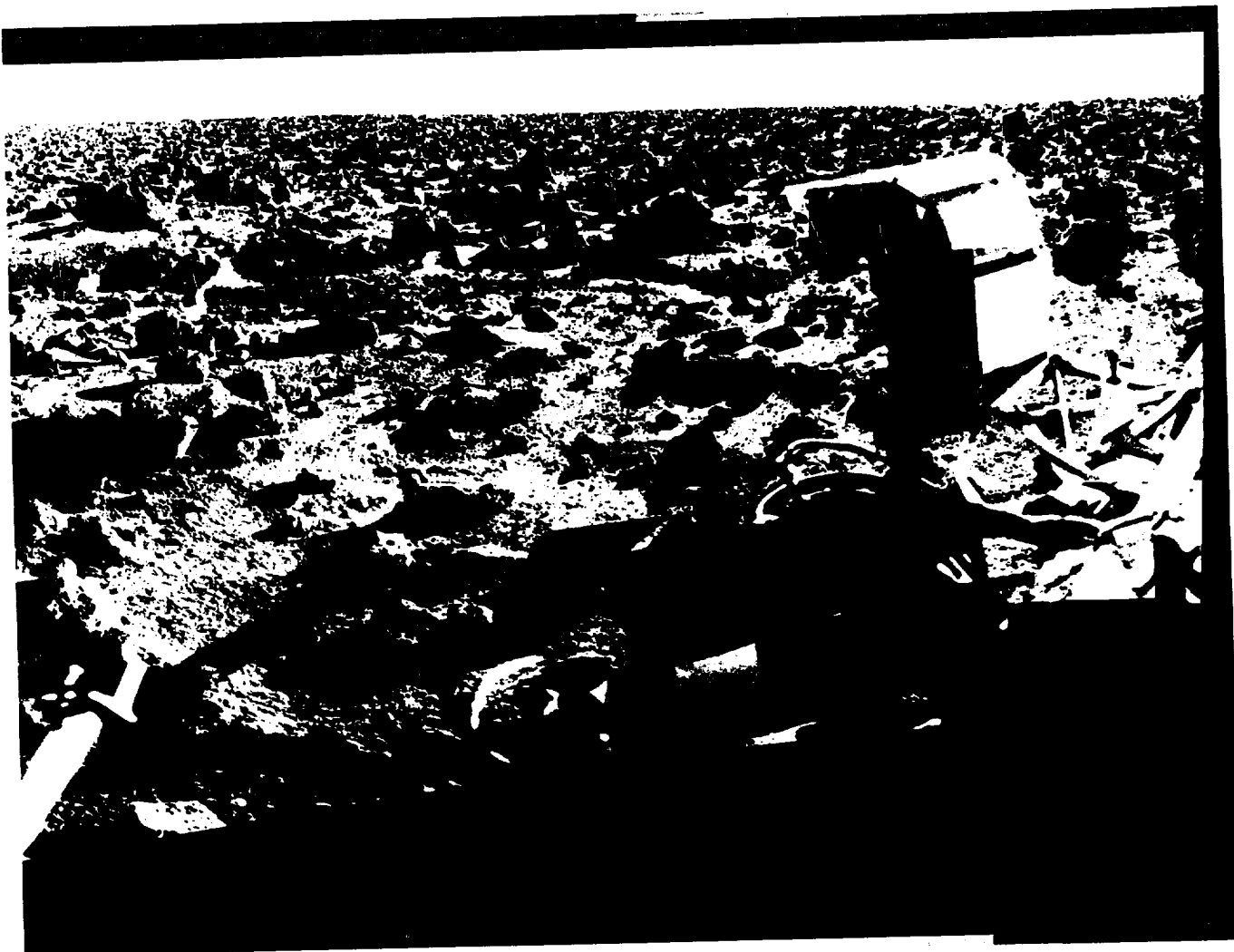


Fig. 61. Vertical Profile Overlay Mosaic: Camera 1 (left), front right quadrant —
from IPL PIC ID 79/10/11/060732.

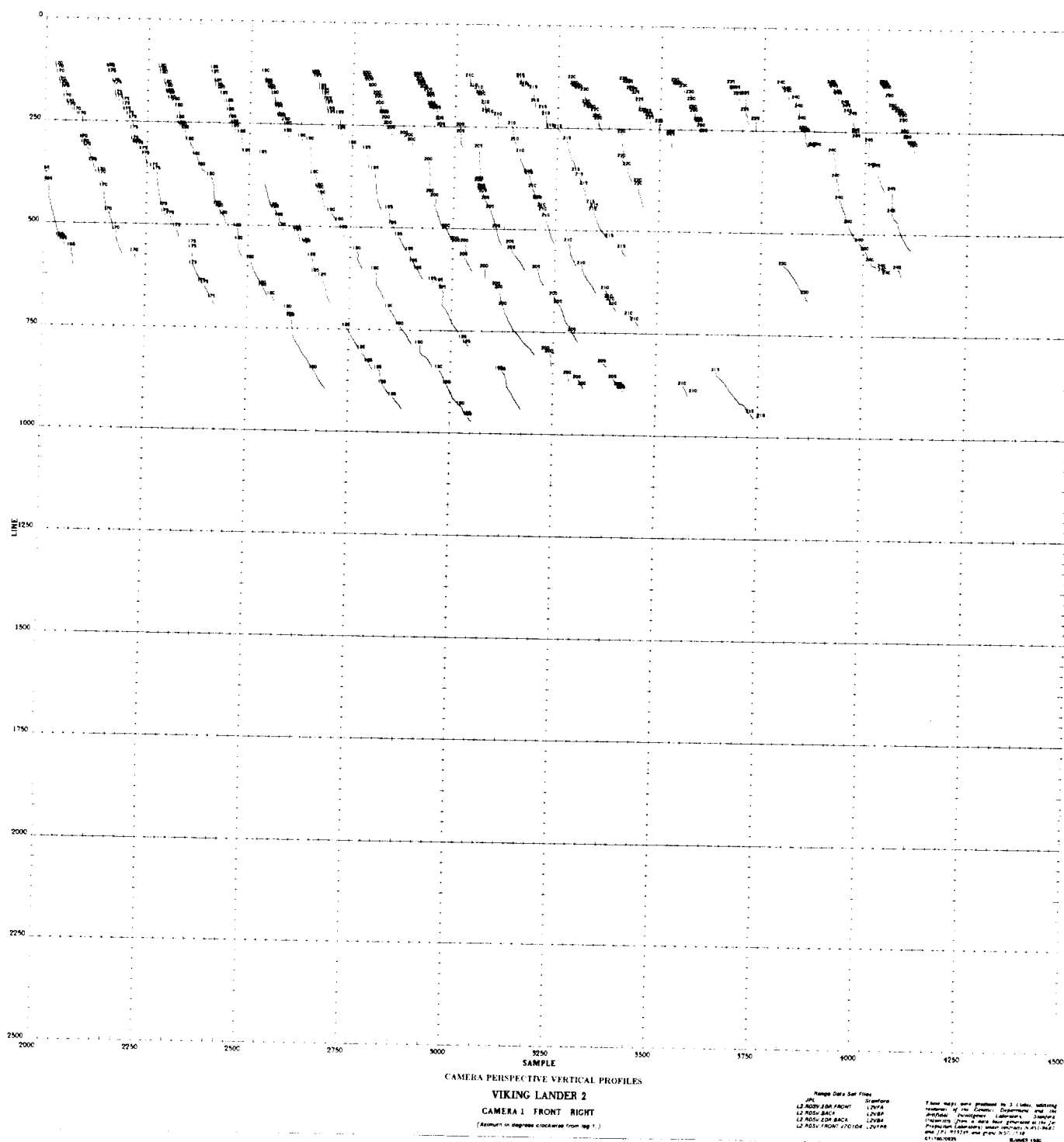


Fig. 62. Camera Perspective Annotated Elevation Vertical Profile Lines:
 Camera 1 (left), front right quadrant.

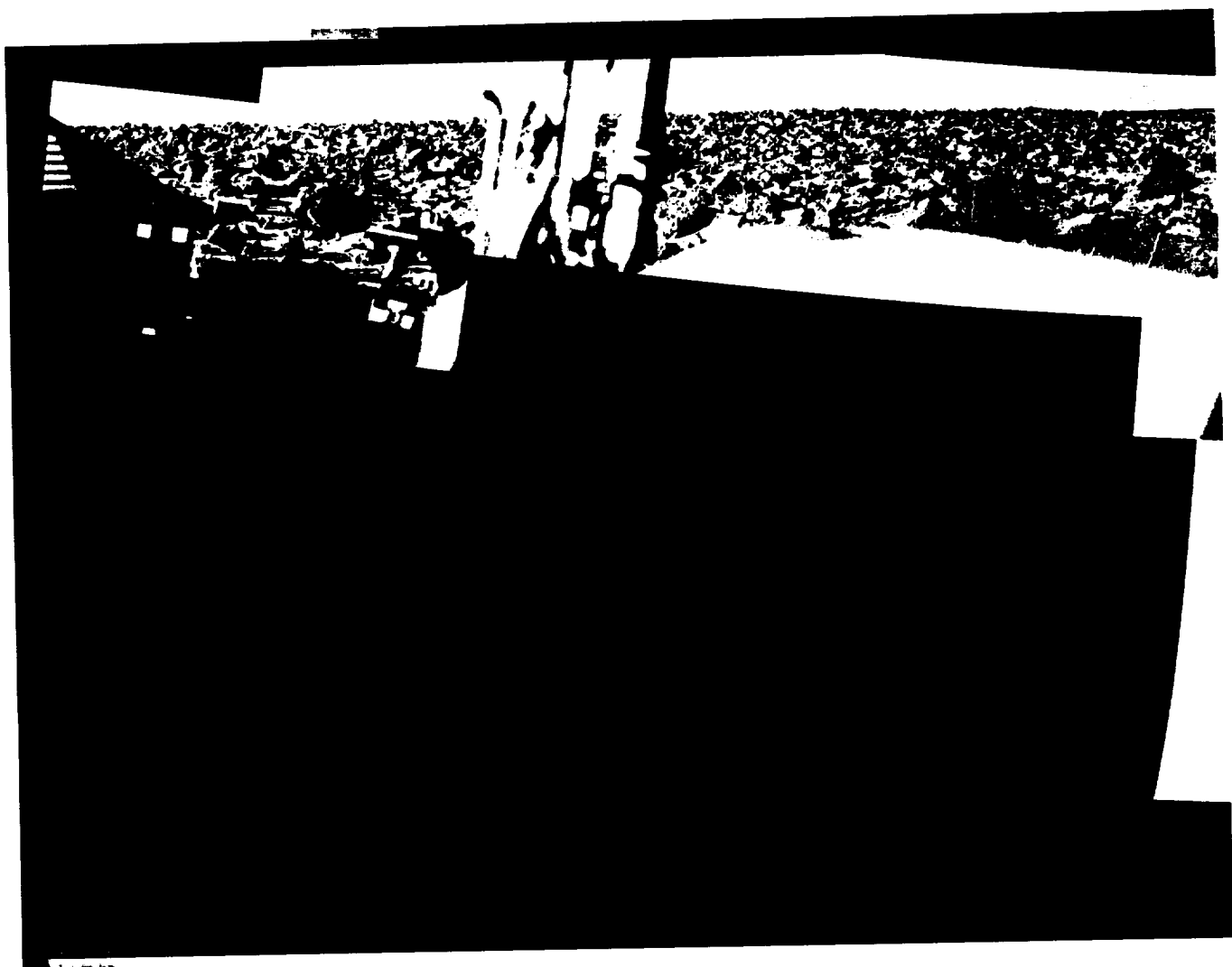


Fig. 63. Vertical Profile Overlay Mosaic: Camera 1 (right), back left quadrant —
from IPL PIC ID 79/04/04/041347.

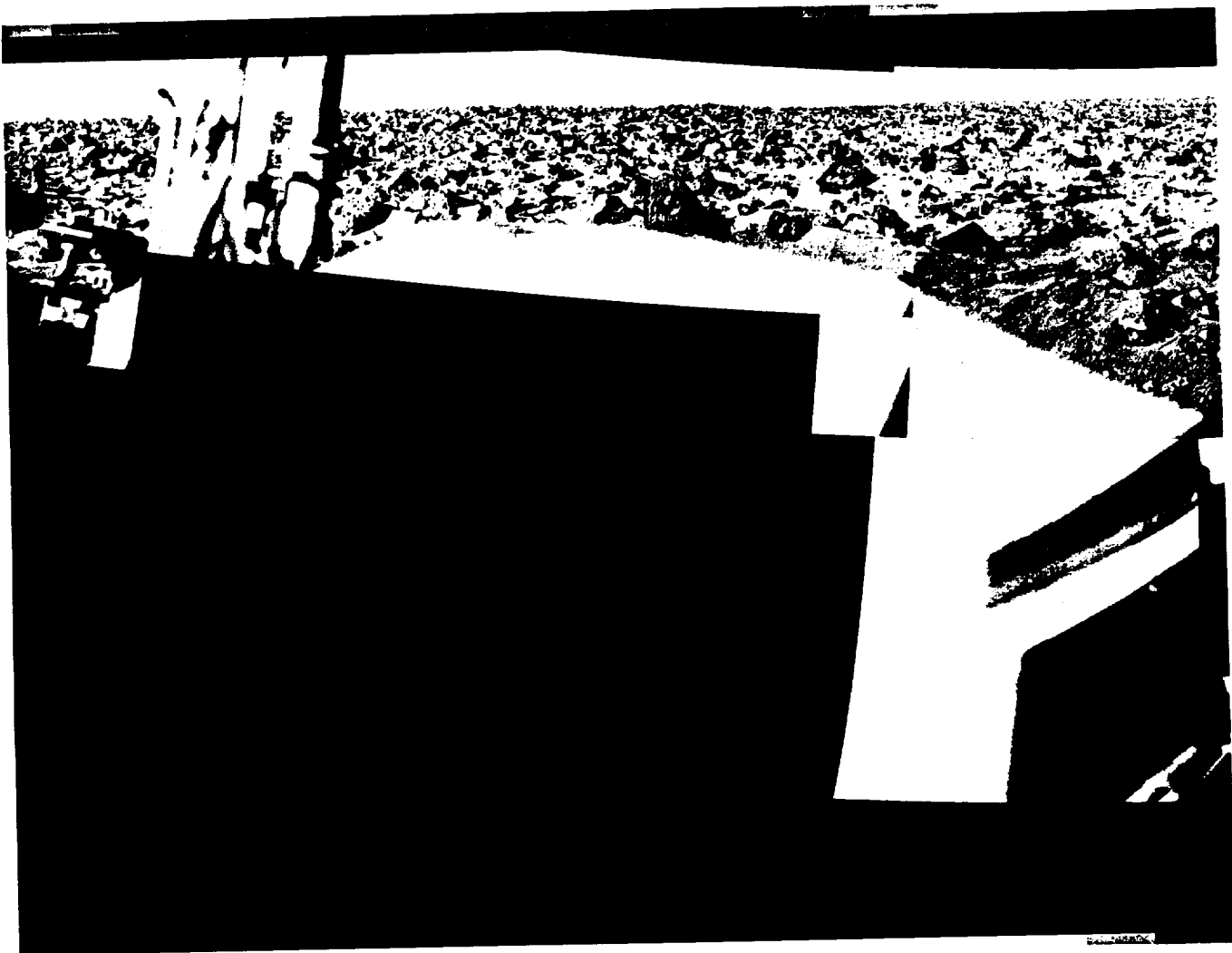
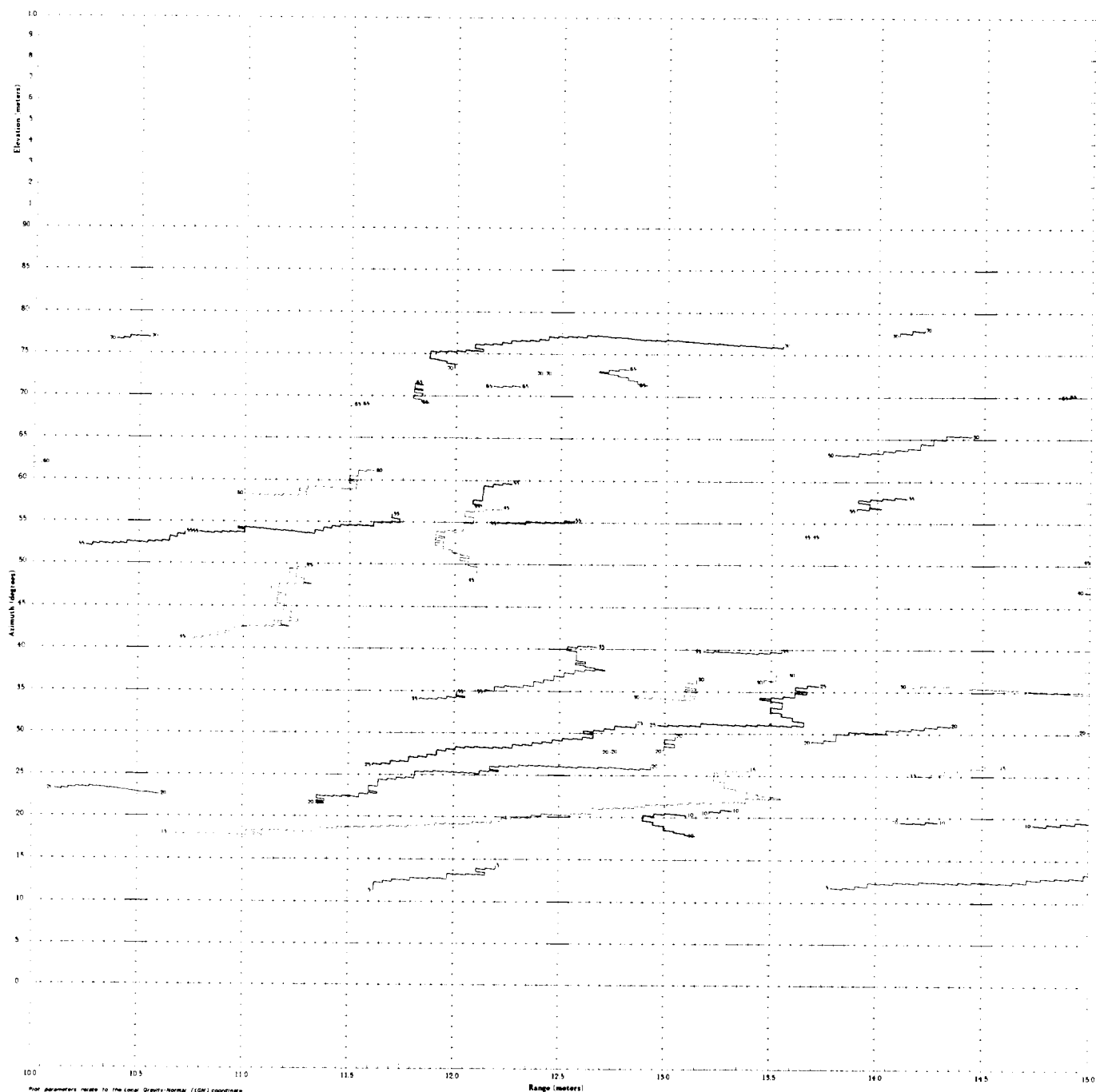


Fig. 65. Vertical Profile Overlay Mosaic: Camera 1 (right), back right quadrant —
from IPL PIC ID 79/04/04/041347.

7.3 Vertical Profile Sheet Collection

This section contains the vertical profile sheet collection. No tabulation of the kind preceding the elevation contour maps has been prepared. The sheets are ordered as follows. First, they are grouped by quadrant, in clockwise order starting from the

back left quadrant (0 - 90°), then within each quadrant by order of decreasing scale number, and within each scale by order of increasing range.

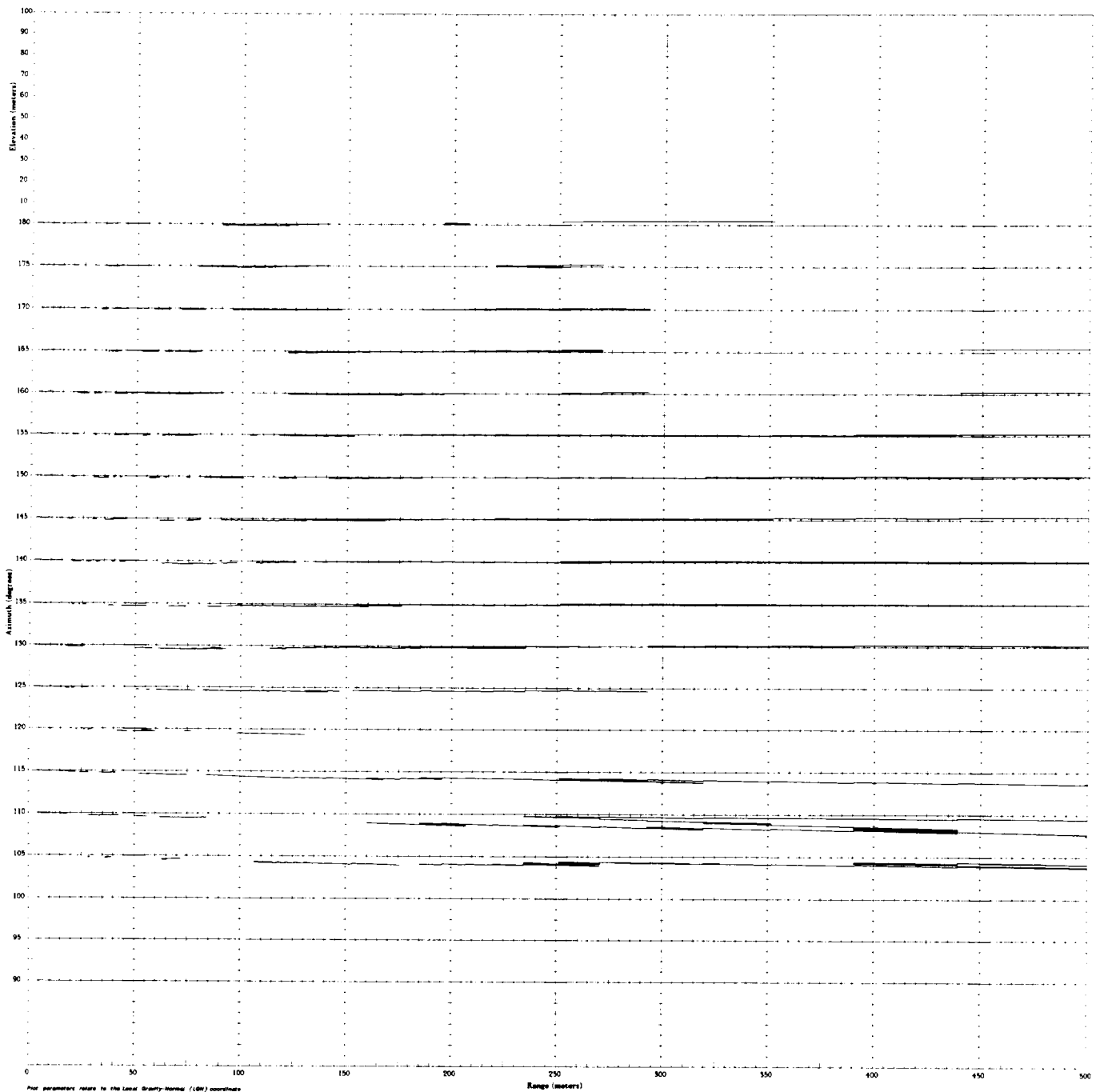


For parameters refer to the local Gravity-Normal (LGN) coordinate system. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lander-Aided Coordinate System (LACS). The Z axis points toward the zenith. The X axis points in the direction of the projection of the LACS Z axis onto the horizontal plane through the origin. The Y axis (azimuth) is the value of the LGN Z coordinate. The Y axis of elevation for each profile is observed by the associated profile azimuth angle. Azimuth is measured clockwise from above relative to a line of reference along the LGN negative Y axis.

VIKING LANDER 2
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 10 TO 15 METERS 0 TO 90 DEGREES

Range Data Set Files
L2 POSN FROM FRONT L2 POSN FROM BACK L2 POSN FROM V-L

These maps were produced by 5 Lander software programs. The Lander software and the Lander software are available from the Viking Lander Project. The Lander software is available from the Viking Lander Project. The Lander software is available from the Viking Lander Project.

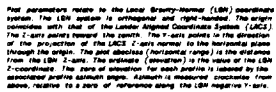


Plot parameters relate to the Local Gravity Network (LGN) coordinate system. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lander Aligned Coordinate System (LACS). The Z axis points toward the planet. The Y axis points in the direction of the projection of the LACS Z axis normal to the horizontal plane through the origin. The horizontal range is the distance from the LGN Z axis. The vertical (elevation) is the height of the LGN Z coordinate. The zero of separation for each profile is labeled by the associated profile azimuth angle. Azimuth is measured clockwise from above, relative to a zero of reference along the LGN negative Y axis.

VIKING LANDER 2
SYSTEMATIC VERTICAL PROFILES
SCALE 1:1000 RANGE 0 TO 500 METERS 90 TO 180 DEGREES

Range Data Set Files
JPL
LEADSV FOR FRONT L2VFR
LEADSV BACK L2VBR
LEADSV FOR BACK L2VBR
LEADSV FRONT V L2VFR

These maps were produced by S. L. Lander, utilizing
measurements of the Lander's altimeter and the
altimeter. The altimeter was used to measure
elevation from a data base generated at the JPL
Planetary Laboratory under contract NAs 11-104-7
and JPL 931249 and grant N51-7132.
CS 100000 N MAP 124 100

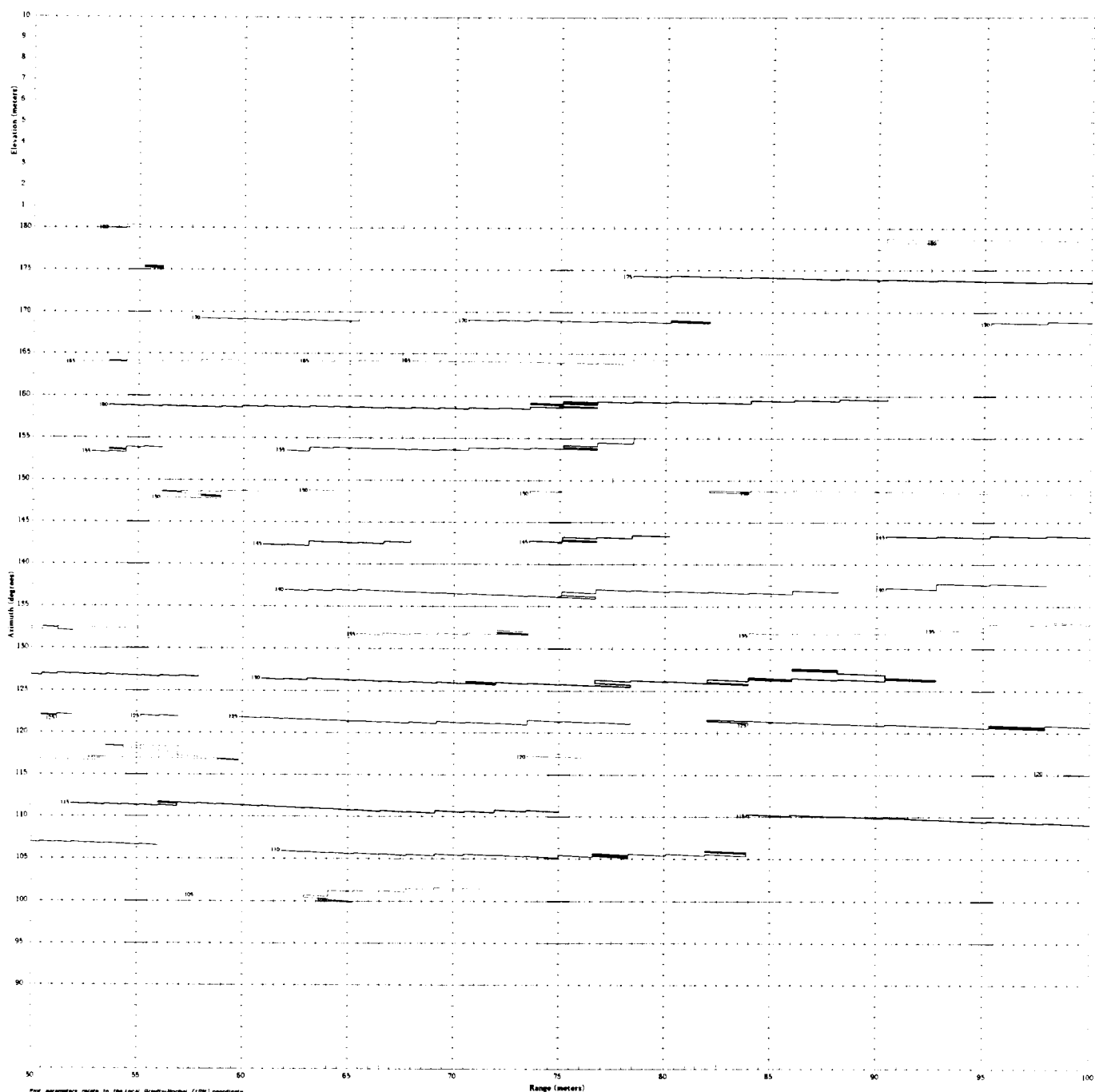


SCALE 1:100 RANGE 0 TO 50 METERS 90 TO 180 DEGREES

Range Date Set From	Standard
JPL	
L2.A03V.EDR.FRONT	L2VFA
L2.A03V.SACR	L2VSB
L2.A03V.EDR.SACR	L2VSB
L2.A03V.FRONT V.	L2VFA

These maps were produced by E. Lieber, with the assistance of the Census Department and the Artificial Intelligence Laboratory, Stanford University (from a data base generated at the Jet Propulsion Laboratory), under contracts N31-7632 and JPL 93529, and grant NSC 7538.

0913027313 MAR 28 1980



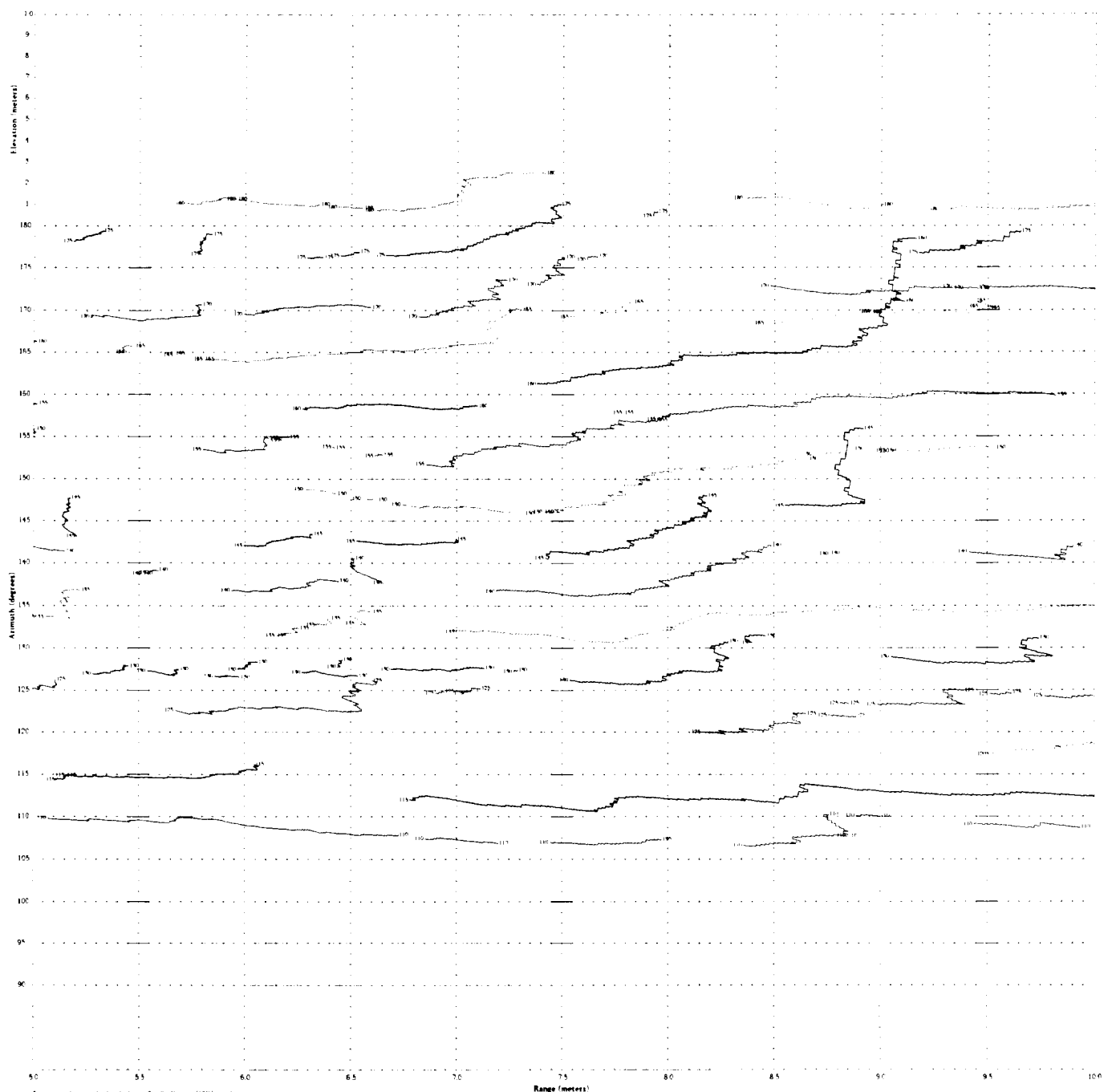
Plot parameters relate to the Local Gravity-Normal (LGN) coordinate system. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lander Angular Coordinate System (LACS). The Z-axis points toward the zenith. The Y-axis points in the direction of the projection of the LACS Z-axis normal to the horizontal plane through the origin. The plot abscissa (horizontal range) is the distance from the LGN Z-axis. The ordinate (elevation) is the value of the LGN Z-coordinate. The line of observation for each profile is oriented by the associated profile azimuth angle. Azimuth is measured clockwise from above, relative to a zero of reference along the LGN negative Y-axis.

VIKING LANDER 2 SYSTEMATIC VERTICAL PROFILES

SCALE 1 : 100 RANGE 50 TO 100 METERS 90 TO 180 DEGREES

Range Data Set Files
APL: Standard
L2 POSN 50M FRONT: LVFR
L2 POSN 50M BACK: LVBP
L2 POSN 100M FRONT: LVFR
L2 POSN 100M BACK: LVBP

These maps were produced by 3 Lines utilizing
images of the Lander's cameras and the
Profile Horizontal Lander Stereo
Coordinate System. The maps were generated by the JPL
Planetary Laboratory under contract N 001-74-52
and JPL 71-171 and were prepared by JPL
DR/USC/111 MARCH 1980

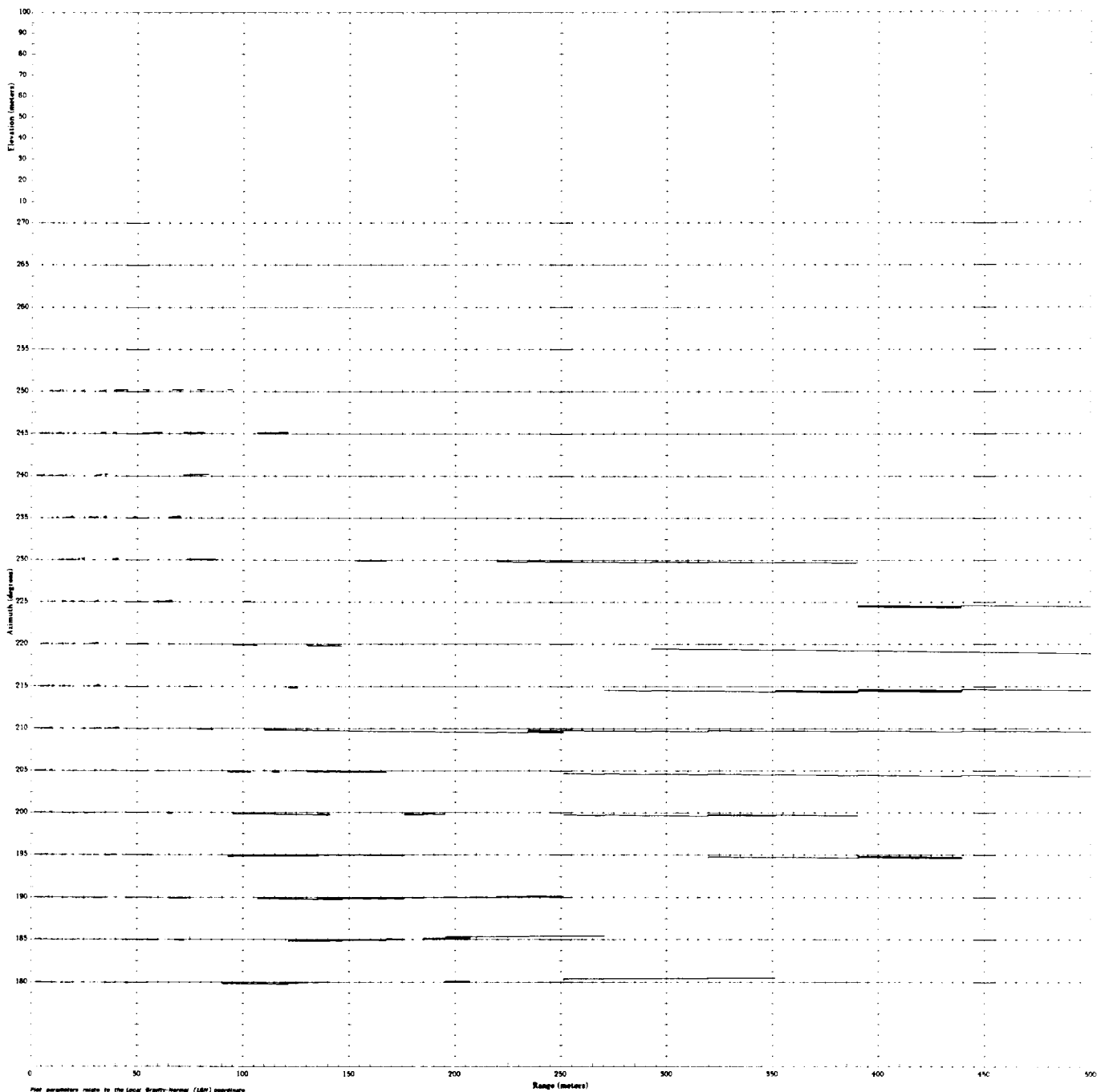


For parameters used in the Lander's coordinate system, the LOR system is orthogonal and right handed. The origin coincides with that of the Lander's coordinate system (LACS). The Z axis points toward the south. The Y axis points in the direction of the direction of the LACS Z axis normal to the horizontal plane through the origin. The X axis points in the direction of the LACS X axis. The elevation (elevation) is the value of the LOR Z coordinate. The range of elevation for each profile is indicated by the associated profile's elevation range. Elevation is measured clockwise from above, relative to a zero of reference along the LOR negative Z axis.

VIKING LANDER 2
SYSTEMATIC VERTICAL PROFILES
SCALE 1:10 RANGE 5 TO 10 METERS 90 TO 100 DEGREES

Range Data for Files
LP: ROLL DATA FRONT 120F8
LP: ROLL DATA BACK 120B8
LP: ROLL DATA BACK 120B8
LP: ROLL DATA BACK 120B8

These maps were produced by the Lander's mapping system. The Lander's mapping system is a computer system that takes data from the Lander's cameras and the Lander's altimeter and produces a map of the Lander's position and the Lander's orientation. The Lander's mapping system is a computer system that takes data from the Lander's cameras and the Lander's altimeter and produces a map of the Lander's position and the Lander's orientation.

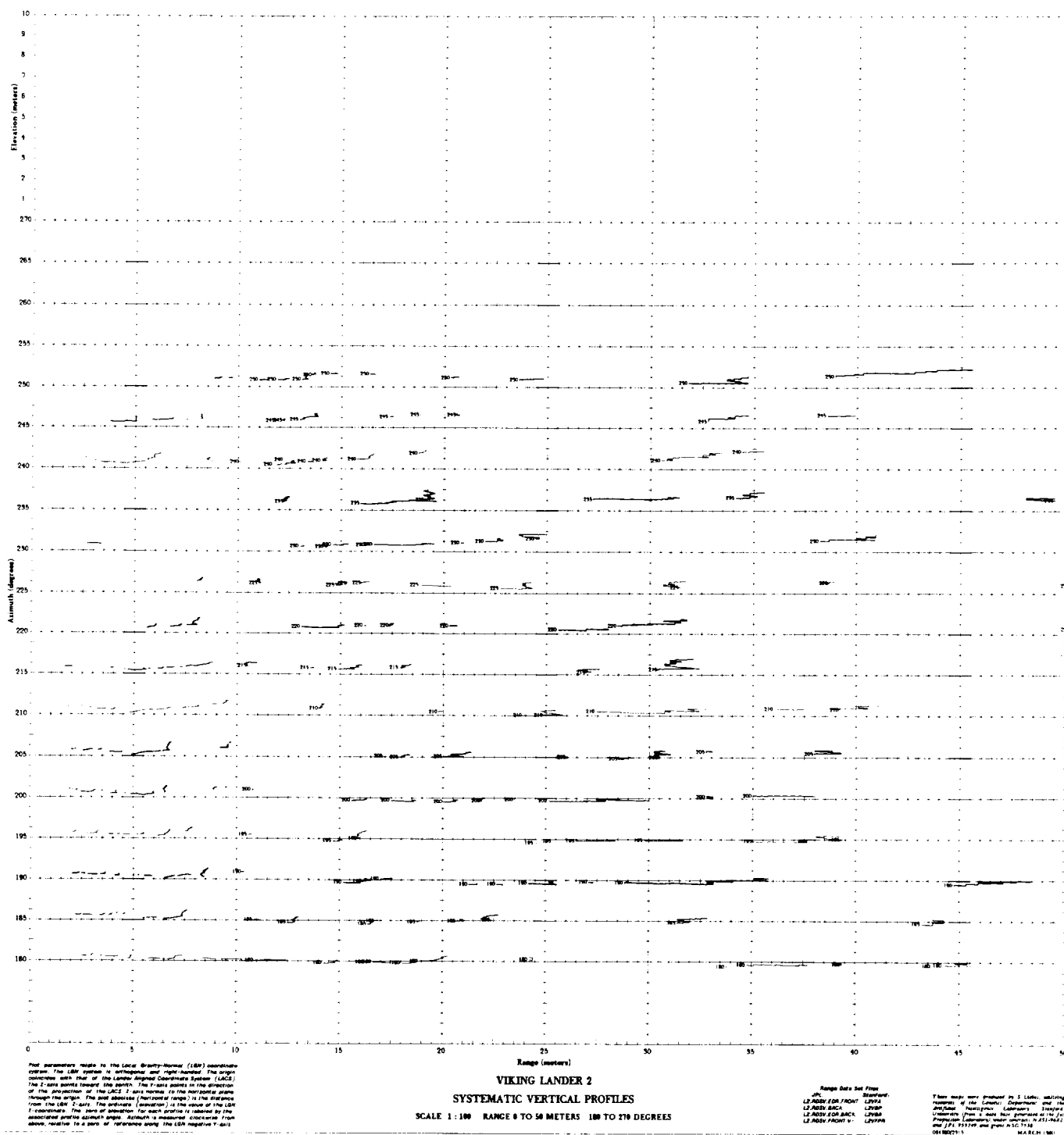


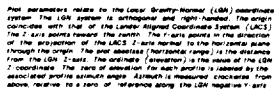
Plot parameters relate to the Local Gravity-Inertial (LGI) coordinate system. The LGI system is antipodal and right-handed. The origin coincides with that of the Lunar-Signed Coordinate System (LSCS). The Z-axis points toward the center. The Y-axis points in the direction of the projection of the LSCS Z-axis normal to the horizontal plane through the origin. The plot parameter (horizontal range) is the distance from the LGI Z-axis. The ordinate (elevation) is the value of the LGI Z-coordinate. The line of observation for each profile is labeled by the associated profile azimuth angle. Azimuth is measured clockwise from above, relative to a zero of reference along the LGI negative Y-axis.

VIKING LANDER 2
SYSTEMATIC VERTICAL PROFILES
SCALE 1:1000 RANGE 0 TO 500 METERS 180 TO 270 DEGREES

Range Data Set File
JPL
L2 ADSS: 20H / FRONT L2VFA
L2 ADSS: 20H / BACK L2VFB
L2 ADSS: 20H / BACK L2VBA
L2 ADSS: 20H / BACK L2VBB

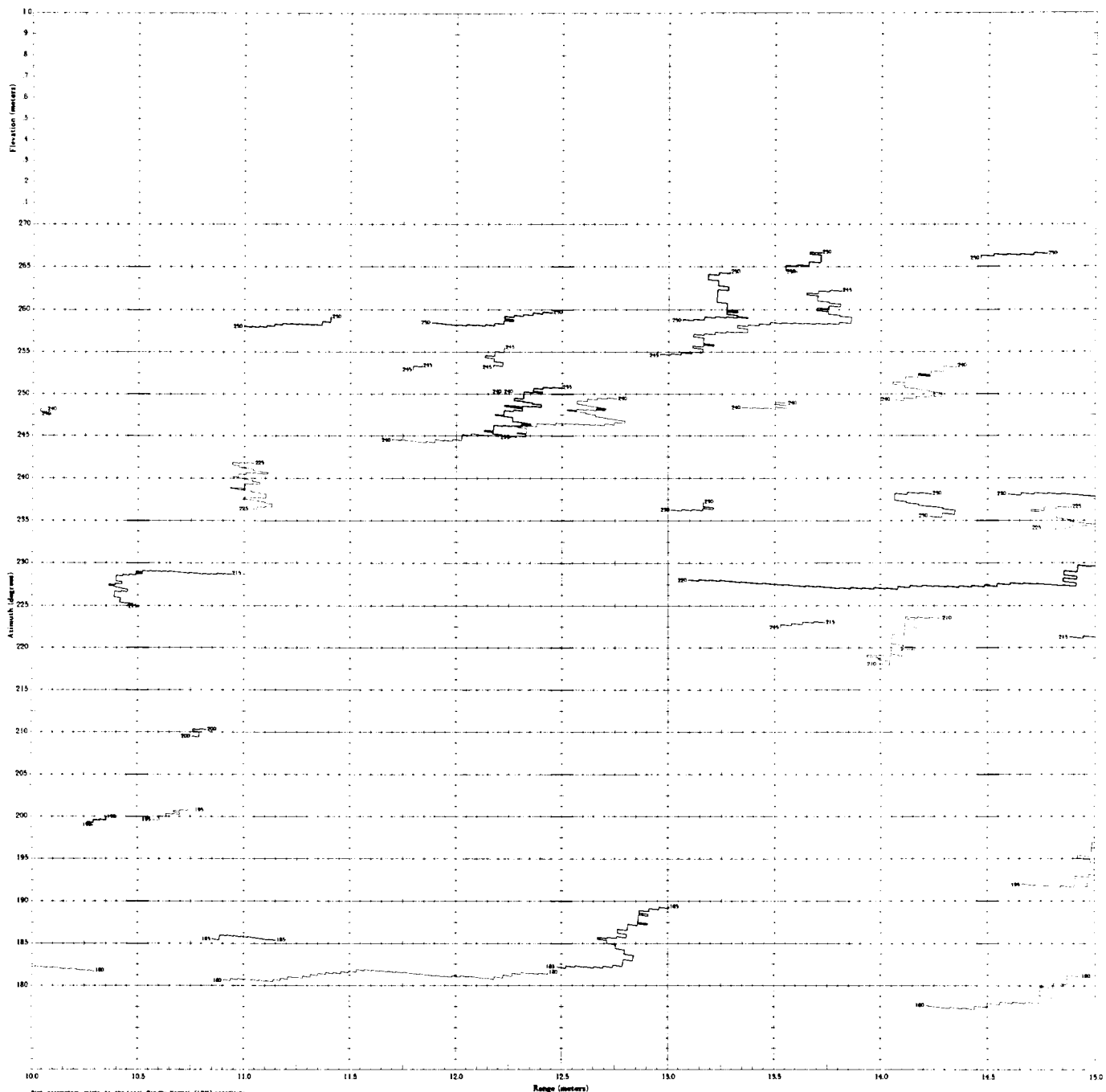
These maps were produced by 3 Lunar Orbiter
reprints of the Lunar Orbiter and the
Apollo Lunar Module's Lander's Lander's
reprints from a data base generated at the
Photographic Laboratory, under contract to NASA,
and JPL staff and from NIT, 1978
(1) March 1980





Range Data Set Files	
JPL	Standard
L2.ADSV.FDR.FRONT	L2VFA
L2.ADSV.BACK	L2VBP
L2.ADSV.FDR.BACK	L2VBA
L2.ADSV.FRONT.V	L2VFP

These maps were produced by S. Lieber, visiting researcher of the Genetics Department and the Biological Population Laboratory, Stanford University (from a data base generated at the JPL Population Laboratory), under contracts N-AS1-90-21 and JPL 85549 and grant NSC 7436.

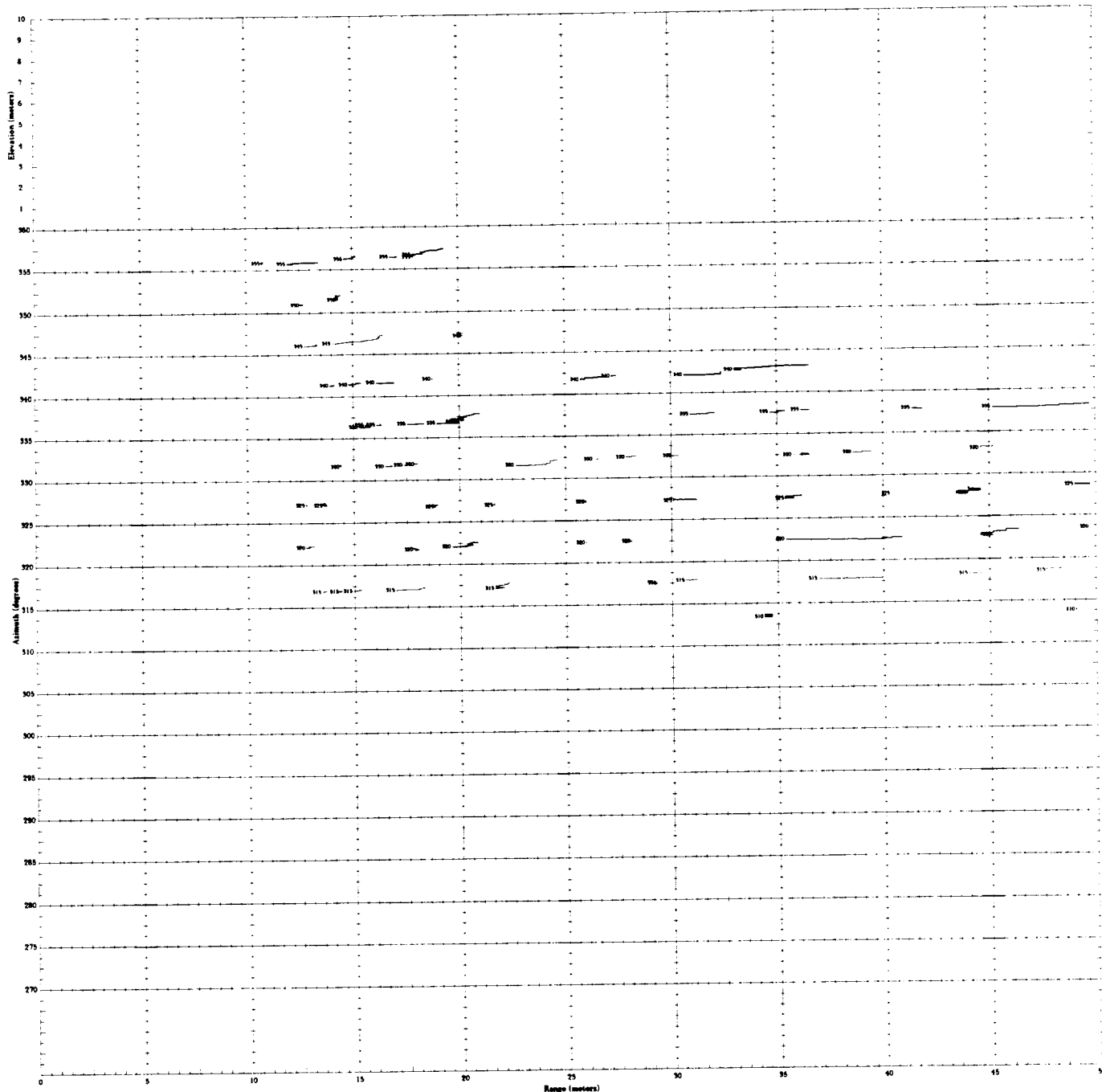


For parameters refer to the Local Gravity Network (LGN) coordinate system. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lunar Apogee Coordinate System (LACS). The x-axis points toward the point. The y-axis points in the direction of the projection of the LACS z-axis normal to the horizontal plane through the origin. The profile azimuth (horizontal range) is the distance from the LGN z-axis. The ordinate (elevation) is the value of the LGN z-coordinate. The range of elevation for each profile is labeled in the associated profile azimuth angle. Azimuth is measured clockwise from above, relative to a zero of reference along the LGN negative y-axis.

VIKING LANDER 2
SYSTEMATIC VERTICAL PROFILES
SCALE 1:11 RANGE 10 TO 15 METERS 180 TO 270 DEGREES

Range Data Set Files
JPL, Stanford
L2-RDSV-LGN-FRONT, L2R1F
L2-RDSV-LGN-BACK, L2R1B
L2-RDSV-LGN-BACK, L2R1B
L2-RDSV-FRONT-Y, L2R1F

These maps were produced by S. Loken, utilizing
measurements of the Gravity, Altimeter, and the
Photometer, Photometer, Lander, and the
Lander and the data from the Lander and the
Photometer (Lander) under contract N45-74-627
and JPL, 1974-1975 and given as 10-71-10
09110075-1 MARCH 1980



Plot parameters relate to the Local Gravity-Reference (LGR) coordinate system. The LGR system is referenced and right-handed. The origin coincides with that of the Lander Regional Coordinate System (LRCS). The Z-axis points toward the north. The Y-axis points in the direction of the perpendicular of the LRCS Z-axis toward the horizon plane through the origin. The plot abscissa (horizontal range) is the distance from the LGR Z-axis. The ordinate (elevation) is the value of the LGR Z-coordinate. The zero of elevation for each profile is labeled by the associated profile abscissa number. Abcissa is measured, read-left, from above, relative to a zero of reference along the LGR negative Y-axis.

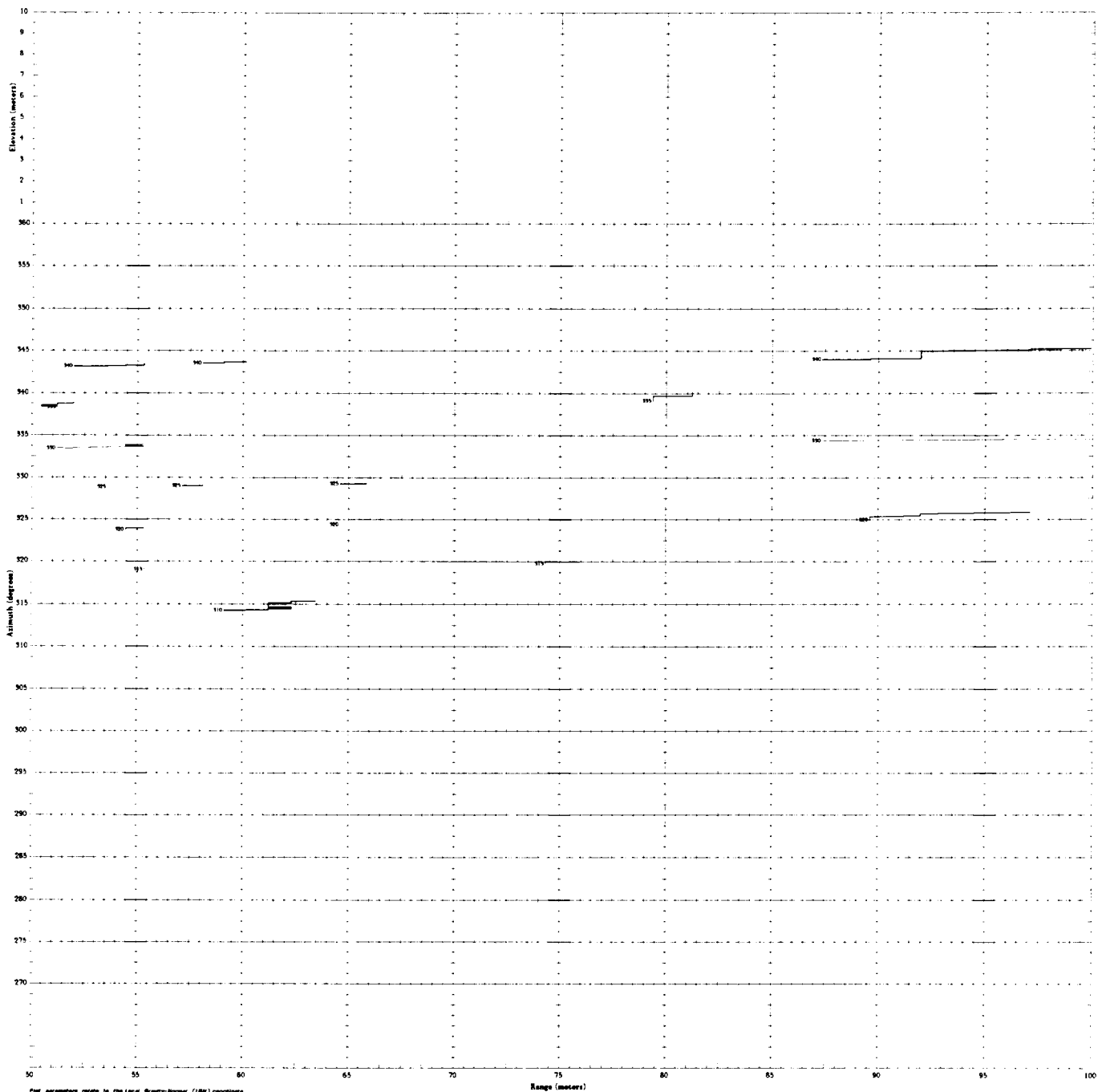
VIKING LANDER 2

SYSTEMATIC VERTICAL PROFILES

SCALE 1:100 RANGE 0 TO 50 METERS 270 TO 360 DEGREES

Range Data Set File
 L2 ADDV.FDA FRONT L2VPA
 L2 ADDV.BACK L2VPA
 L2 ADDV.FDA BACK L2VPA
 L2 ADDV.FRONT V L2VPA

This page was produced by 3 Lander-robotic
 cameras of the Lander's Department and the
 Lander's Department. The Lander's Department
 cameras (from 4 data files generated in the
 Lander's Department) were used to produce this
 page. PPL 9112M and photo ASC 7118
 10/10/71

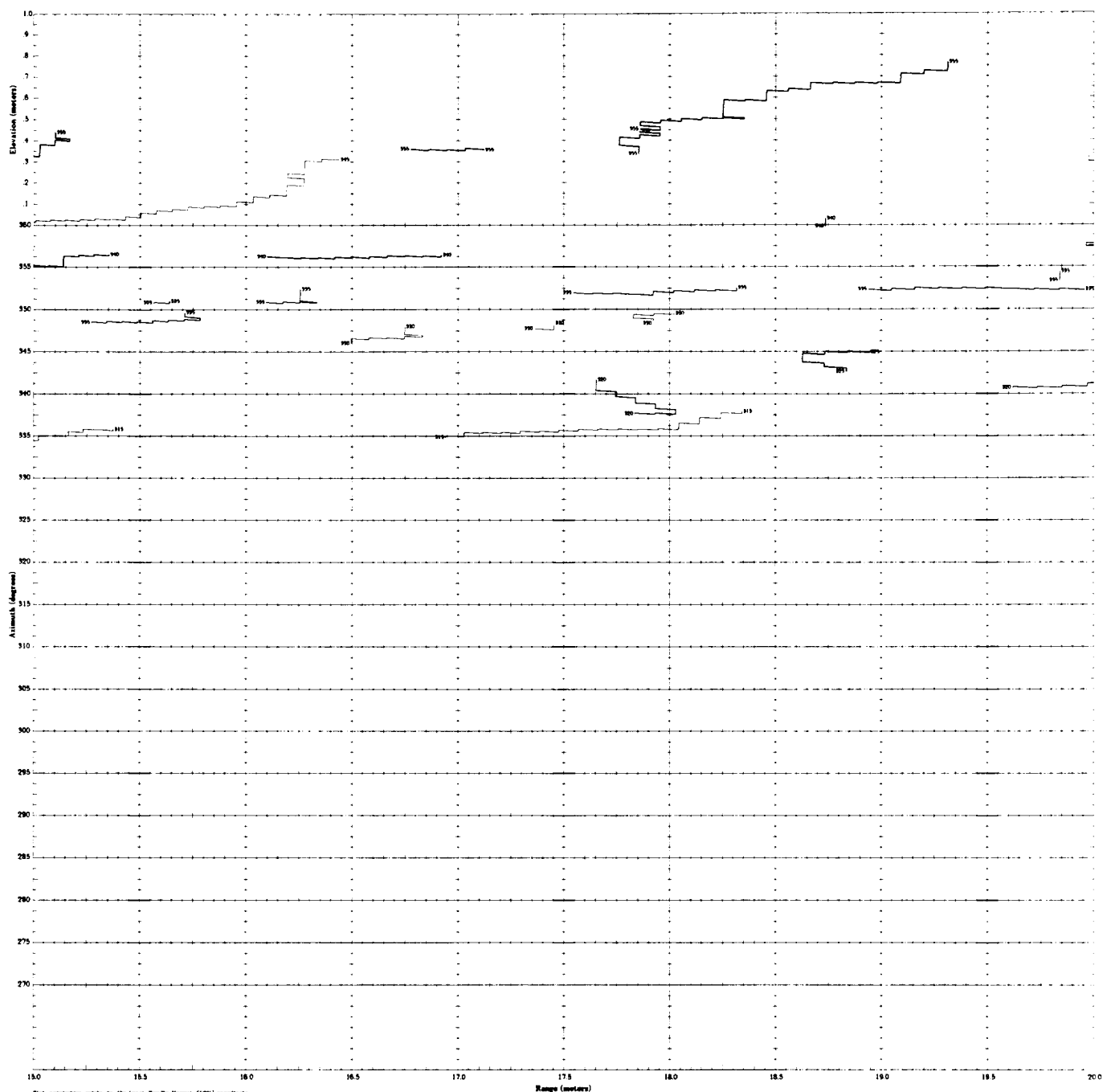


Plot parameters relate to the Local Gravity-Normal (LGN) coordinate system. The LGN system is orthogonal and right handed. The origin coincides with that of the Lander-Surface Coordinate System (LSCS). The Z-axis points toward the center. The Y-axis points in the direction of the projection of the LSCS Z-axis normal to the horizontal plane through the origin. The point distance (horizontal range) is the distance from the LSCS Z-axis. The elevation (elevation) is the height of the LGN Z coordinate. The zero of elevation for each profile is labeled by the associated profile azimuth angle. Azimuth is measured clockwise from above, relative to a zero of reference using the LGN negative Y axis.

VIKING LANDER 2
SYSTEMATIC VERTICAL PROFILES
SCALE 1:100 RANGE 50 TO 100 METERS 270 TO 360 DEGREES

Range Data Set Files
JPL: Scanfile
LE ADSV FOR FRONT: LQVFA
LE ADSV BACK: LQVFB
LE ADSV FOR BACK: LQVBA
LE ADSV FRONT V: LQVFAV

These data were produced by 3 Lines utilizing resources of the JPL/Caltech Department and the JPL/Caltech Department. Lander's location (elevation) from 0 meter base generated at the JPL/Caltech Department under contract N0001-80-2-0001 and JPL 89709 and from N.S.C. 118. ON MARCH 1981



Plot parameters relate to the Local Gravity-Normal (LGN) coordinate system. The LGN system is orthogonal and right-handed. The origin coincides with that of the Lander Aligned Coordinate System (LACS). The Z-axis points toward the zenith. The Y-axis points in the direction of the projection of the LACS Z-axis normal to the horizontal plane through the origin. The east direction (horizontality) is the direction from the LGN Z-axis. The azimuth (azimuth) is the angle of the LGN Z-coordinates. The range elevation for each profile is obtained by the measured profile azimuth angle. Azimuth is measured clockwise from above surface to a line of reference, over the LGN normal Z-axis.

VIKING LANDER 2

SYSTEMATIC VERTICAL PROFILES

SCALE 1 : 10 RANGE 15 TO 20 METERS 270 TO 360 DEGREES

Range Data Set Files	Standard
ADSV_EOR_FRONT	LYNFA
ADSV_BACA	LYVBP
ADSV_EOR_BACK	LYVBA
ADSV_FRONT_V-	LYVFR

These maps were produced by S. Linder, utilizing resources of the Central Department and the Artificial Intelligence Laboratory, Stanford University (from a data base generated at the JPL Population Laboratory), under contract NASS-94-02 and JPL 955249, and from ALC 7950.

8. GLOSSARY

CCW	Counter Clockwise
CW	Clockwise
IPL	Image Processing Laboratory, JPL
JPL	Jet Propulsion Laboratory, 4800 Oak Grove Drive, Pasadena, CA 91103
LACS	Lander Aligned Coordinate System
LGN	Local Gravity-Normal coordinate system (equivalent to Lander Gravity-Normal System)
LMS	Local Mars coordinate System
RDS	Range Data Set (the fundamental Stereo System product)
RANGER	Stanford Stereo Station computer program (at IPL)
T^{SR}	Rotation Transformation matrix, from LACS ("R" Reference coordinate system) to LMS ("S")

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APPENDIX A COORDINATE SYSTEMS AND LANDER ORIENTATION

A1 Lander Aligned Coordinate System (LACS)

Figure A1 schematically indicates the placement of the Viking lander cameras. The coordinate system shown in the figure is the Lander Aligned Coordinate System (LACS). The LACS was selected by mission management to be the lander prime reference coordinate system. The curious placement of the origin, not only outside of the lander but even below the

nominal landing plane, apparently suited the convenience of engineers having lander/orbiter system integration responsibilities. The coordinate system employed for recording all range information within the RDSs is the LACS. This of course means that, since the landers are tilted, no value of coordinate remains fixed when tracing an elevation contour.

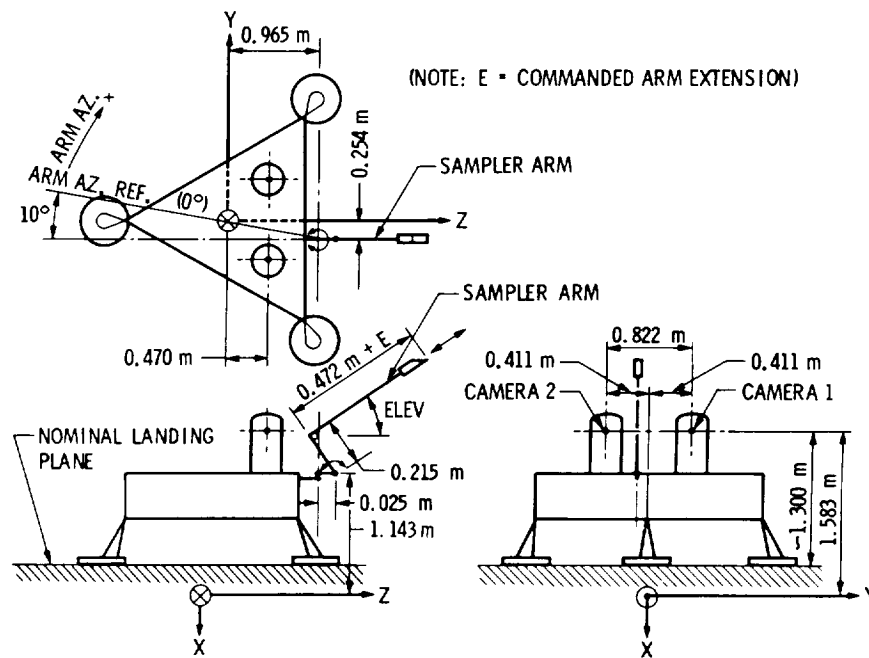


Fig. A1. Lander, camera, and LACS coordinate system configuration.

A2 Local Mars System (LMS)

For the purpose of supporting map generation, a coordinate system called the Local Mars System (LMS) was introduced within RANGER. This was not a unique system, but one that was variously defined according to the circumstances of the particular mapping operation. For the purpose of generating systematic elevation contour maps, the LMS was defined as follows. It was right-handed, the LMS z-axis pointed directly upward, toward the zenith, the LMS y-axis pointed toward the north, and the LMS x-axis pointed toward the east. The matrix within RANGER that defines the transformation from the LACS to the LMS is referred to as the T^{SR} (from "R", Reference LACS, to "S", LMS.) The values of these transformation matrices for the two landers are given in sections 4.1 and 4.2 of the present appendix.

A3 Local Gravity-Normal System (LGN)

The borders of the few preliminary systematic maps plotted at JPL were aligned with the LMS; the cardinal compass directions were parallel to edges of these maps. However, since the cameras were located at equal elevations above the lander deck, they were directly or indirectly responsible for substantial mutual obscurations of the Martian surface. Furthermore, stereo ranging accuracy degenerates as the viewing directions approach a parallel to the intercamera baseline. These two factors contribute an inherent geometrical symmetry to the mapping characteristics relative to a plane passing perpendicular to the intercamera baseline through its midpoint. For this reason a coordinate system was introduced specifically for plotting the systematic elevation contour and vertical profile maps. This system is variously referred to as the Lander Gravity-Normal coordinate system (in early informal memoranda of A. A. Schwartz), and in this report as the Local Gravity-Normal (LGN) system. Though, formally, the LGN is defined in terms of the LACS, the system, in fact, may be obtained from the LMS by a rotation about the zenith. The LGN system is defined as follows: It is orthogonal and right-handed. The origin coincides with that of the Lander Aligned Coordinate System (LACS). The z-axis points toward the zenith. The y-axis points in the direction of the projection of the LACS z-axis normal to the horizontal plane through the origin. The transformation matrices from either LACS or LMS to LGN, and vice versa, may be obtained from the T^{SR} given in sections 4.1 and 4.2, below.

To the extent that lander tilt may be considered small, the LGN may be visualized as having its x-axis approximately antiparallel to the LACS y-axis (and thus pointing roughly parallel to the direction from the "left-hand" camera, or camera number 1, toward the "right-hand" camera, or camera number 2.) The LGN y-axis points roughly parallel to the LACS z-axis, or in the horizontal direction outward and approximately forward from the lander, and roughly perpendicular to the intercamera baseline.

A4 Lander Orientation

The lander orientation information utilized in the preparation of the systematic maps, and represented by the T^{SR} matrices, was obtained from the inertial reference units onboard the landers. A comparison of lander performance with a computation of lander orientation based on output from the onboard guidance systems indicates that the standard deviation of the lander orientation parameters is $\pm 0.4^\circ$ [Villyard, 1977; K. W. Villyard, private communication, 1981]. Independent evidence is consistent with the inertial reference unit data. The Viking 2 landing site is very flat. We do note that when the Stanford Stereo System is commanded to encircle the lander with a remote gravity-aligned elevation contour line, in the neighborhood of zero elevation, the contour line falls very closely upon the visual horizon. We furthermore note that the contour lines of the maps for Viking lander 2 hover very closely around zero elevation. Additionally, the orientation information obtained from the inertial reference units is consistent with rough independent checks performed by the lander imaging team [Mutch, 1976a].

The systematic map T^{SR} matrices for the two landers, as well as the lander orientation parameters derived from these matrices are indicated in the sections below.

A4.1 Lander 1

The systematic map T^{SR} matrix for Viking lander 1 is

$$T^{SR} = \begin{pmatrix} +.0503457 & +.7858010 & +.6164240 \\ -.0136545 & +.6176890 & -.7862990 \\ -.9986370 & +.0311701 & +.0418279 \end{pmatrix}.$$

The orientation angles obtained from the lander 1 T^{SR} are shown below. The large number of significant figures is shown only for compatibility with the T^{SR} . As noted in above, the standard deviation error in tilt and orientation of the lander relative to north is estimated to be $\pm 0.4^\circ$ (Note: When viewing normal to the deck of the lander, the LACS negative z-axis coincides with the projection of leg number 1, the aft pointing leg.):

North: 38.09489° cw from leg 1.

Tilt magnitude: 2.991820°

Tilt azimuth (downslope): 323.2693° cw from leg 1.

Alternatively,

North is 141.90511° ccw about the LGN z-axis, from the LGN y-axis.

Deck tilted downward 2.991820° at azimuth 285.1744° cw from north (LMS y-axis).

A4.2 Lander 2

The systematic map T^{SR} matrix for Viking lander 2 is

$$T^{SR} = \begin{pmatrix} +.1414660 & -.8623340 & +.4861690 \\ -.0191174 & +.4886360 & +.8722730 \\ -.9897580 & -.1326930 & +.0526407 \end{pmatrix}.$$

The following orientation parameters may be obtained from the T^{SR} matrix for lander 2 (Note: When viewing normal to the deck of the lander, the LACS negative z-axis coincides with the projection of leg number 1, the aft pointing leg.):

North: 150.8665° cw from leg 1.

Tilt magnitude: 8.207321°

Tilt azimuth (downward): 68.56271° cw from leg 1.

Alternatively,

North is 29.1335° ccw about the LGN z-axis, from the LGN y-axis.

Deck tilted downward 8.207321° at azimuth 277.6962° cw from north (LMS y-axis).

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APPENDIX B

SELF-OBSCURATION OF MARS BY THE LANDER

B1 Self-Obscuration

If the policy in lander design been exclusively to satisfy the interests of the lander imaging team, much more of the landing site would have been visible to the cameras. In reality, there were of course many competing engineering and scientific interests that had to be considered in configuring the lander. It will be recalled, from figure A1, that the two cameras were jointly placed closest to the side of the lander on which the surface sampler arm assembly was mounted. This was to enable the cameras to support surface sampler activity. This side of the lander side was referred to as the front. The extent to which the final configuration of the lander and the placement of the cameras thereon influenced the visibility of the landing site to the cameras is indicated in the "Skyline Drawings" of figure B1. The top half of the figure indicates the view from camera 1, the "left-hand" camera, and bottom half that from camera 2, the "right-hand" camera. The figure indicate the fields of view of each of the cameras, from 40° above the horizon to 60° below, through 360° of azimuth. The nominal horizon is located along the fiducial line just above center in each half of the figure. The

canister-like structure at the azimuthal extreme in each of the views is the other camera.

The figure indicates the appearance of the various pieces of lander hardware as well that of a rectangular grid on the nominal landing plane in front of the lander. The front of the lander is in the right half of the camera 1 drawing and in the left half of that for camera 2. It is evident from these figures that not much solid angle of the nominal Martian landing plane can be seen behind the lander. In order to perform stereophotogrammetry of the Martian surface, it is of course necessary for the surface to be visible to both of the cameras. (There were at least two mirrors on the lander, but the author is not aware that any significant attempts were made to do perform stereophotogrammetry by pairing real and virtual images.) A careful inspection of the skyline drawings, with the aid of the grid on the nominal landing plane, will indicate that substantial portions of the near field in front of the lander are obstructed from the joint view of the cameras.

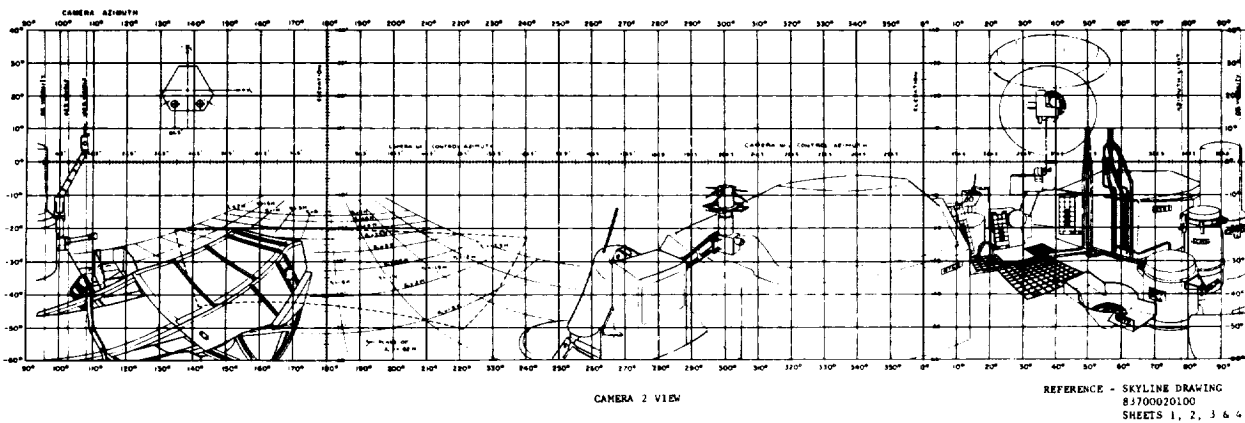
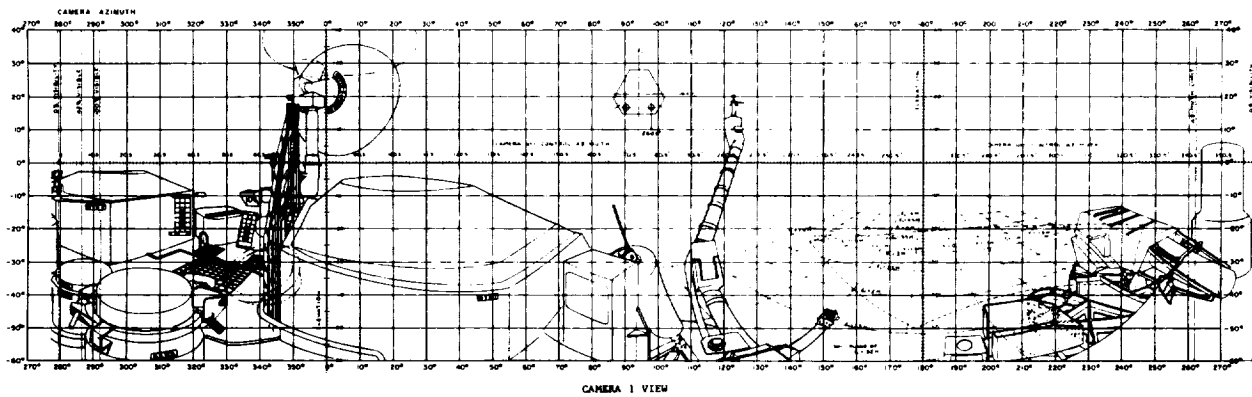


Fig. B1. Skyline Drawings.

APPENDIX C

MAP GENERATION SYSTEM AND PRODUCTS

C1 Ranging and Mapping System

For further details than contained in this report regarding the Stanford Stereo Station and the most commonly utilized features of the program RANGER, the reader is referred to earlier work [Liebes, 1977]. Additional information regarding the program RANGER resides in unpublished documents of JPL's Image Processing Laboratory.

C2 Map Product Types

The present report is concerned with the sets of systematic elevation contour and vertical profile maps that have been produced for the two Viking landing sites. During the mission other types of map products were produced. No systematic distribution of the remainder of the map products has either been undertaken or is planned. The bulk of the these other products were generated in support of the sampler arm activity during the mission. This activity primarily involved soil sample acquisition for delivery to onboard experimental laboratories, and material manipulation in support of physical properties experiments. The map products generated for this support were in the form of arm-specific "vertical" (relative to the lander) profiles. The format of these products has been previously described [Liebes, 1977]. These products, along with photographs of the arm-specific vertical profiles overlayed on unrectified stereo images pairs of the subject surface region were delivered to members of the individual Viking lander science teams, to members of the sampler arm team, and to management during the mission. There has been no other distribution. Miscellaneous other products have been generated. Listed below is a summary of the types of range data set products have been generated:

1. Systematic contour map data (intersections with Mars of planes oriented perpendicular to the local Mars zenith), extending from the immediate foreground to the remote limits of ranging capability, in excess of 100 m range, for the front and back of both landers. Half-size versions of all of these products are contained in the present report.
2. Systematic vertical profiles (intersections with Mars of planes radiating out from the lander and containing the local Mars zenith), at 5°-azimuth intervals, from the immediate

foreground to the remote limits of ranging capability, for the front and back of both landers. Half-size versions of all of these products are contained in the present report.

3. An orthographic projection of ridge and trough center lines for the front of lander 1. This information was used to aid in the preparation of a geological map of the Viking 1 landing site [Binder, 1977] (a drafting error resulted in the labeling of a scale bar in figure 5 of that report as 10 m in length, instead of the correct value of 20 m).

4. An asystematic contour mapping of a drift field in the left front of lander 1. This data has been published [Mutch, 1976b].

5. Numerous detailed sampler-arm-specific lander aligned "vertical profiles" (intersections with Mars of planes perpendicular to the lander deck and radiating out at various arm azimuth angles from the vertical pivot axis of the sampler arm) as required to support sample acquisitions, "rock-rollings," trenching operations, and site disturbance analysis during the active part of the mission. Much of the mission support sampler-arm-specific stereophotogrammetry was performed by Ray Jordan of the U.S. Geological Survey, Flagstaff, AZ. The format of these profile maps, and examples of the products have been reported [Liebes, 1977].

6. Detailed vertical profile and contour map data of selected "rock-roll" candidates. The profiles were identical to those referenced in 5., above. An example of a detailed profiling of a "rock-rolling" candidate has been published [Moore, 1978].

C3 RDS File Names

The file names for all of the RDSs used in generating the contour and vertical profile maps for both landers are indicated below. As mentioned in sect. 3.3 of this report, the differences in computer types at the two institutions necessitated different formats for the names. Operational procedures lead, as a rule, to the division of the RDSs into pairs for the front and again for the back of each lander. The lander 2 contour set is an exception, in having three sets in front.

C3.1 Viking Lander 1

C3.1.1 Lander 1 Systematic Elevation Contour RDS File Names

Stanford University	Jet Propulsion Lab.
L1CFP.DAT[CON,VL1]	L1.RDSC.FRONT.VT1025
L1CFA.DAT[CON,VL1]	L1.RDSC.EDR.FRONT
L1CBP.DAT[CON,VL1]	L1.RDSC.BACK
L1CBA.DAT[CON,VL1]	L1.RDSC.EDR.BACK

C3.1.2 Lander 1 Systematic Vertical Profile RDS File Names

Stanford University	Jet Propulsion Lab.
L1VFP.DAT[PRO,VL1]	L1.RDSV.FRONT3
L1VFA.DAT[PRO,VL1]	L1.RDSV.EDR.FRONT
L1VBP.DAT[PRO,VL1]	L1.RDSV.BACK
L1VBA.DAT[PRO,VL1]	L1.RDSV.EDR.BACK

C3.2 Viking Lander 2

C3.2.1 Lander 2 Systematic Elevation Contour RDS File Names

Stanford University	Jet Propulsion Lab.
L2CFP.DAT[CON,VL2]	L2.RDSC.FRONT
L2CFP2.DAT[CON,VL2]	L2.RDSC.FRONT2
L2CFA.DAT[CON,VL2]	L2.RDSC.EDR.FRONT
L2CBP.DAT[CON,VL2]	L2.RDSC.BACK
L2CBA.DAT[CON,VL2]	L2.RDSC.EDR.BACK

C3.2.2 Lander 2 Systematic Vertical Profile RDS File Names

Stanford University	Jet Propulsion Lab.
L2VBP.DAT[PRO,VL2]	L2.RDSV.BACK
L2VBA.DAT[PRO,VL2]	L2.RDSV.EDR.BACK
L2VFA.DAT[PRO,VL2]	L2.RDSV.EDR.FRONT
L2VFPR.DAT[PRO,VL2]	L2.RDSV.FRONT.V70104

C4 Map Plotter

All of the maps were formatted and produced by computer. The output device was a Varian Statos 4222 electrostatic printer/plotter, with 4224 styli spaced at 0.005 inch intervals over a 21.12 inch writing width.

APPENDIX D

STEREO SYSTEM RANGING ACCURACY

D1 Ranging Accuracy

The ranging accuracy of the stereophotogrammetry system applied to the Viking lander camera images has been previously discussed [Liebes, 1977]. A plot of the theoretical point ranging error as a function of range is reproduced, from this reference, in figure D1. The error curves are plotted for three different values of azimuth, relative to a 0° reference direction perpendicular to the intercamera baseline. The plots are made for 0.04° resolution imagery data that characterizing the high resolution mode employed in creating the mosaics used in producing the systematic maps.

The error curves have been constructed as follows. Consider a plan view of the stereo camera system, such as indicated schematically in figure A1 (a). Imagine there to be associated with a given field point a pair of wedges radiating out from the cameras, each of apex angle equal to the angular resolution of the cameras, and crossing at the field point to form a diamondlike zone of intersection. Roughly speaking, a field point located anywhere within this overlap zone will be recorded at the same fixed pixel locations in the two images. The diamond therefore indicates the zone of uncertainty for the range. For the purpose

of the plot, the range uncertainty is taken to be the length of the diamond in the direction away from the lander. This is a conservative figure from the point of view of an ideal system, since a) the extreme variability of range relative to a central point in the diamond is only half this magnitude, and b) the shape of the diamond will render the rms error even smaller than this. However, there are uncertainties regarding the true pointing directions of the cameras. These reflect prelaunch calibration determinations [Wolf, 1981], nonlinearities in the cameras, settling of the lander and diurnal thermal warp of the hardware, etc. Furthermore, the budget for generating the map data corresponded to a commitment of a certain amount of mapping time. A compromise had to be struck between the conflicting desires for density of mapping coverage and accuracy of the product. It is judged that that the uncertainties shown in the figure, are a reasonable measure of the absolute accuracy of the map data. Local relative potential accuracy of the system is estimated to be roughly an order of magnitude better than that plotted in the figure.

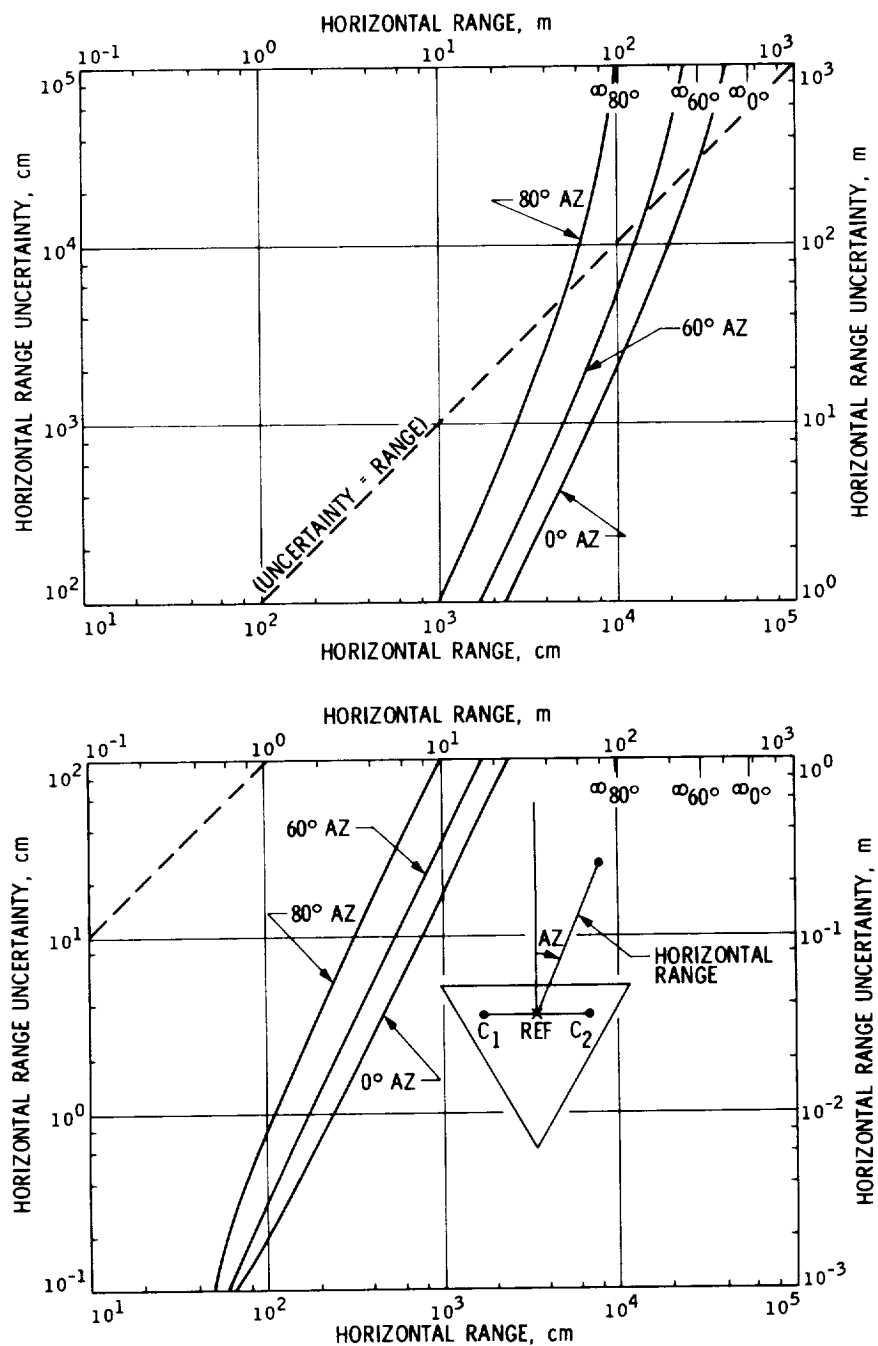


Fig. D1. Ranging Accuracy as a Function of Range.

APPENDIX E

RAGGEDNESS AND DISTORTION

E1. Jaggies and Irregularities

We have described in section 2.1 the principle of operation of the stereo mapping system. The contours/profiles were generated by manually commanding a 3-D overlay cursor to follow the perceived intersection of the appropriate mathematical surface of constraint with the Martian relief. The hand-input device for performing cursor control was a trackball. This device consists of a baseball-sized sphere, roughly the top third of which protrudes through a hole in the top of a box. Inside of the box, the ball rests upon a pair of orthogonally orientated angular encoders, the positions of which may be read by the computer. One may freely define the parameterization between trackball angular position and the location of the 3-D cursor on the surface of constraint in the Martian model. The parameterization adopted was approximately the following. Lateral motions of the ball were translated linearly to angular displacements of the 3-D cursor about the LGN z-axis, i.e. the zenithal direction through the lander. Rotation of the top surface of the ball away from or toward the operator was translated into nonlinear variation of the range of the cursor from the lander. Specifically, the range parameterization established a linear relation between angular displacement of the trackball and variation of the convergence angle from the two cameras to the field point. This parameterization was selected to enable the operator to move to remote field points without having to make an inordinate number of rotations of the trackball. The parameterization was also qualitatively consistent with the degradation of ranging resolution with range.

This selection of parameterization resulted in the introduction of an unanticipated artifact in the maps. Before explaining this, we make several introductory comments, dealing first with an aspect of mammalian, and in particular human, stereo performance. Humans use a variety of clues other than stereoscopic in judging the absolute and relative range of objects and features. Among these are convergence, foreshortening, diminution, shadow, shading, contrast, inference from familiar kinds of features, and the like. The author personally performed all of the stereophotogrammetry represented by the maps of this report. It was not an uncommon experience, while generating contour/profile lines, for him to experience a lag in awareness that the 3-D cursor had been carried from a region where stereo data existed into a region where, as a result of obscuration to one camera, only monoscopic data existed. These transgressions were invariably discovered after a few pixels of trespass into forbidden territory. Under such a circumstance, it would be necessary to edit out the faulty data from the range data set; a convenient means was provided for doing this. The point is that the nonstereoscopic clues to depth can be very strong, and even though there may be a degree of qualitative validity to the relative range determinations, their intentional use was out of order

in the context of the present work.

The reader may have noted that when looking at an outdoor scene, there is no conspicuous range at which one has a sense that one's stereo perception is breaking down and other clues to depth are becoming dominant. Such was also the author's experience in mapping remote features, where the convergence angle began to approach the resolution of the cameras. Under such circumstances, one often had a distinct sense of local depth variation from other than stereoscopic clues; e.g. if a rock were being examined, brightness and shadow variations were powerful cues.

We have remarked that the cameras were point scanning devices, and that the data radioed back to earth consisted of distinct brightness values for a raster of digitized points in the scene. These individual digitized points were resolved on the video screen. Now, it was also true that the two component points of light, overlayed on the images on the video monitors to generate the 3-D cursor were written to a one-bit raster corresponding 1:1 to the raster locations of the displayed image points. This led to a frustrating appearance of the 3-D cursor stitching back and forth in depth through the relief as the cursor was carried across the scene. It seems that the human being is able to establish an average range for micro-facets of extended textured objects to greater precision than the point ranging accuracy associated with the resolution of the cameras and display. Thus, there was the inherent capability of the human operator to place the cursor more accurately in local average depth than the system would permit.

The trackball had an inherent angular resolution, corresponding to the quantized angular rotational increment required to generate an output pulse. RANGER recorded and calculated range changes at the resolution of the trackball. In addition to a compiled-in trackball parameterization, there was a commandable scale factor controlling the sensitivity of displacements in 3-space to the rotational sensitivity of the trackball. A qualitative attempt was made to maintain the relationship between the quantized reading sensitivity of the trackball, corresponding 1:1 with the detail of the map data, a bit finer than the ranging resolution associated with the display raster. This enabled interpolation of range to finer detail than theoretical point ranging accuracy. The trackball and ranging sensitivity is evidenced in the plots of the map data by the mutually orthogonal elemental line segments, or "jaggies" in the lines.

Enough for the introductory remarks. The map data was generated and plotted out to ranges where the ranging error was substantial. The outer mapping limits were generally well beyond the range where cosmetic considerations alone might dictate calling it quits. The reason for working this far out was to milk the data for all it was worth. It was appreciated that the

audience for these maps would be sophisticated. We were anxious to capture whatever information we were able to, consistent with the resources committed to the task. As there is no post editing, detectible smoothing, or other form of cosmetizing applied to the data, indications of degradation, such as size of the jaggies, are evident to the investigator. He can do his own discounting.

Consider the case where we are trying to map a feature, the extent of which in the range direction is small compared to the ranging resolution of the 1-bit raster-limited 3-D cursor. We have explained above that the trackball was read to finer granularity than that corresponding to a displacement of the displayed 3-D cursor. This, in effect, enabled us to interpolate the data in order to capitalize on the ability of the human to locally average over an extended textured feature. The products support the merit of this practise, in that range noise on contour lines running over smooth but textured surfaces was generally less than point ranging accuracy would have lead one to expect. Additionally, adjacent lines in elevation on such surfaces, which should not cross unless the surface is vertical or overhanging, did not generally cross as much as one would expect on the basis of point ranging accuracy.

Now, here is the nub. Consider the the cursor to be resting as closely as it can be visually placed upon a remote rock. As the cursor is drawn across the face of the rock in order to generate a contour or profile line, one has a distinct, and to a significant extent meaningful, sense of range variation. However, the inherent limitations of the 1-bit graphics raster display precluded subpicture element visual feedback as the cursor is played over the small perceived variations in depth. One's instinct in moving the cursor over the contour of e.g. an apparently roughly spherical rock is to rotate the trackball through approximately the same range in lateral as in longitudinal angle. For these small rotations there is not sufficient variation in range for the differential disparity of the left and right components of the cursor to register on the display as a jump in depth. However, the motion of the tackball was recorded and subsequently displayed on the maps. Now, given the nature of the parameterization of the trackball, at relatively great range equal lateral and longitudinal rotations of the trackball lead to much greater jumps in range than in lateral 3-

space displacement. As there was either no or insufficient visual feedback clues to alert the photogrammetrist to the degree of the problem, it was difficult to compensate satisfactorily in the blind. Furthermore, the photogrammetry was conducted at the JPL, and the vast bulk of the map products were not plotted until we brought the data back to Stanford University, well after the photogrammetry was completed. Thus, we were not greatly sensitized to the degree of the effect during the data generation period. The artifacts are evident in the tendency for the contour and profile lines to tend to be excessively v-shaped toward the lander for the relatively more remote features.

This leads to ambiguity in judging just what part of a contour or profile line on a remote feature most accurately represents the range of the feature. If an isolated remote rock were to be contoured, it is likely that the first point of contact with the rock was the most accurate indication of its range. However, it is not possible, in looking at the map data to determine this point. In some instances the cursor might have been carried to the near nose of the rock, in "no-draw" mode, then moved back along the side in no-draw mode, via a judgement call on the range change; then the contour could be swept across the face. In this instance, the range of the nose would tend to be the most accurate value for the rock. If, however, the cursor were initially placed to the side of the rock in no-draw mode, and then the contour traced, the measure at the side of the rock would tend to be the more valid. There is no way to know which method was employed in any given situation.

Had the magnitude of this effect become evident to us sooner, we likely would have introduced an interactive means for calling into play, as necessary, a 1:1 parametric relationship between 3-D displacement and trackball rotation.

We might also note that one approach to improving the resolution of the 3-D cursor, and the overlay contour/elevation graphics as well, would be to introduce a spatially extended multi-bit gray-level graphics supporting interpolation between raster elements.

Rather than progress further into finer points of detail, it is perhaps best to let the matter rest with a warning to the reader to interpret clearly noisy data with care.

APPENDIX F

DATA AVAILABILITY

F1. General Lander Image Products

The NASA Contactor Report on "The Mosaics of Mars" [Levinthal, 1980] provides detailed references and suggestions for accessing the diverse Viking lander camera related data products.

F2. Systematic Maps Products

F2.1 Systematic Maps and Mosaic Overlays

Individual copies of the true-scale versions of the maps, as well as copies of the mosaic overlays, contained in this report may be obtained from the sources indicated below. In the case of Viking lander 1, reference should be made to data number NSSDC ID 75-075C-06T, and in the case of Viking lander 2 to NSSDC ID 75-083C-06T.

Researchers within the United States should contact:

The National Space Science Data Center
Code 601.4
Goddard Space Flight Center
Greenbelt, Maryland 20771

Researchers outside the United States should contact:

World Data Center A
Rockets and Satellites
Code 601
Goddard Space Flight Center
Greenbelt, Maryland 20771

F2.2 Range Data Sets

The Range Data Sets are stored on magnetic tape. There are a total of 9 nine-track/800 BPI tapes for elevation contour and vertical profile RDSs. There are additionally 18 nine-track/800 BPI magnetic tapes of arm-specific and miscellaneous RDSs. RDSs are not available from the National Science Data Center. They are stored at the following locations:

Image Processing Laboratory
Jet Propulsion Laboratory
4800 Oak Grove Drive
Pasadena, CA 91103

Regional Planetary Image Facility
Jet Propulsion Laboratory
4800 Oak Grove Drive
Pasadena, CA 91103

Additionally, individual members of the Viking Lander Imaging Team have received copies of these tapes. For the names and addresses of these recipients, refer to [Levinthal, 1980].

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